



The Dark Energy Survey in 10 mins

**“ONE OF THE MOST AMBITIOUS
astronomical surveys in history ...”**

Yuanyuan Zhang, Fermilab
Schramm postdoctoral fellow
New Perspectives 2016



The Dark Energy **Survey** in 10 mins

A gigantic sky
“photography”
project.

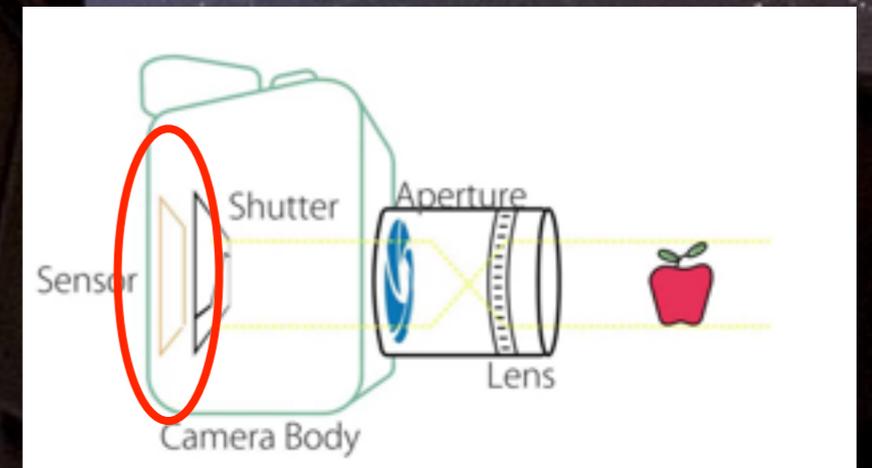
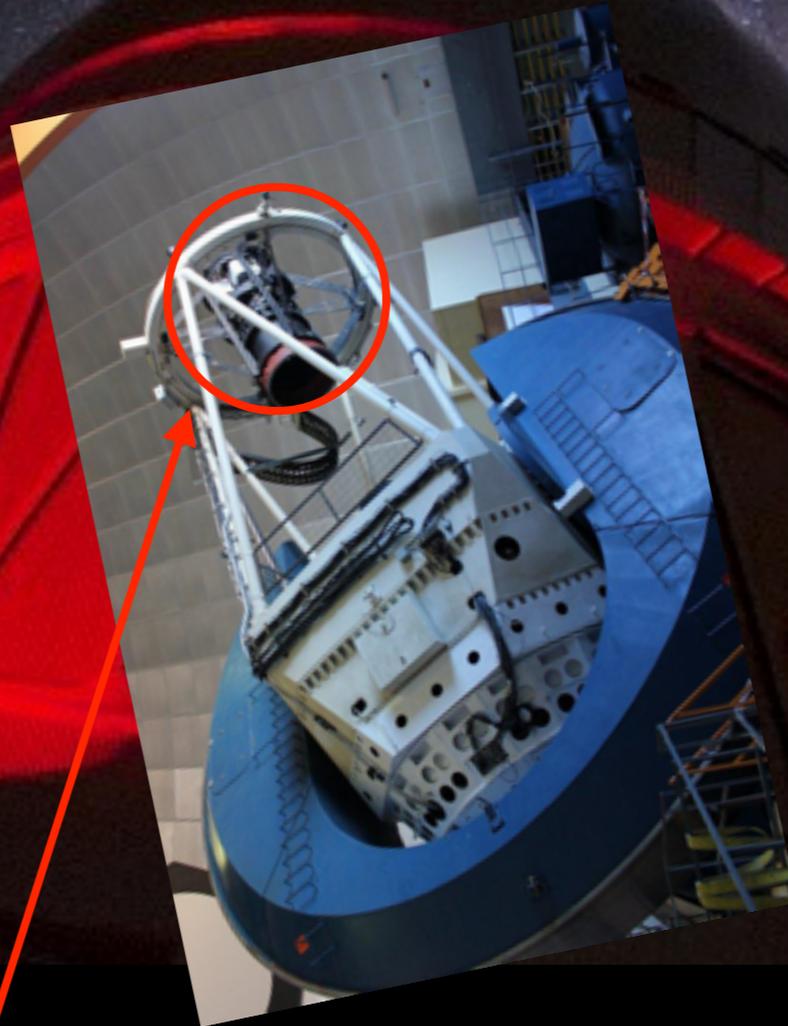
An extremely wide-field deep optical imaging survey using the Dark Energy Camera mounted on the Blanco telescope.

