

New results for $\nu_\mu \rightarrow \nu_e$ oscillations in MINOS



Ryan Patterson
Caltech

Joint Experimental-
Theoretical Seminar

FNAL

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Absolute neutrino mass?

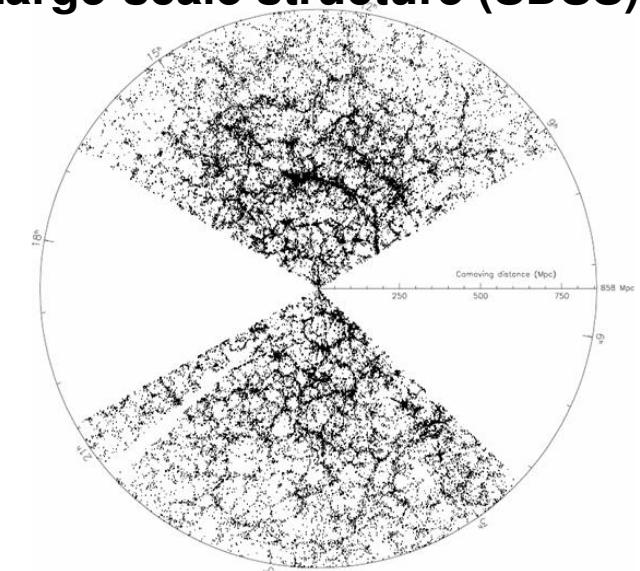
Unmeasured, but known **small**:

→ **less than ~ 0.7 eV**

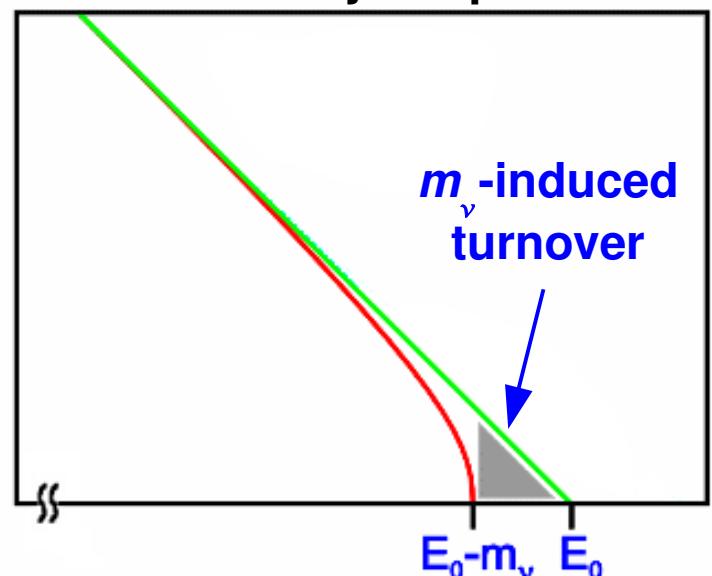
(sum of ν masses inferred from cosmological observations)

→ **less than ~ 2 eV**

(direct kinematic measurements of ν_e effective mass)



Beta decay endpoint:



Or, notably: $m_\nu \ll m_{EW}$

“See-saw” mechanism puts the physical neutrino mass at:

$$m_\nu \sim \frac{m_{EW}^2}{m_{GUT}}$$

Neutrino sector may provide a window to new physics

Neutrino mixing

mass eigenstates \neq flavor eigenstates:

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i}^* |\nu_i\rangle$$

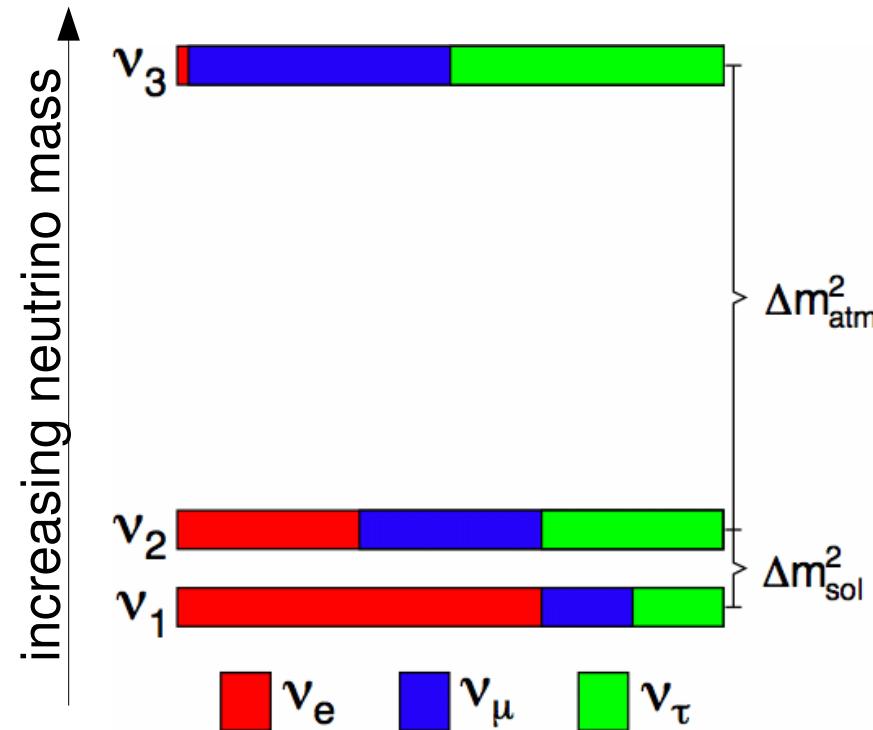
$$\alpha = (e, \mu, \tau)$$

flavor composition of the neutrino
can **change as it propagates**:

$$P(\nu_\alpha \rightarrow \nu_\beta) = |\langle \nu_\beta | \nu_\alpha(L) \rangle|^2$$

$$= \delta_{\alpha\beta} - 4 \sum_{i>j} \Re(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2[1.27 \Delta m_{ij}^2 L/E]$$

$$+ 2 \sum_{i>j} \Im(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2[2.54 \Delta m_{ij}^2 L/E]$$



two-neutrino case: $P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2(2\theta) \sin^2(1.27 \Delta m^2 L/E)$

$$\left(1.27, 2.54 \text{ in units of } \frac{\text{GeV}c^4}{\text{eV}^2\text{km}} \right)$$

Neutrino mixing

Two oscillation regimes:

- **solar, reactor experiments:**

$$\Delta m_{\text{sol}}^2 = (7.6 \pm 0.2) \times 10^{-5} \text{ eV}^2$$

- **atmospheric, accelerator experiments:**

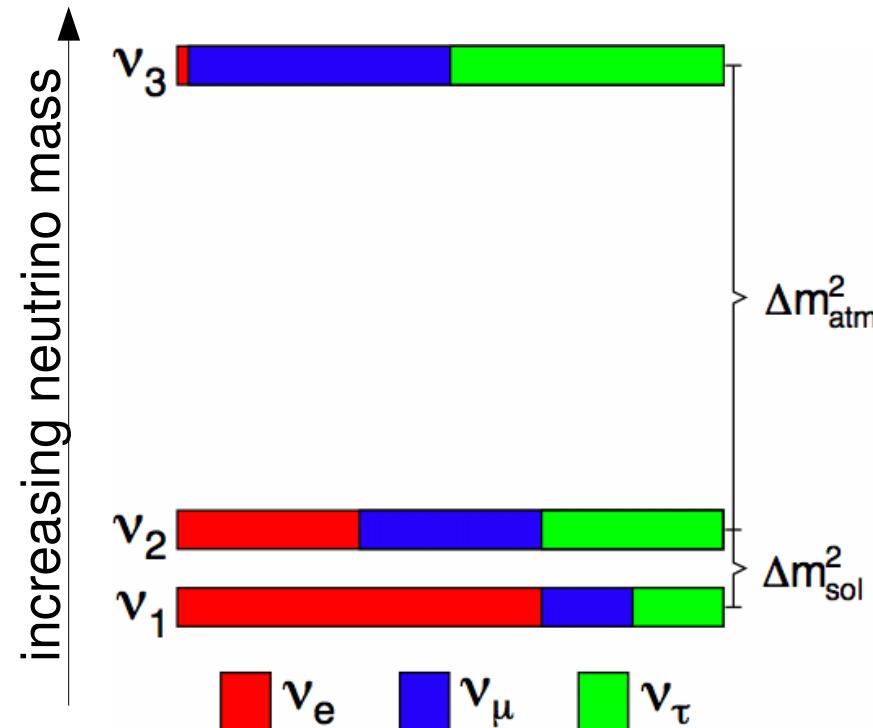
$$\Delta m_{\text{atm}}^2 = (2.43 \pm 0.13) \times 10^{-3} \text{ eV}^2$$

Mixing angles extracted so far...

$$\theta_{12} = 34^\circ \pm 3^\circ$$

$$\theta_{23} = 45^\circ \pm 5^\circ$$

$$\theta_{13} < 11^\circ \text{ (90% C.L.)}$$



$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Accessible by MINOS
 Solar sector
 Majorana phases

Many questions

- Why is θ_{23} near maximal?
- What is θ_{13} and why is it small?
- Why so different than quark mixing?

Quarks:

$$U_{\text{CKM}} \sim \begin{pmatrix} 1 & 0.2 & 0 \\ 0.2 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

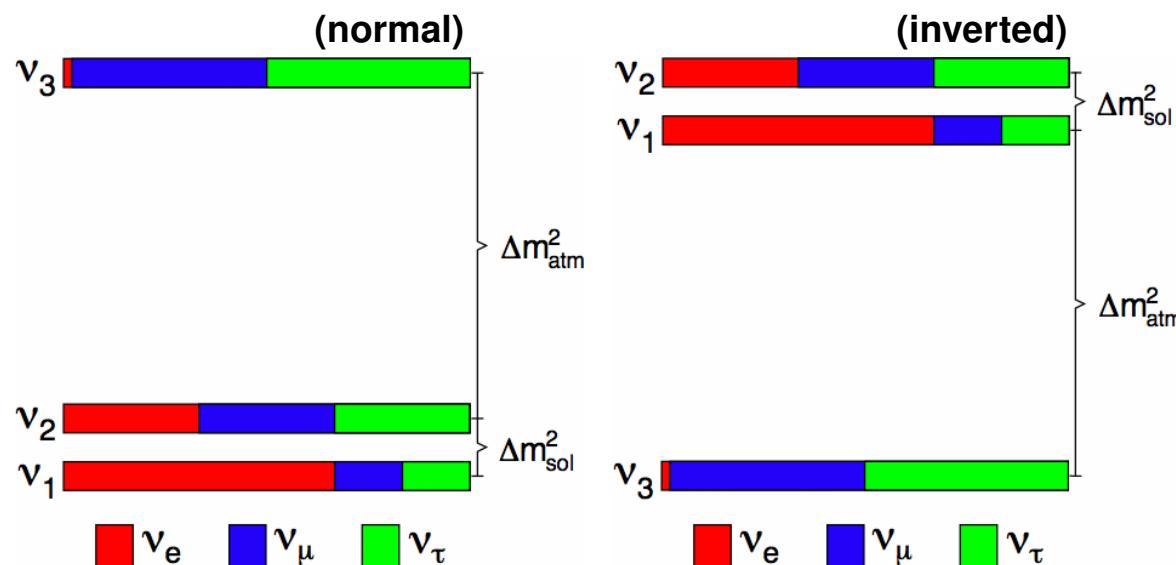
Neutrinos:

$$U_{\text{MNS}} \sim \begin{pmatrix} 0.8 & 0.6 & 0 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$$

*underlying symmetries,
unification implications...*

- Is CP violation present?
- Are neutrinos Majorana?
- What is the hierarchy of neutrino masses?

*leptogenesis, interpreting
CMB & LSS observations, $0\nu\beta\beta$*



Measuring θ_{13}

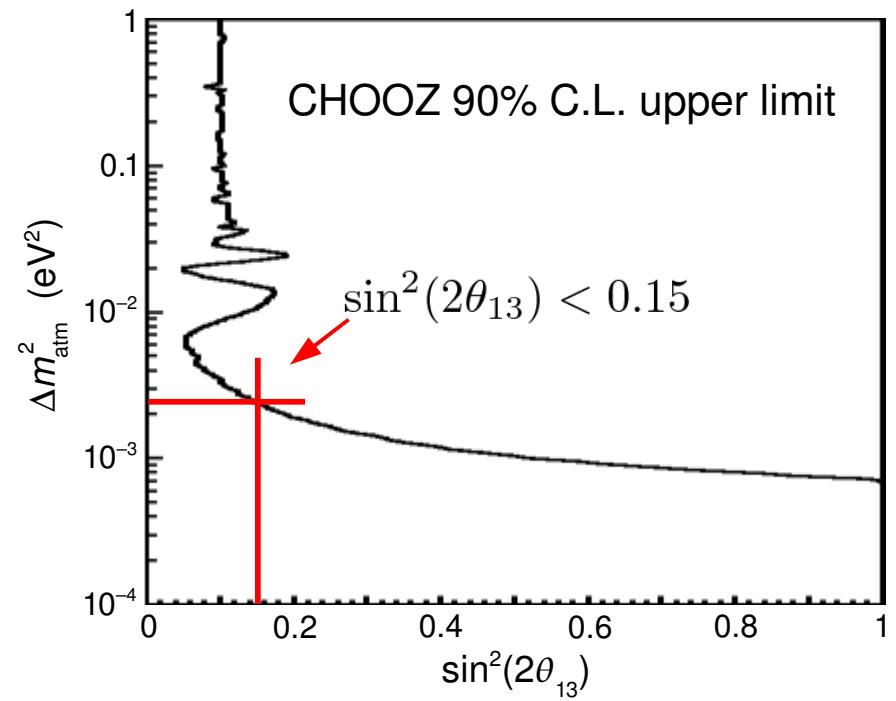
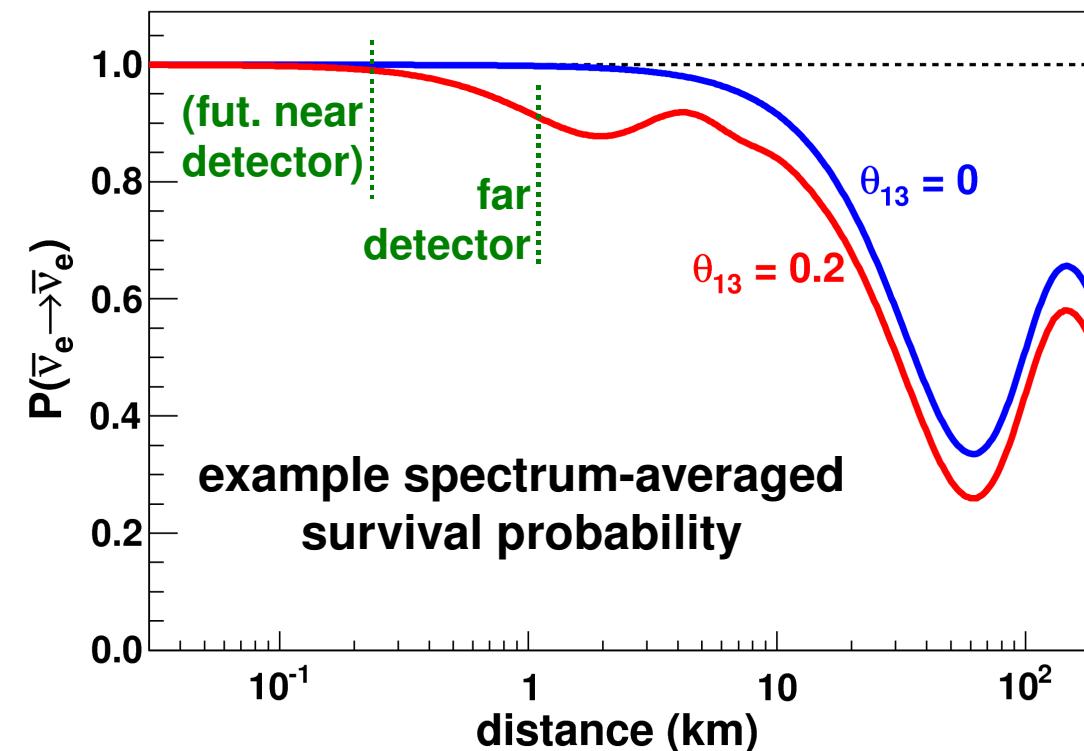
- Need a neutrino source: **accelerators or nuclear reactors**
- **CHOOZ reactor experiment** holds best limit
- ***Reactor principle:***