Background credit: Reidar Hahn



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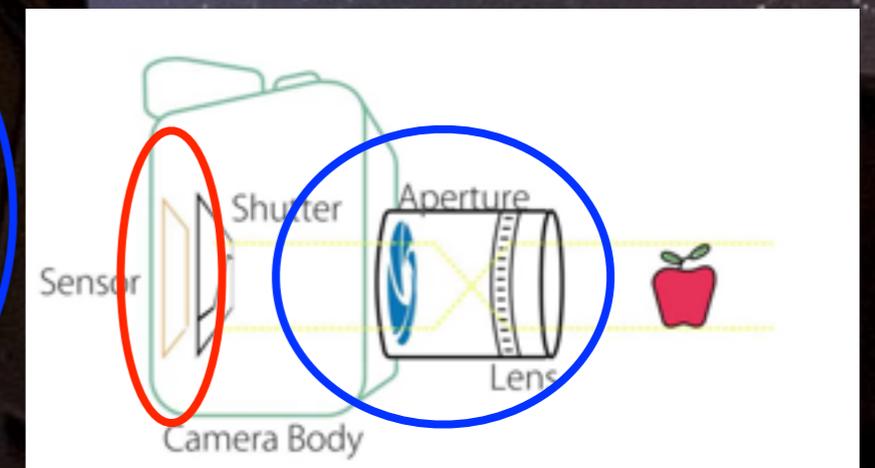
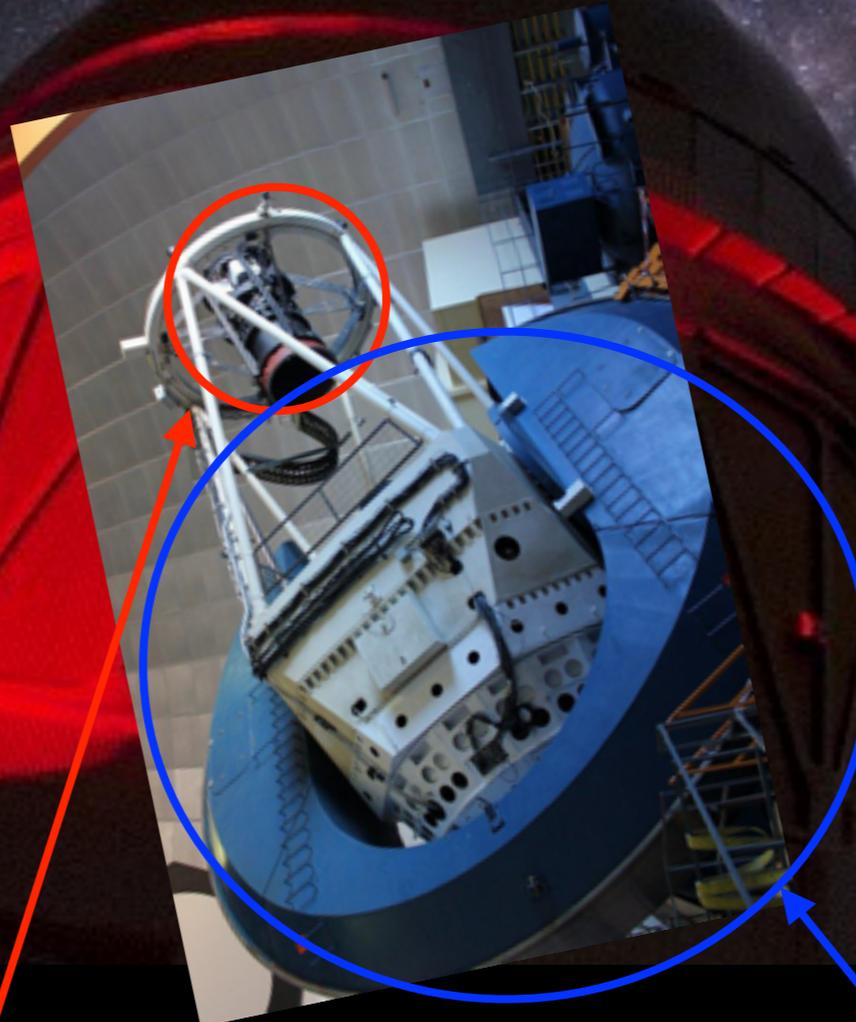


An extremely wide-field deep optical imaging survey using
the Dark Energy Camera mounted on **the Blanco telescope**.