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \approx 1 - \frac{\sin^2(2\theta_{13}) \sin^2(1.27 \Delta m_{\text{atm}}^2 L/E)}{\text{plus smaller solar term}}$$

plus smaller
solar term

$L \approx 2 \text{ km}$

$E \approx 4 \text{ MeV}$



Measuring θ_{13}

- A conventional **accelerator-based** neutrino source provides ν_μ
- Look for **appearance** of ν_e

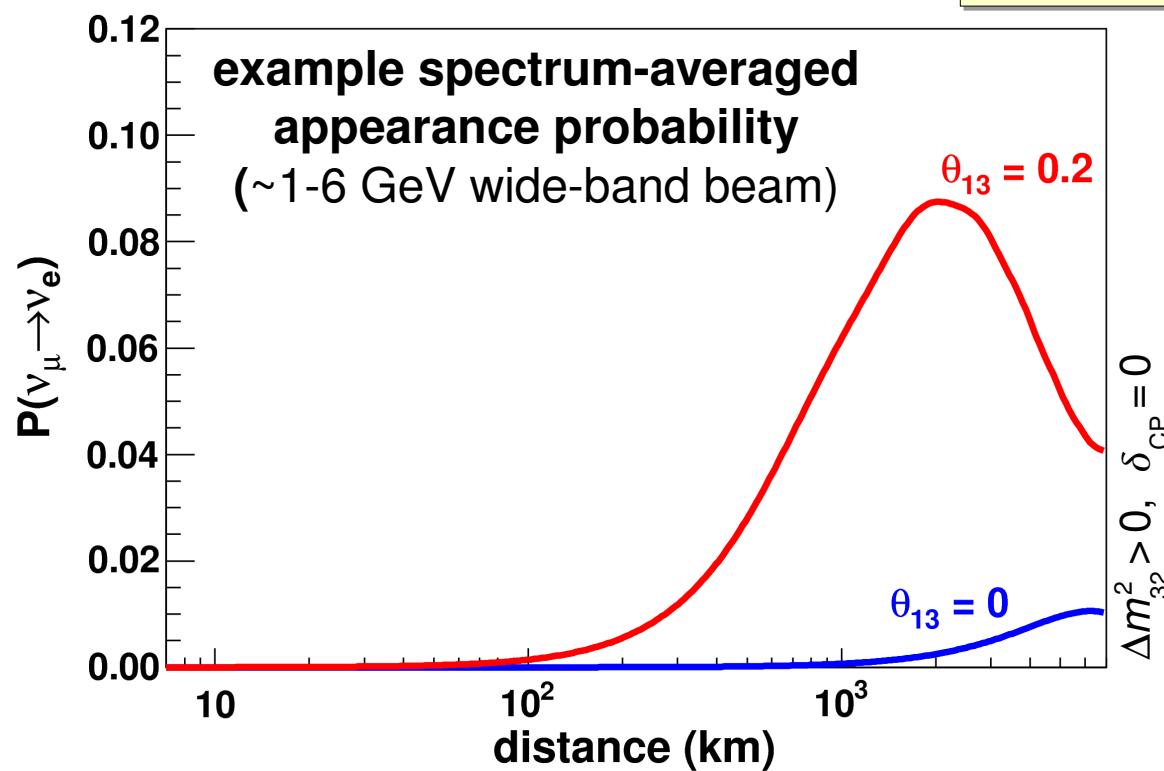
$$P(\nu_\mu \rightarrow \nu_e) \approx$$

$$\sin^2(2\theta_{13}) \sin^2(\theta_{23}) \sin^2(1.27 \Delta m_{\text{atm}}^2 L/E)$$

plus potentially large modifications (CPV, matter effects)

250 - 2500 km

0.5 - 5 GeV



Disappearance (reactor):

$$\theta_{13}$$

Appearance (accelerator):