The Dark Energy **Survey** in 10 mins

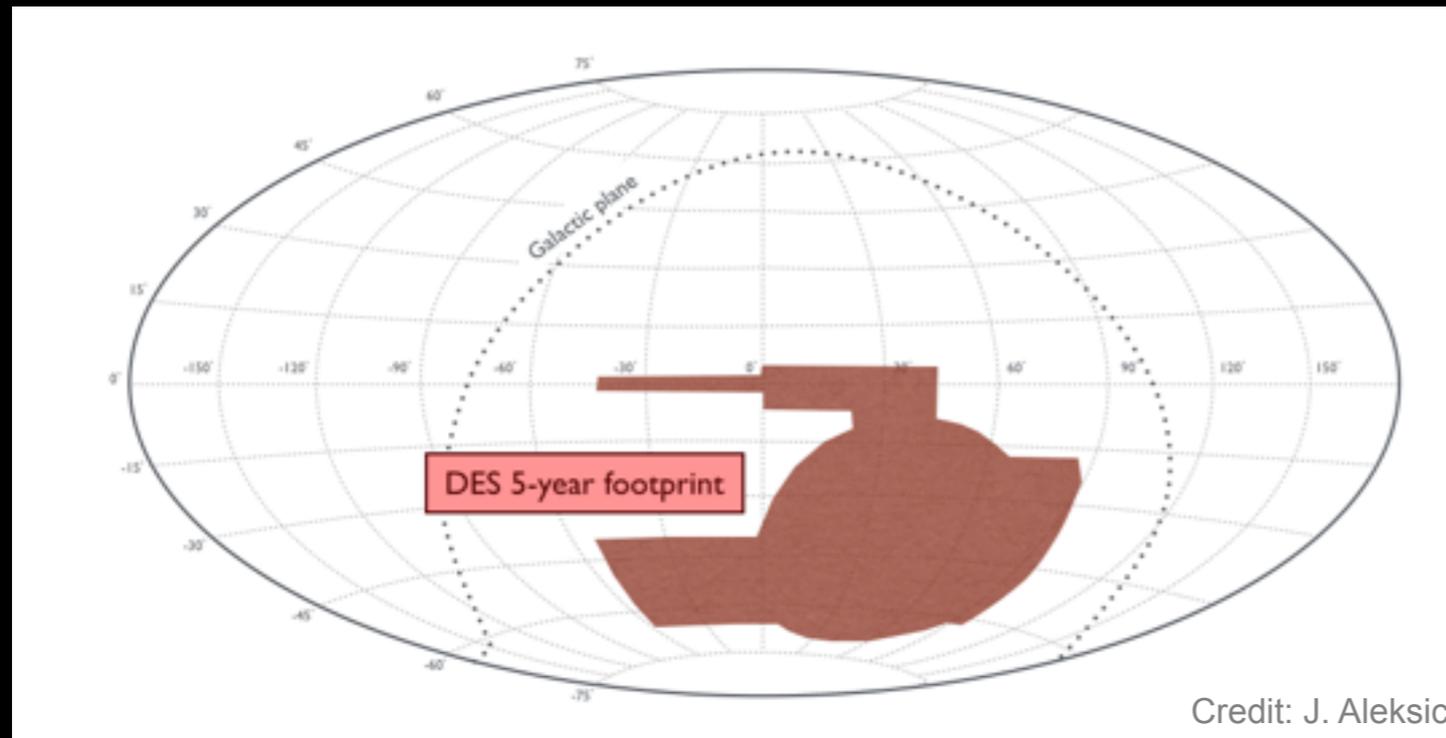
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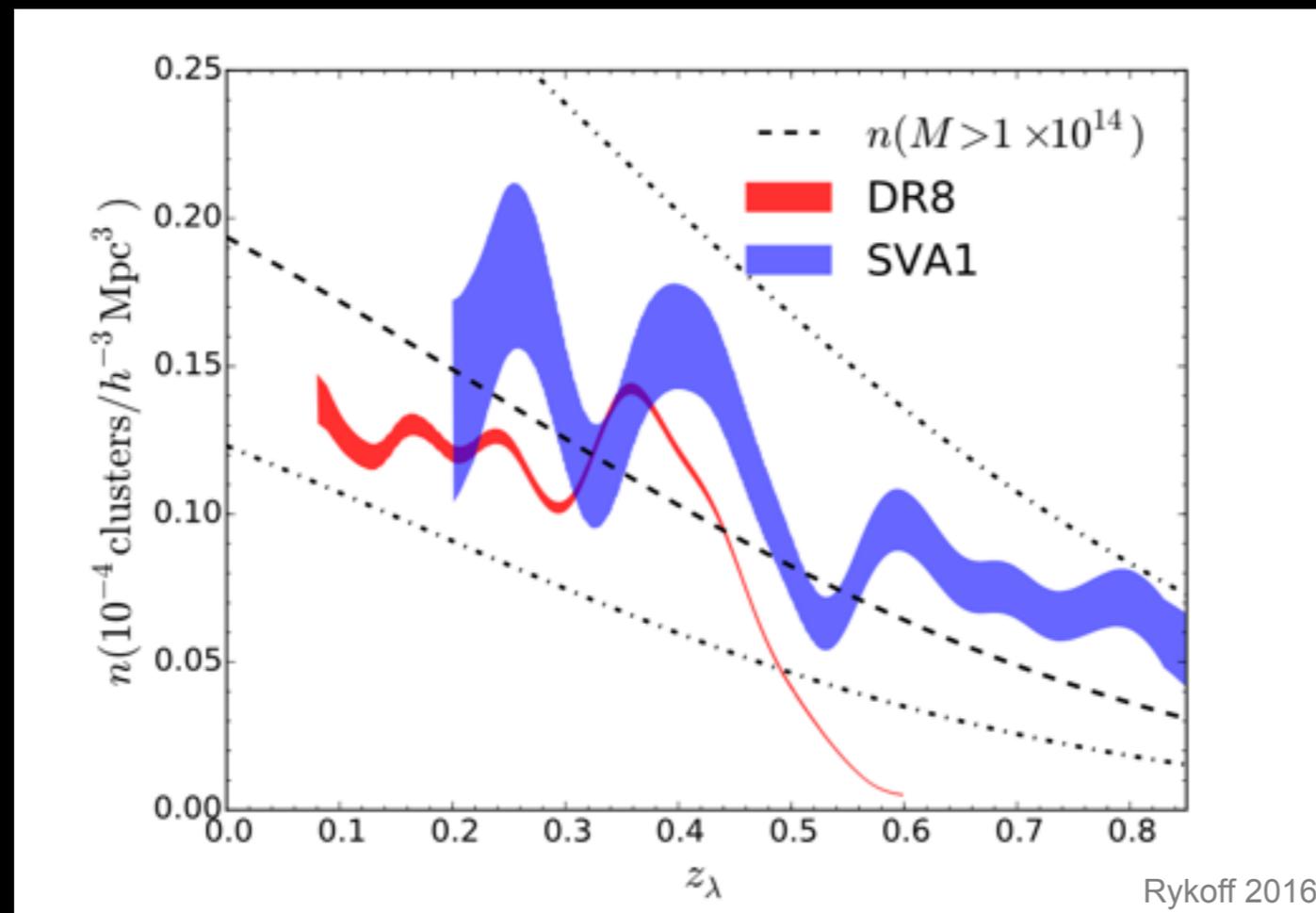
DES covers 5,000 deg^2 of the sky, which is made possible with DEcam's 3.0 deg^2 field of view.

An **extremely wide-field** deep optical imaging survey using the Dark Energy Camera mounted on the Blanco telescope.



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The depth and the wavelength range allow cosmological studies to $z \sim 1$.