$$\theta_{13}, \theta_{23}, \text{CP}, \text{sign}(\Delta m_{32}^2)$$

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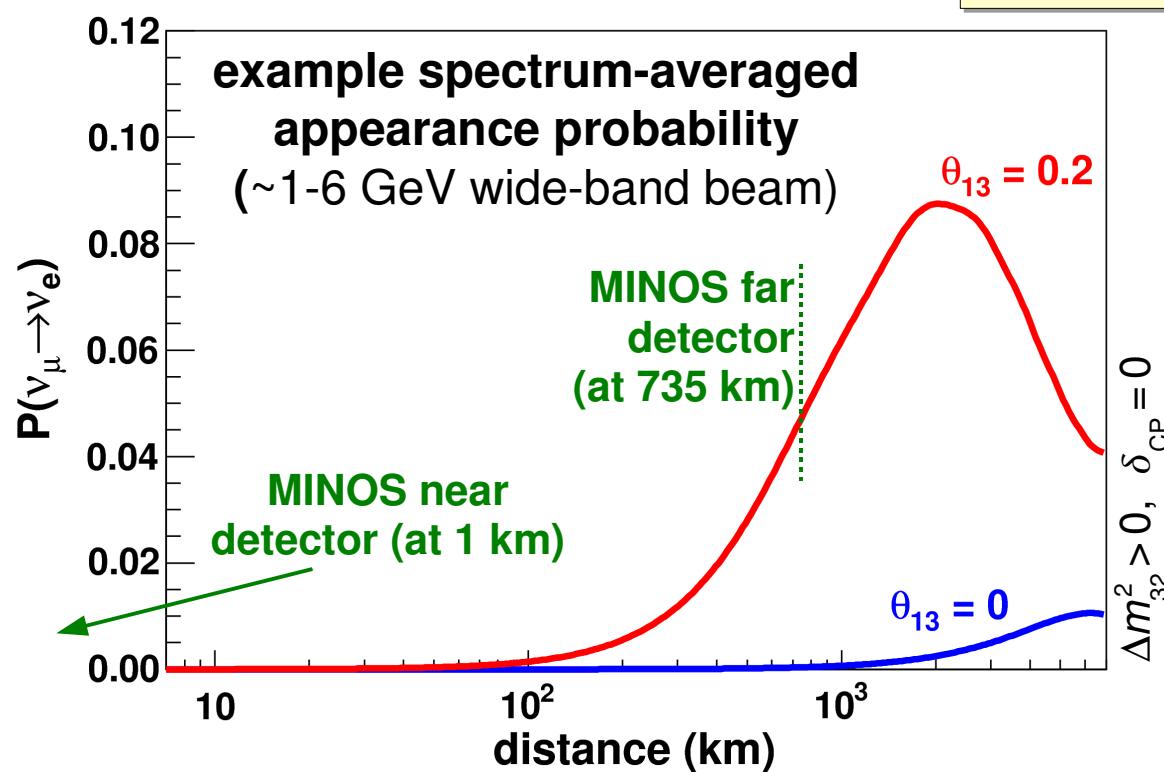
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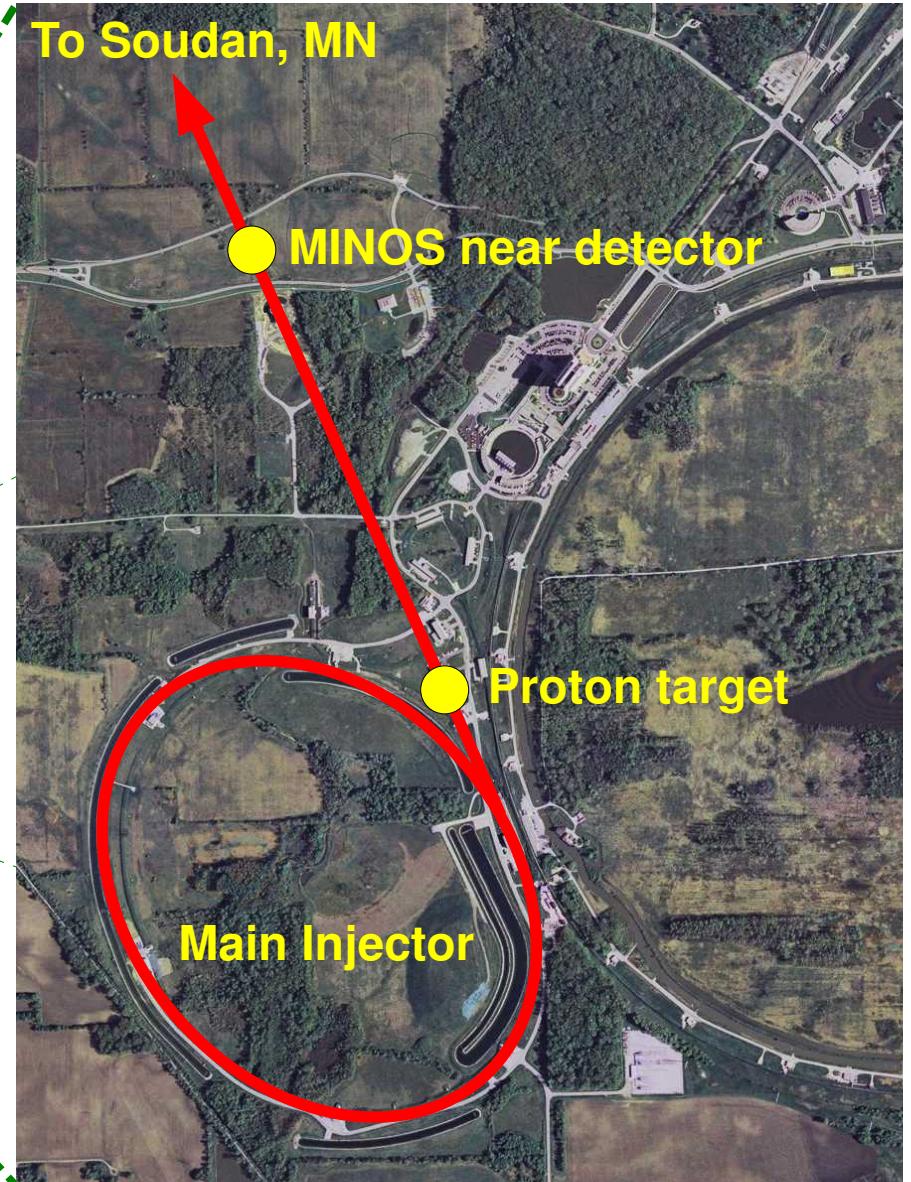
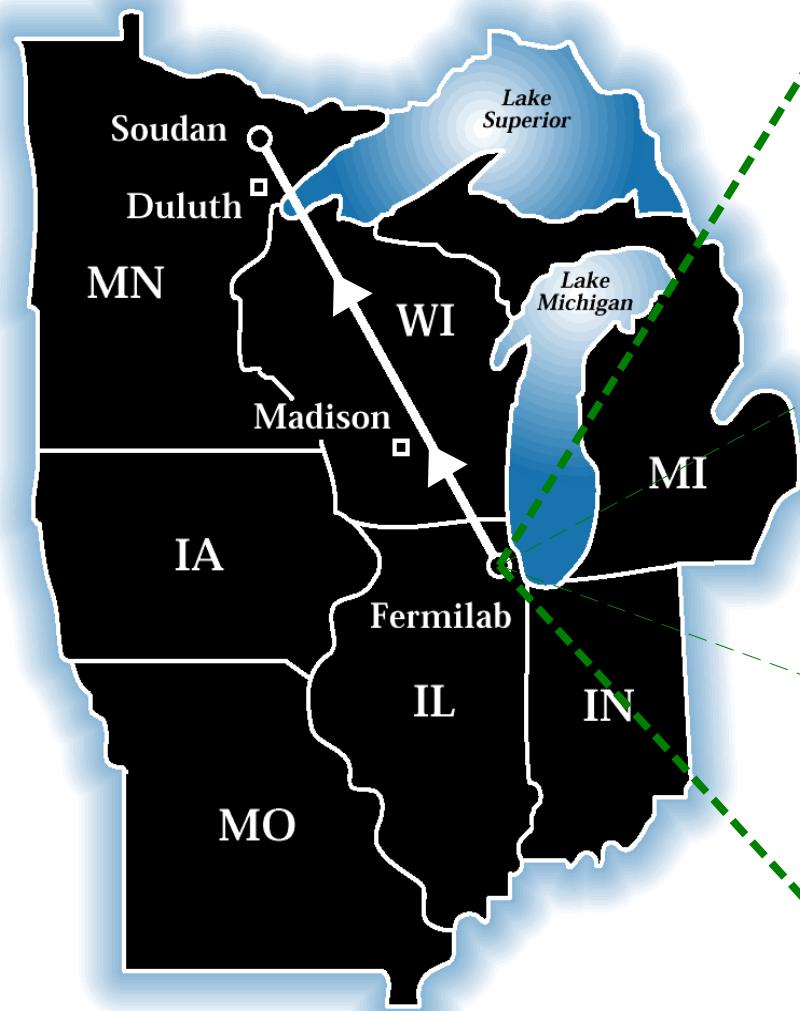
Appearance (accelerator):

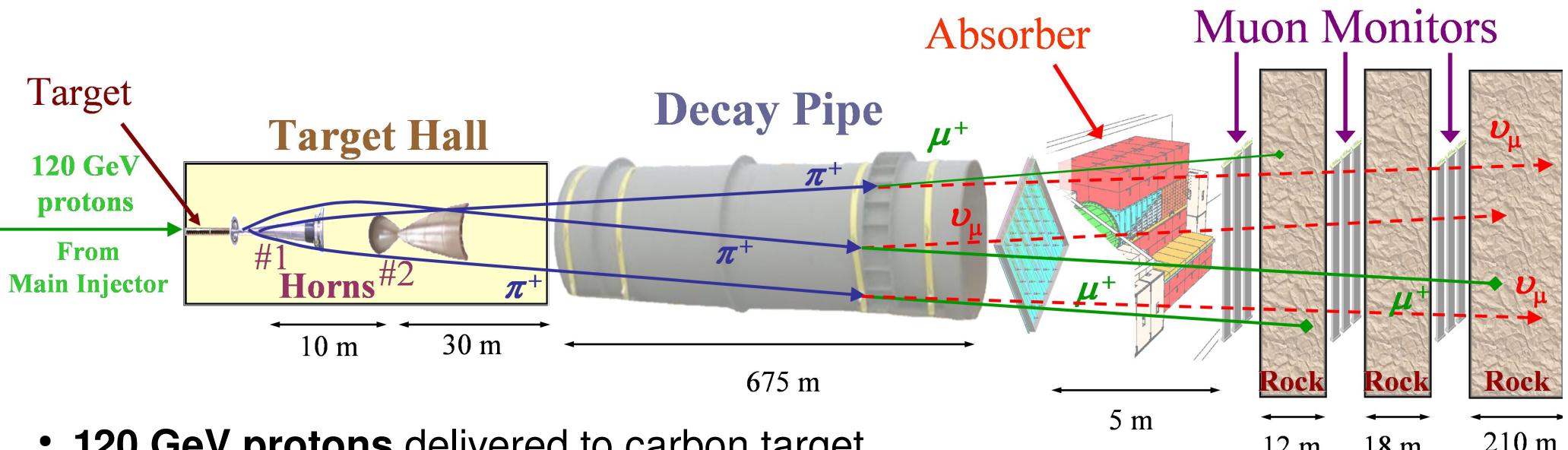
$$\theta_{13}, \theta_{23}, \text{CP}, \text{sign}(\Delta m_{32}^2)$$

$\nu_\mu \rightarrow \nu_e$ in MINOS

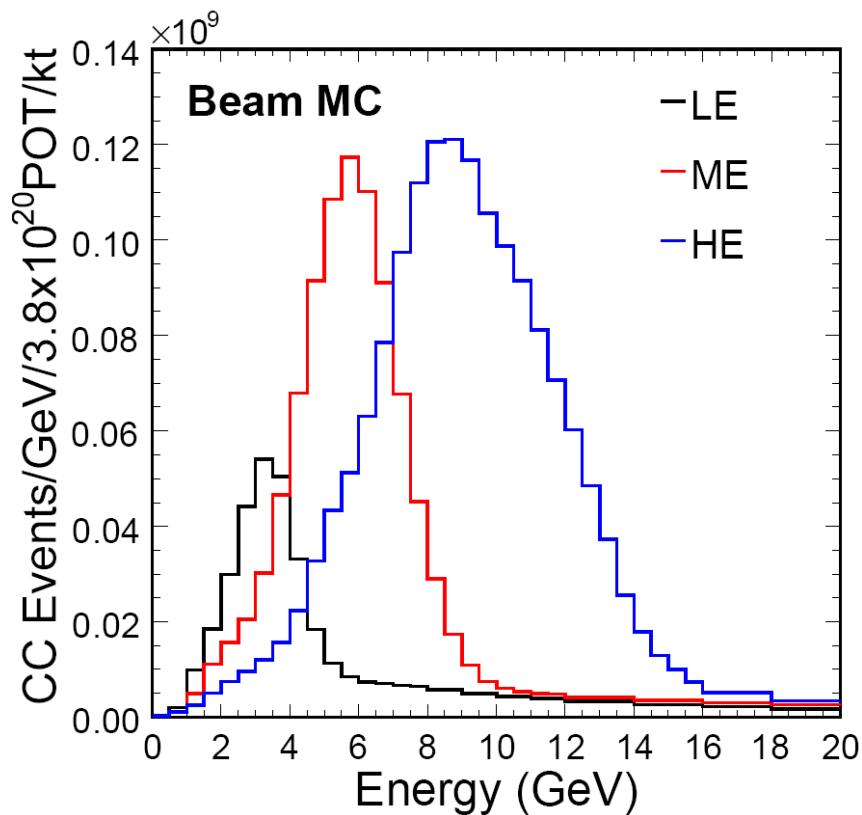
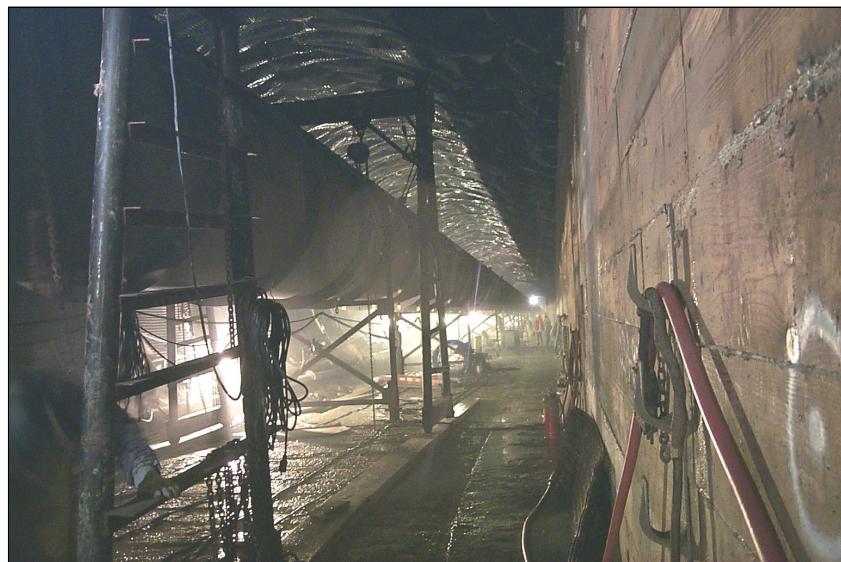
Main Injector Neutrino Oscillation Search (MINOS)

735 km to the far detector



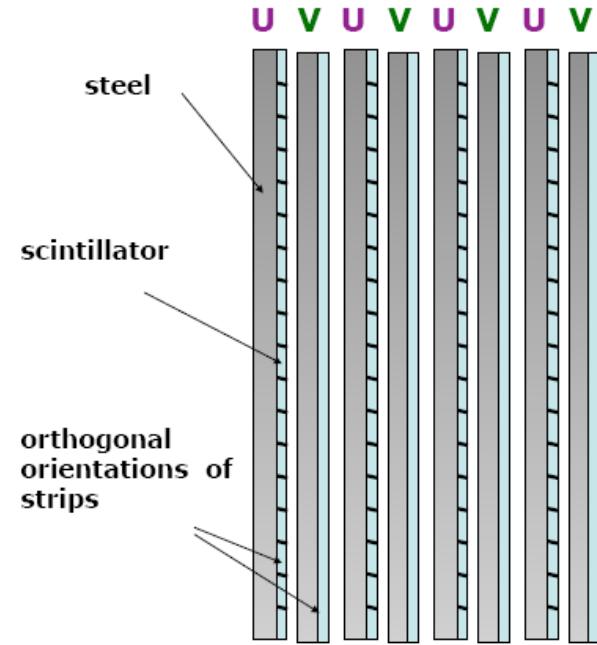


- **120 GeV protons** delivered to carbon target
- Magnetic **focusing horns** send positive secondaries (π^+ mostly) down **decay pipe**
- Target and horn positions are adjustable:
tunable neutrino energy spectrum →



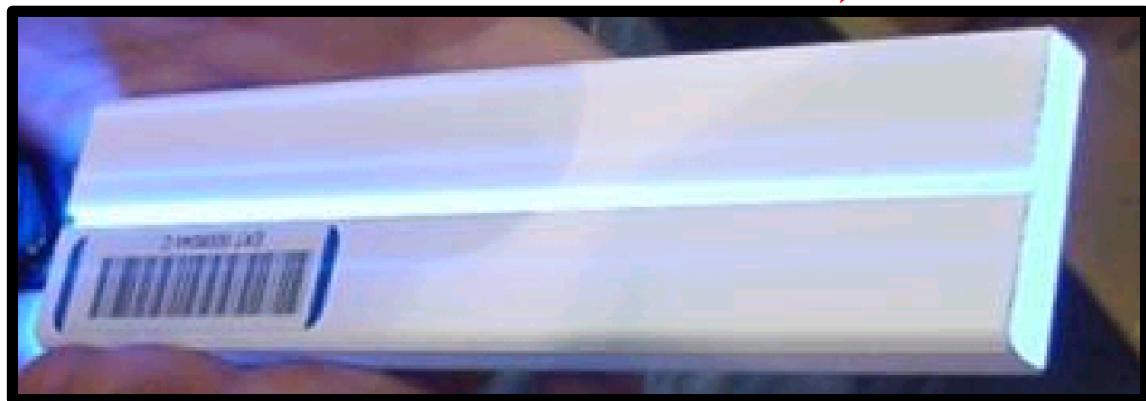
MINOS Detectors

- Near and far detectors:
Magnetized tracking calorimeters (~ 1.3 T field)
- Alternating layers of
steel (1" thick plates)
scintillator (1 cm thick, 4.1 cm wide strips)
- Scintillator layers oriented at $\pm 45^\circ$