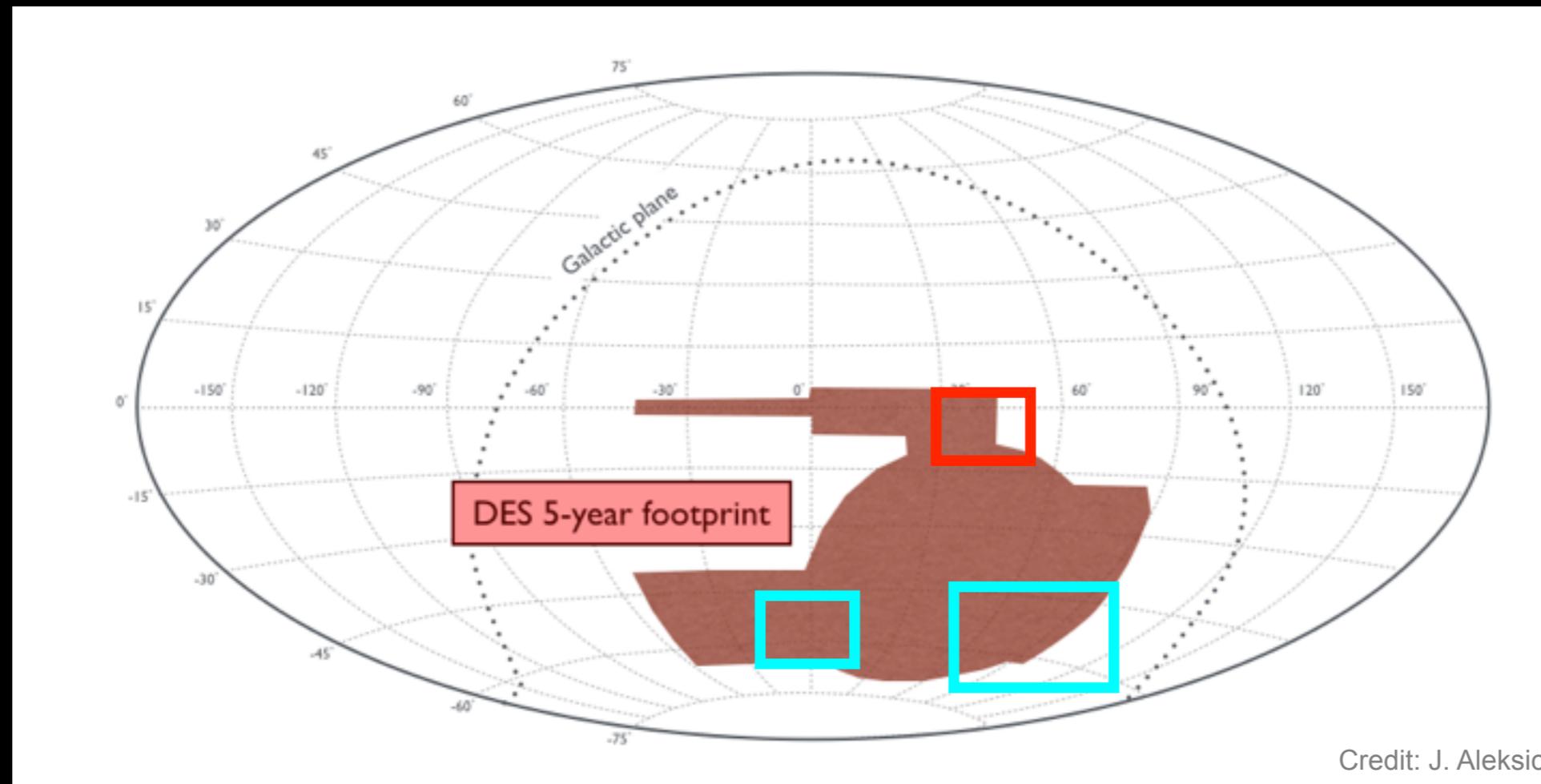
An extremely wide-field **deep optical** imaging survey using the Dark Energy Camera mounted on the Blanco telescope.



The Dark Energy **Survey** in 10 mins

Data taking will continue for another two years, using ~500 nights in total.

- **Science Verification 2012**: small area, full depth, data released. 35+ papers.
- **Official start in 2013**: Data being processed. 9+ papers posted
- **Supernovae survey**: 30 deg², on sky since 2012. 9+ papers.





The **Dark Energy** Survey in 10 mins

Dark energy affects the distance-redshift relation.

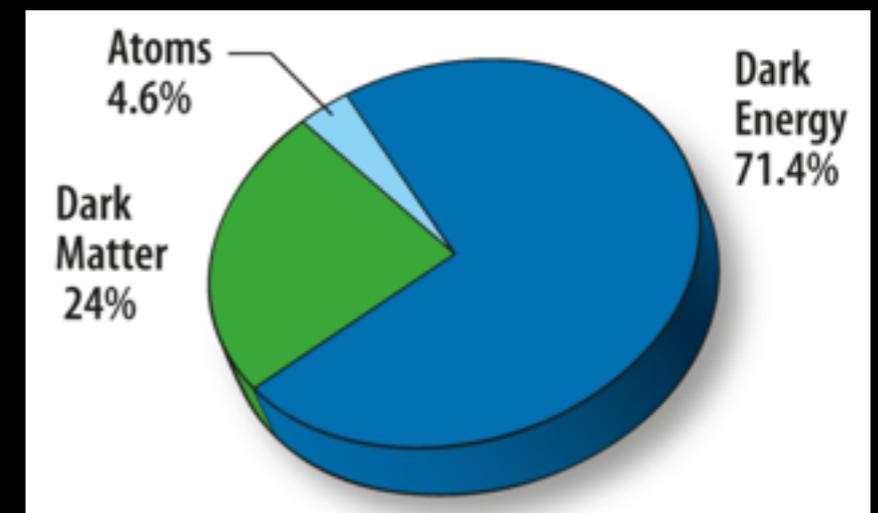
DES measures the “distance” via:

- Type Ia supernovae
- Baryon acoustic oscillation

Dark energy affects cosmic structure formation.

DES quantifies structure formation via:

- weak lensing
- galaxy clusters

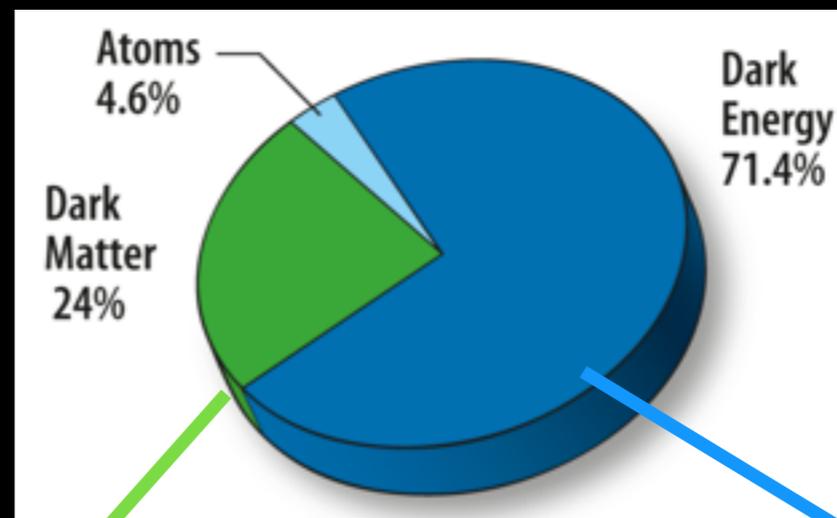


Credit: nasa.gov



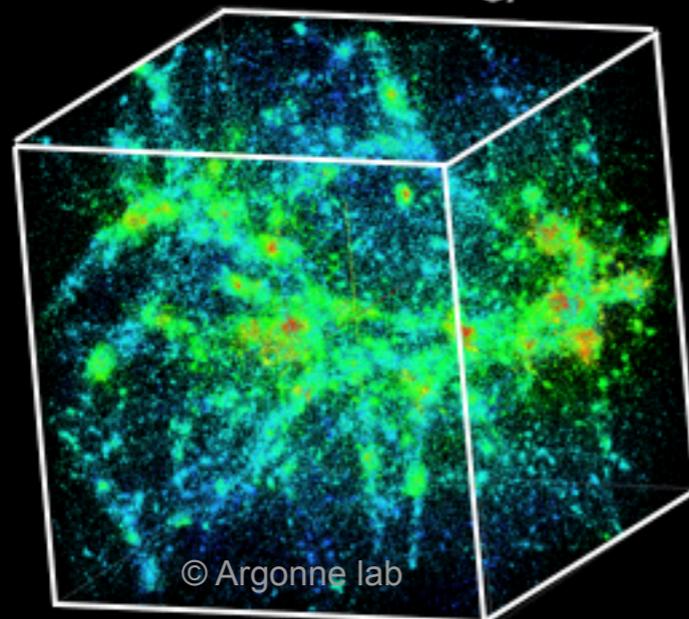
The Dark Energy Survey in 10 mins

Dark energy affects cosmic structure formation.