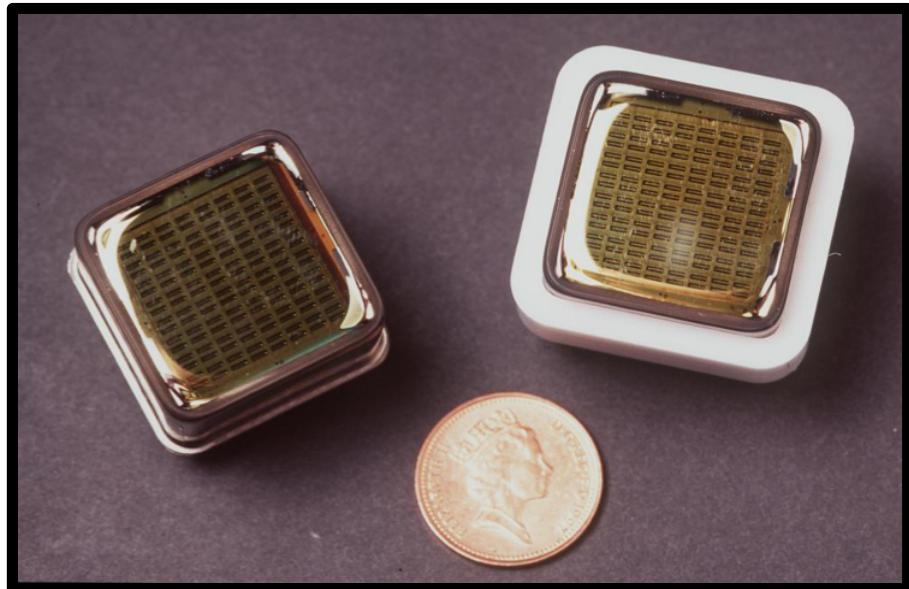
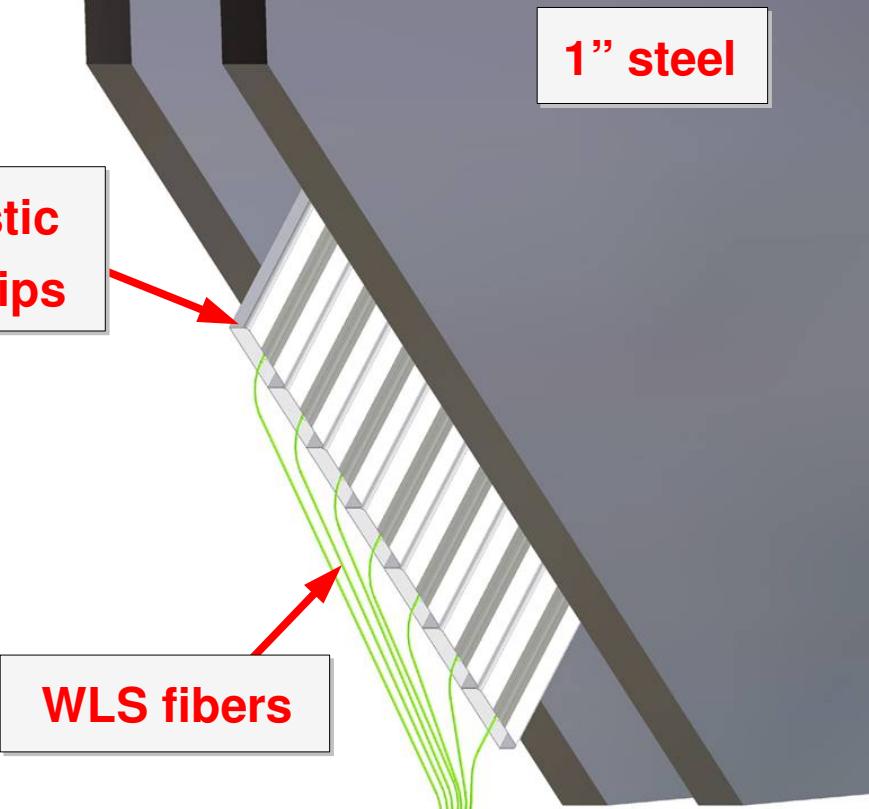


1" steel

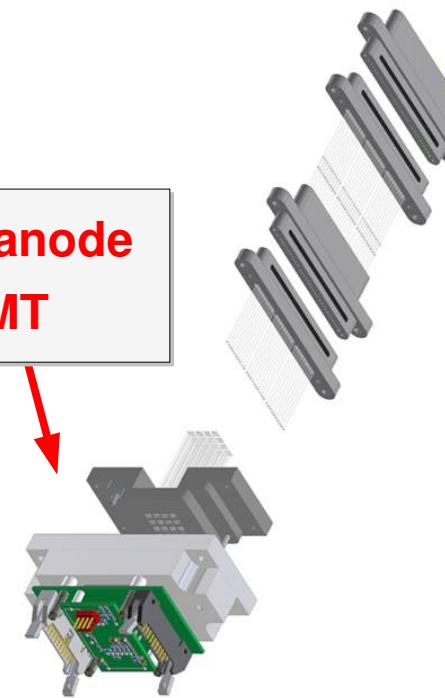
MINOS Detectors



Extruded plastic
scintillator strips

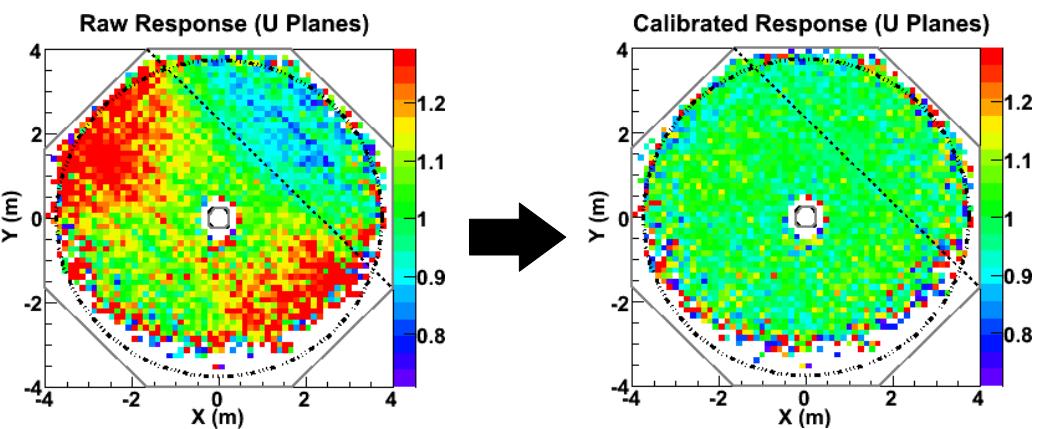


Multi-anode
PMT



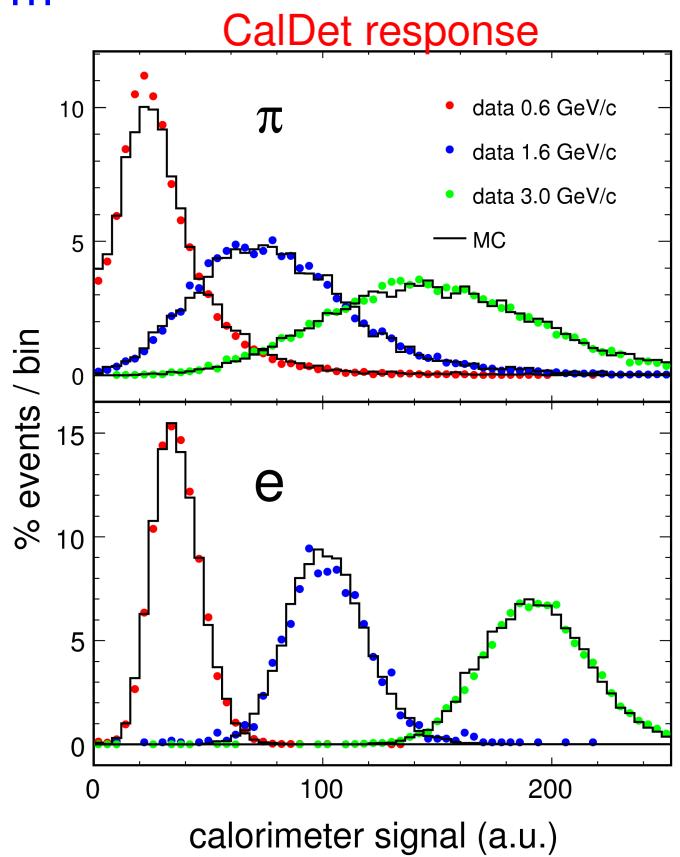
Calibration

Cosmic ray muons:
measure & remove **spatial variations**
(channel differences, attenuation)