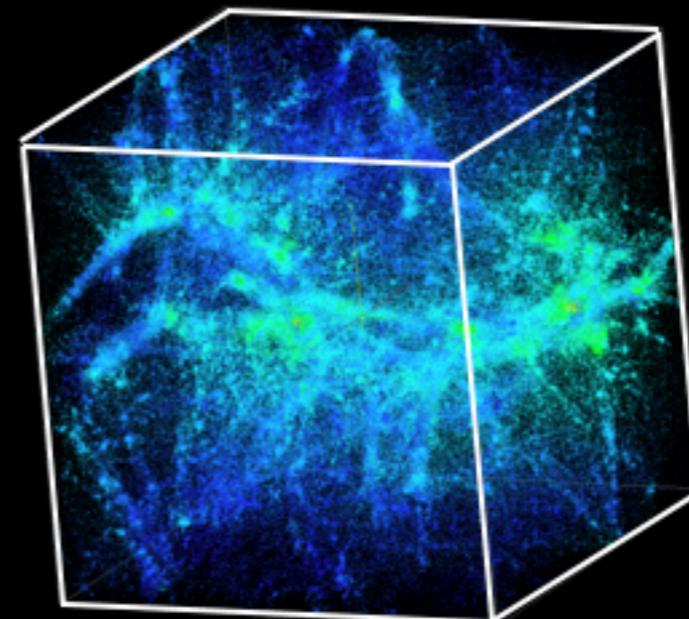


Credit: nasa.gov

No dark energy



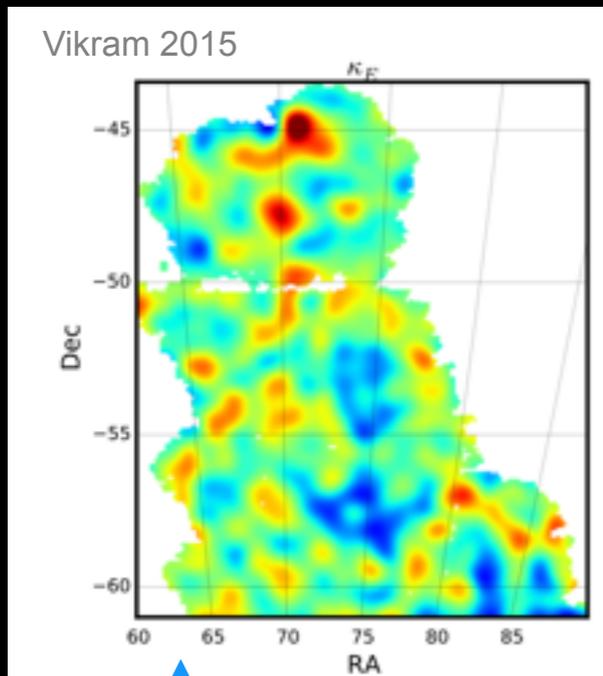
Standard Model





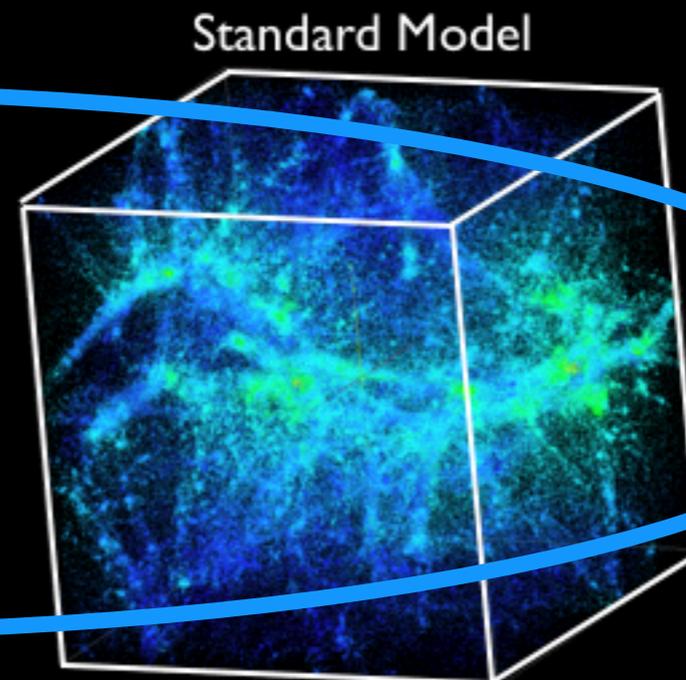
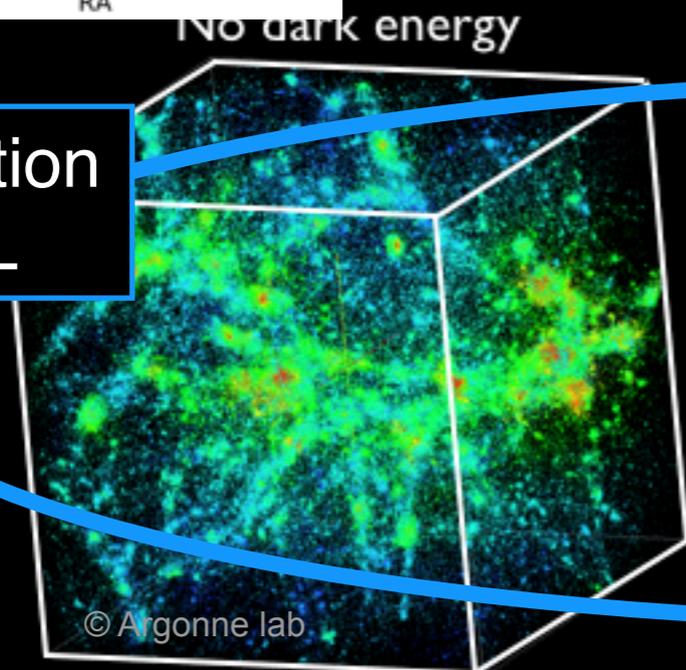
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DES quantifies the structure formation clumpiness with **weak lensing**.



Hear more about weak lensing on Wednesday with Daniel Gruen.

Observation via WL

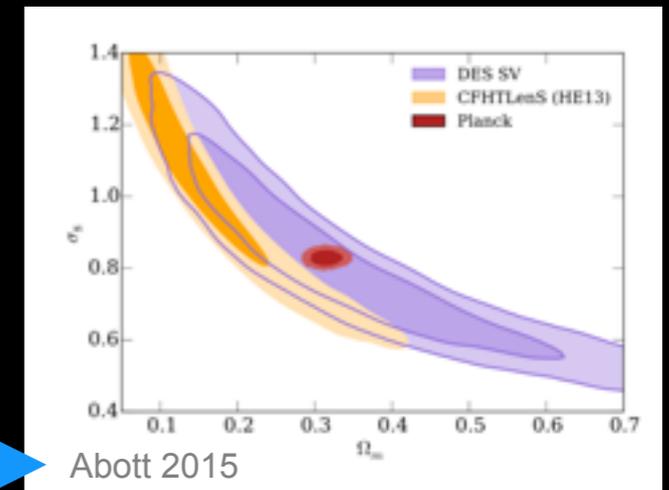
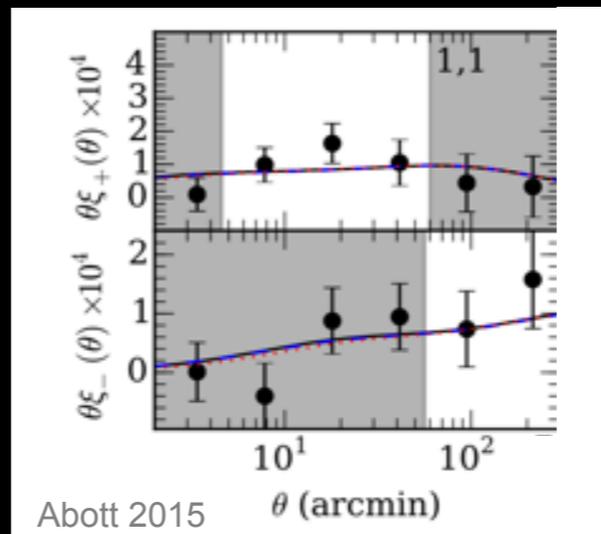
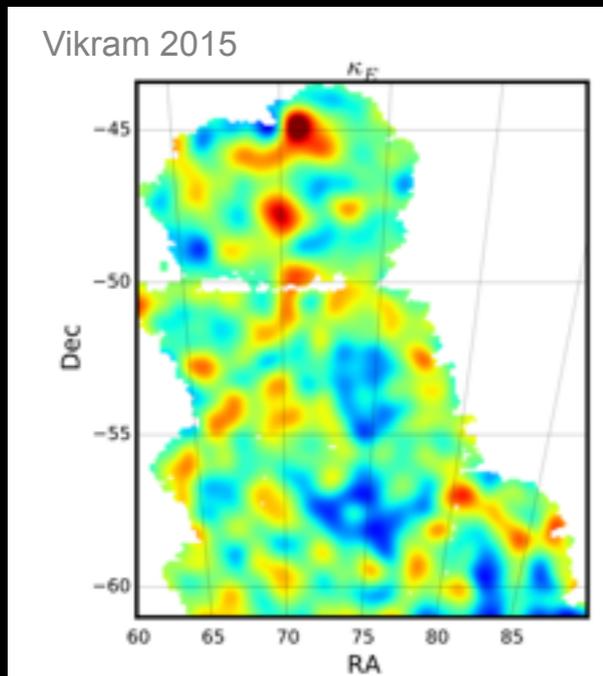


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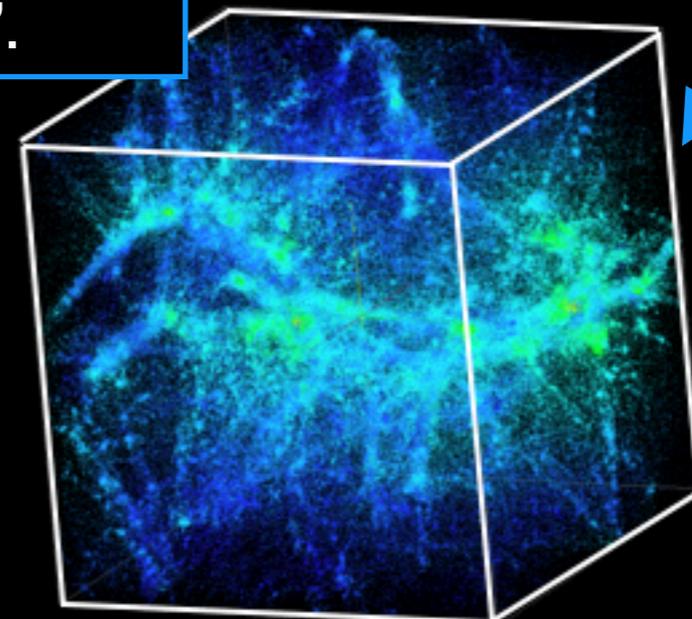
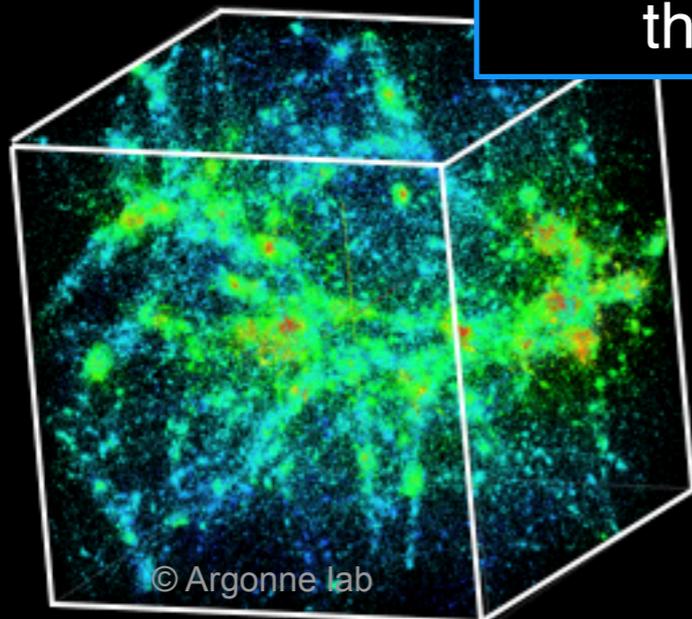


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DES quantifies the structure formation clumpiness with **weak lensing**.



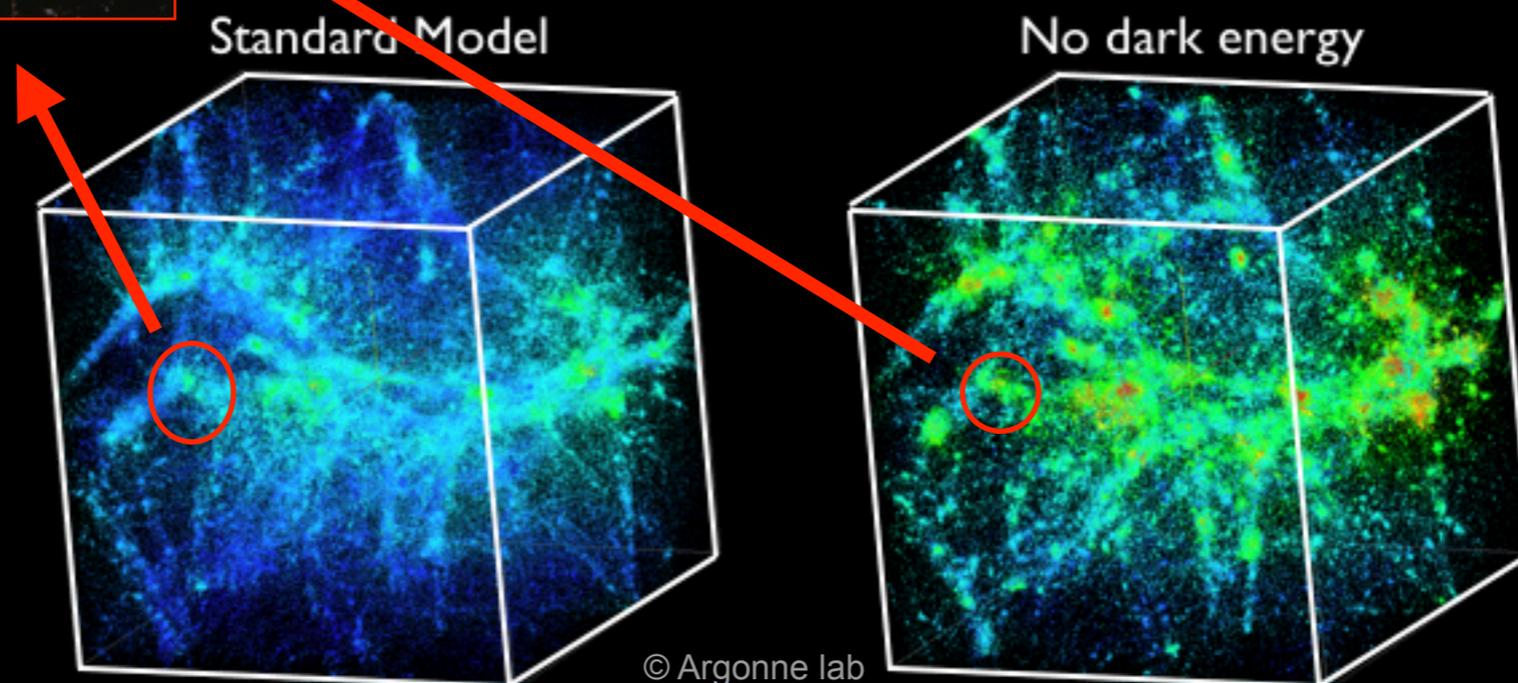
Correlation analysis agrees with the "standard model".