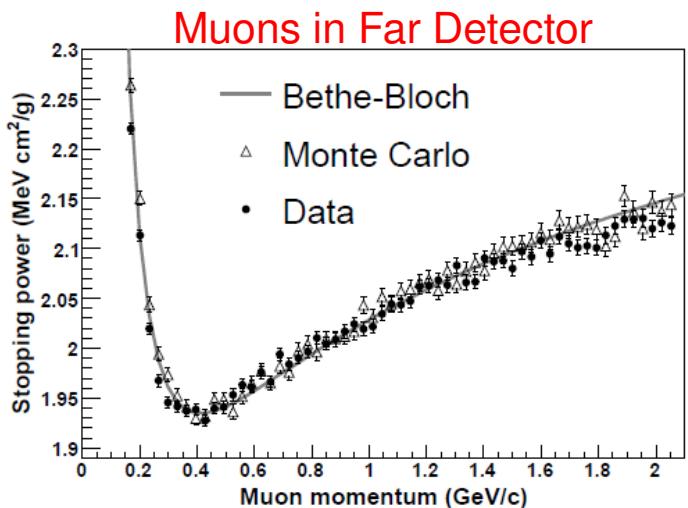


MINOS Calibration Detector (CalDet): E scale

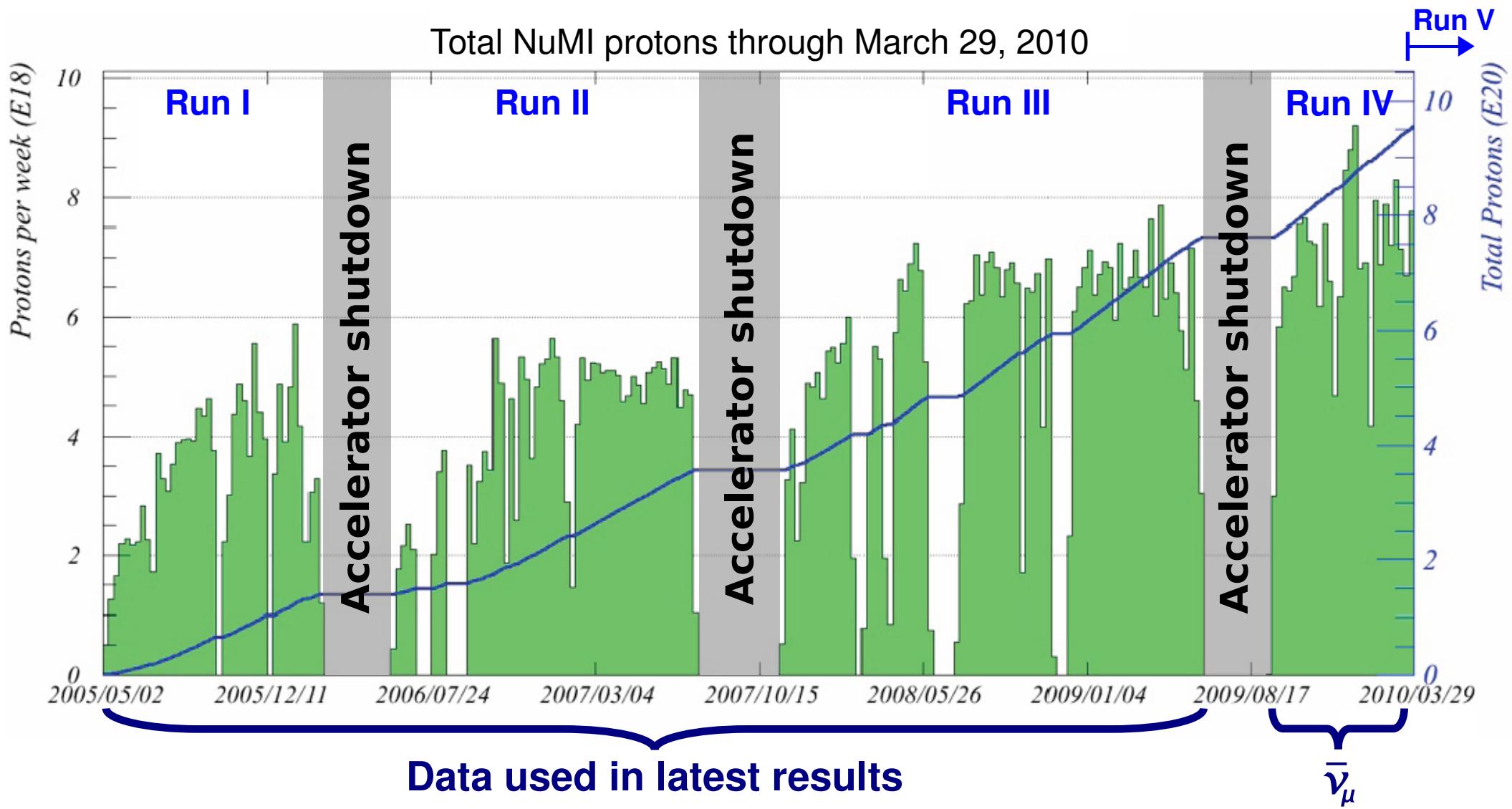
- Exposed to **0.2–10 GeV p , e , μ , π** at CERN
- 60 planes, 1 m x 1 m



**Everything tied together
with **stopping muons**:**

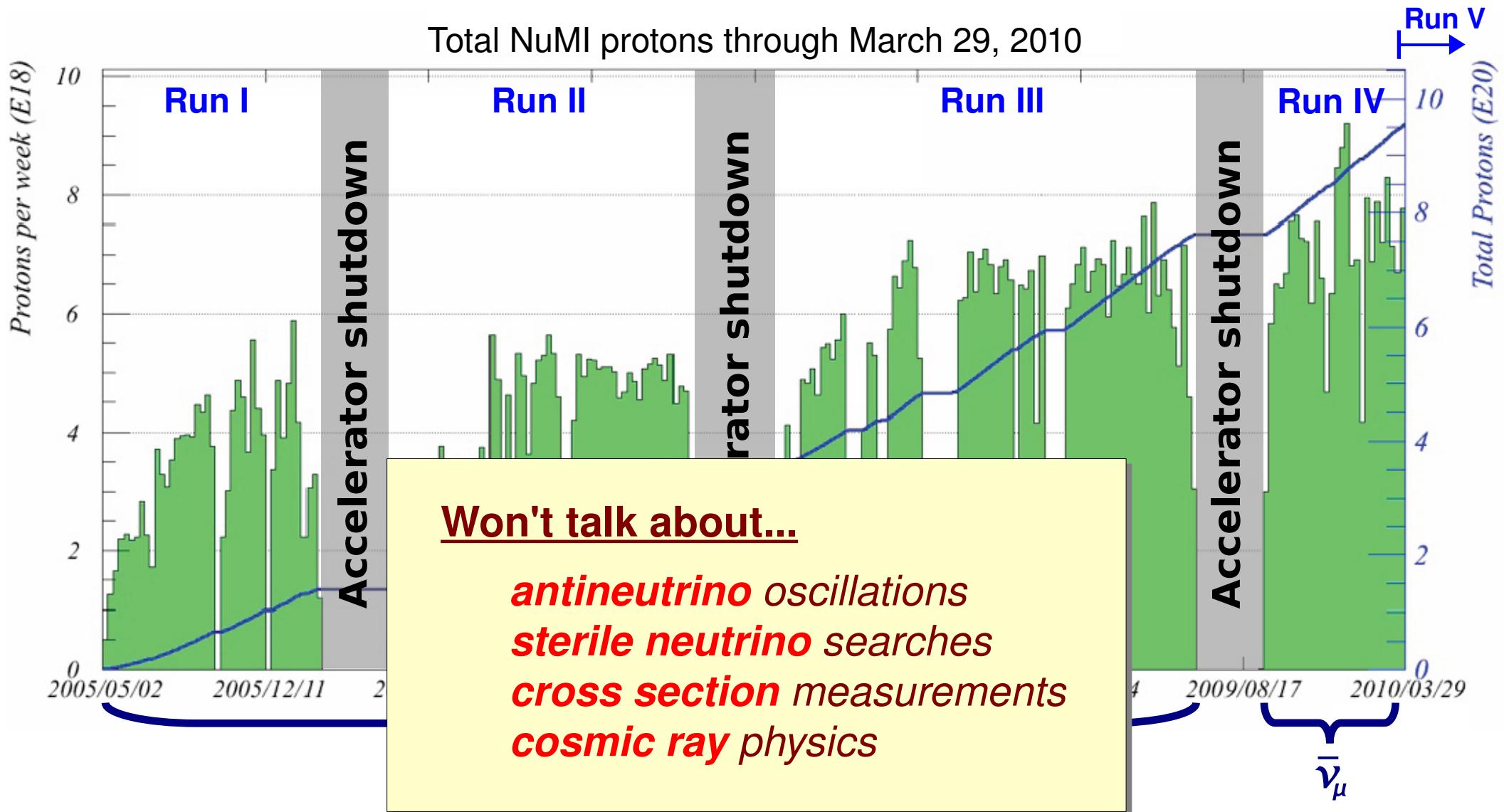


MINOS calorimetry good to:
6% (absolute)
2% (relative near/far)