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DES quantifies the structure formation clumpiness with **galaxy clusters**.



Galaxy Clusters are the peaks of the clumpy density field.

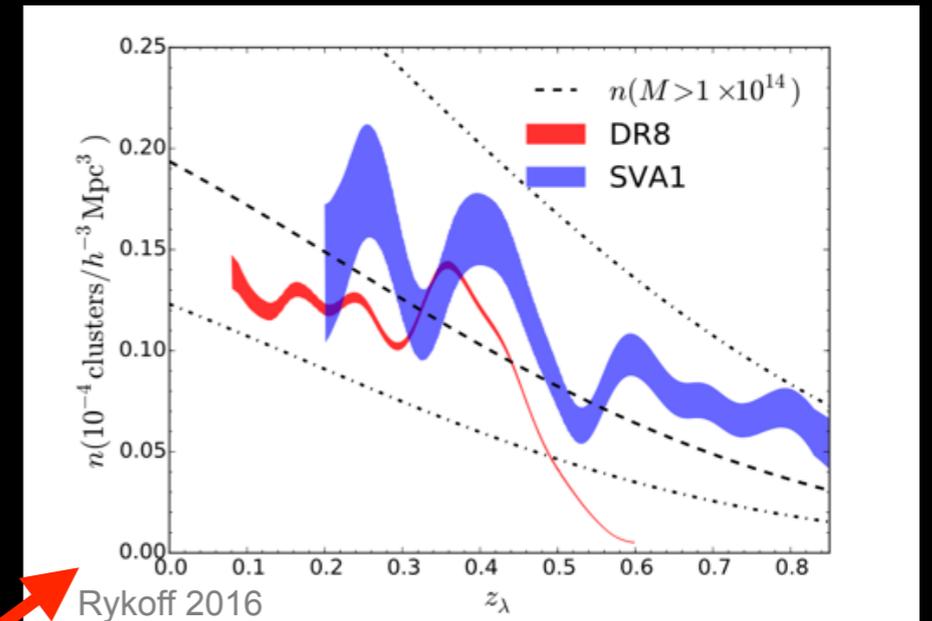
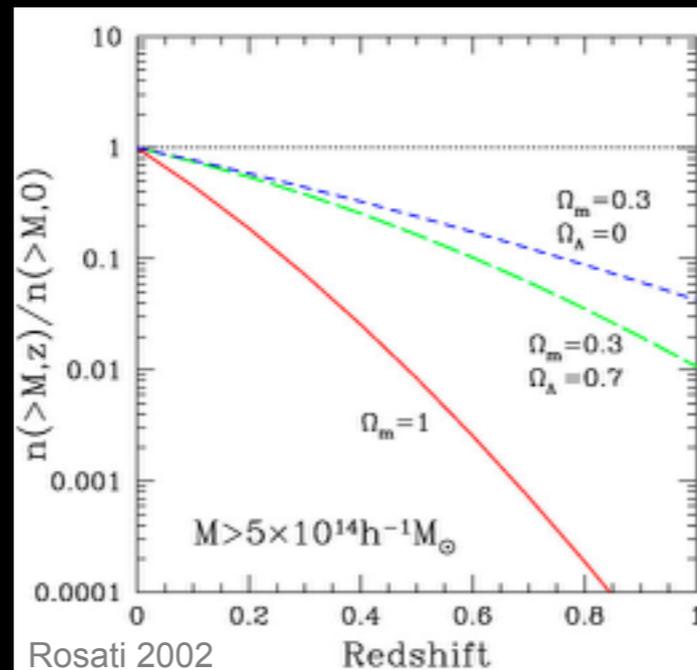


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Number densities sensitive to on cosmology models.



DES observations in progress:
Clusters are identified (Rykoff 2016).
Calibration in progress (Saro 2015, 2016).



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Other than Dark Energy:

Discovery of rare objects:

- Milky Way satellite galaxies
 - **Two talks by Alex Drlica-Wagner.**
- Strong-lensing galaxy clusters
 - **One talk by Brian Nord, end of today.**
- Discovery of Kuiper Belt objects (planet nine?)

Big Data yield high statistical precision:

- Galaxy astrophysics processes.

Making use of DECam

- Follow up gravitational wave events (led by Fermilab).

The Collaboration

 **Fermilab** — The Fermi National Accelerator Laboratory

 **Chicago** — The University of Chicago

 **NOAO** — The National Optical Astronomy Observatory

 **United Kingdom DES Collaboration**

- **UCL** - University College London
- **Cambridge** - University of Cambridge
- **Edinburgh** - University of Edinburgh
- **Portsmouth** - University of Portsmouth
- **Sussex** - University of Sussex
- **Nottingham** - University of Nottingham

 **DES-Brazil Consortium**

 **OSU** — The Ohio State University

 **TAMU** — Texas A&M University

Munich—Universitäts-Sternwarte München

-  **Ludwig-Maximilians Universität**
-  **Excellence Cluster Universe**

 **UIUC/NCSA** — The University of Illinois at Urbana-Champaign

 **LBL** — The Lawrence Berkeley National Laboratory

 **Spain DES Collaboration**

- **IEEC/CSIC** - Instituto de Ciencias del Espacio,
- **IFAE** - Institut de Fisica d'Altes Energies
- **CIEMAT** - Centro de Investigaciones Energeticas, Medioambientales y Tecnologicas

 **Michigan** — The University of Michigan

 **Pennsylvania** — The University of Pennsylvania

 **ANL** — Argonne National Laboratory

 **Santa Cruz-SLAC-Stanford DES Consortium**

- **Santa Cruz** - University of California Santa Cruz
- **SLAC** - SLAC National Accelerator Laboratory
- **Stanford** - Stanford University

 **ETH** Eidgenössische Technische Hochschule Zürich
Swiss Federal Institute of Technology Zurich

 **ETH-Zuerich** — Eidgenoessische Technische Hochschule Zuerich