Excellent NuMI intensity over past two years

ν_e appearance results through Run III ($\sim 7 \times 10^{20}$ p.o.t.)



Excellent NuMI intensity over past two years

ν_e appearance results through Run III ($\sim 7 \times 10^{20}$ p.o.t.)

Neutrino events

ν_μ charged current

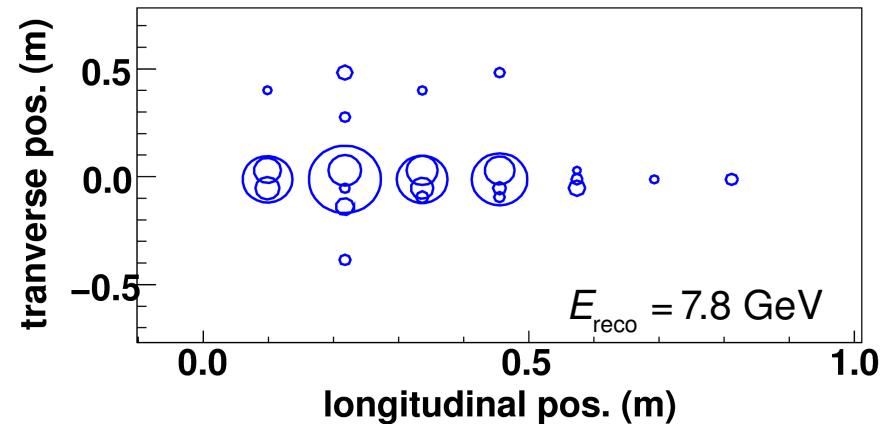
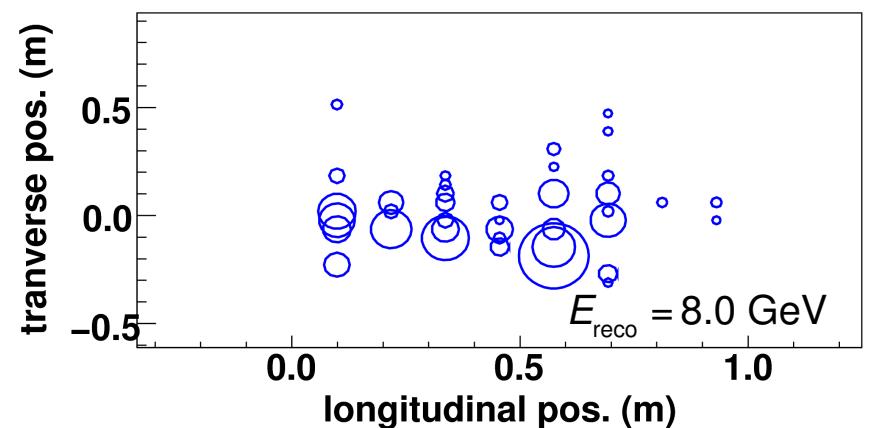
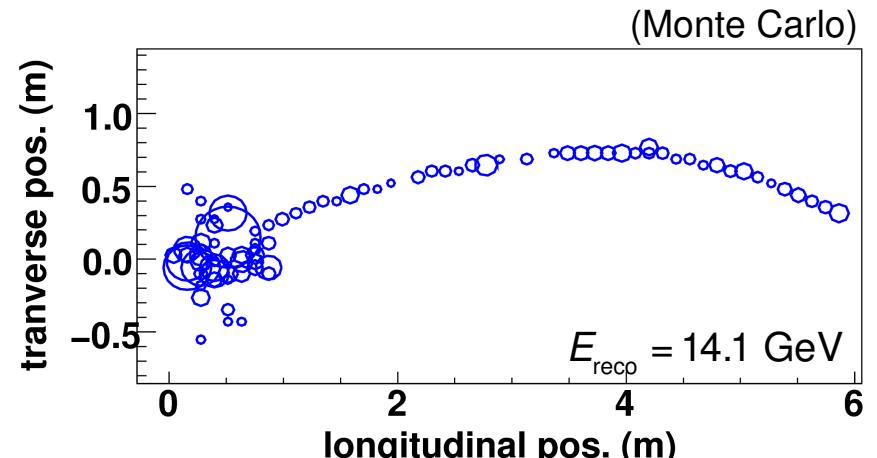
- Clear signature in MINOS: *long track*
- If μ track is very short, event can be mistaken for ν_e CC

neutral current

- Esp. with π^0 , hard to distinguish from ν_e CC
- Energy more transversely distributed

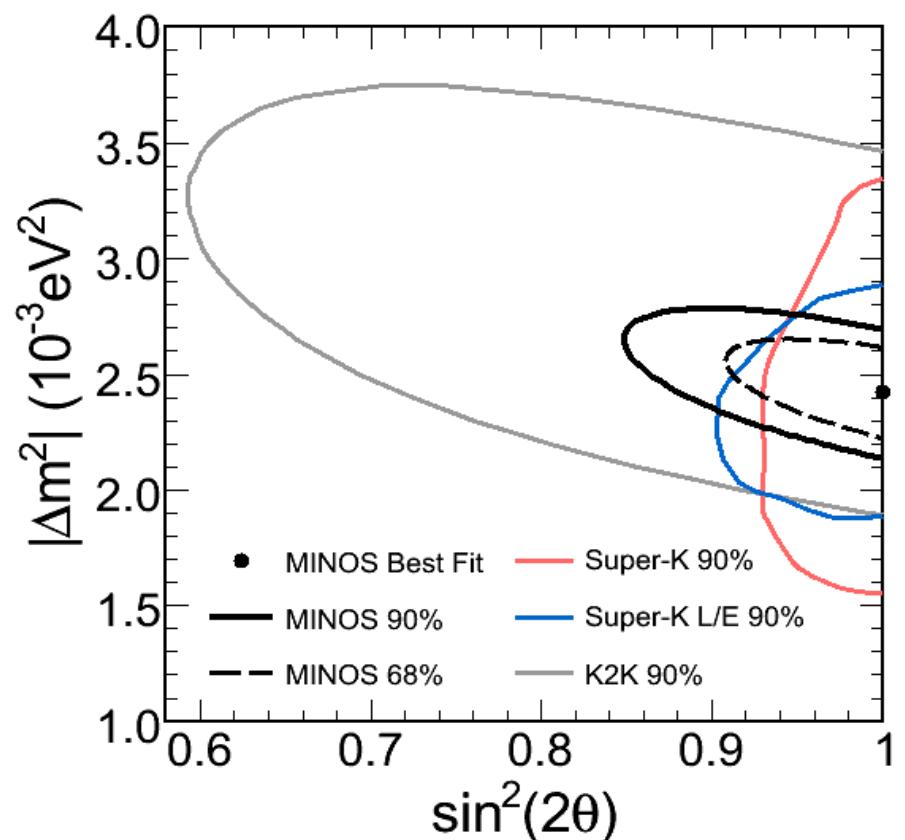
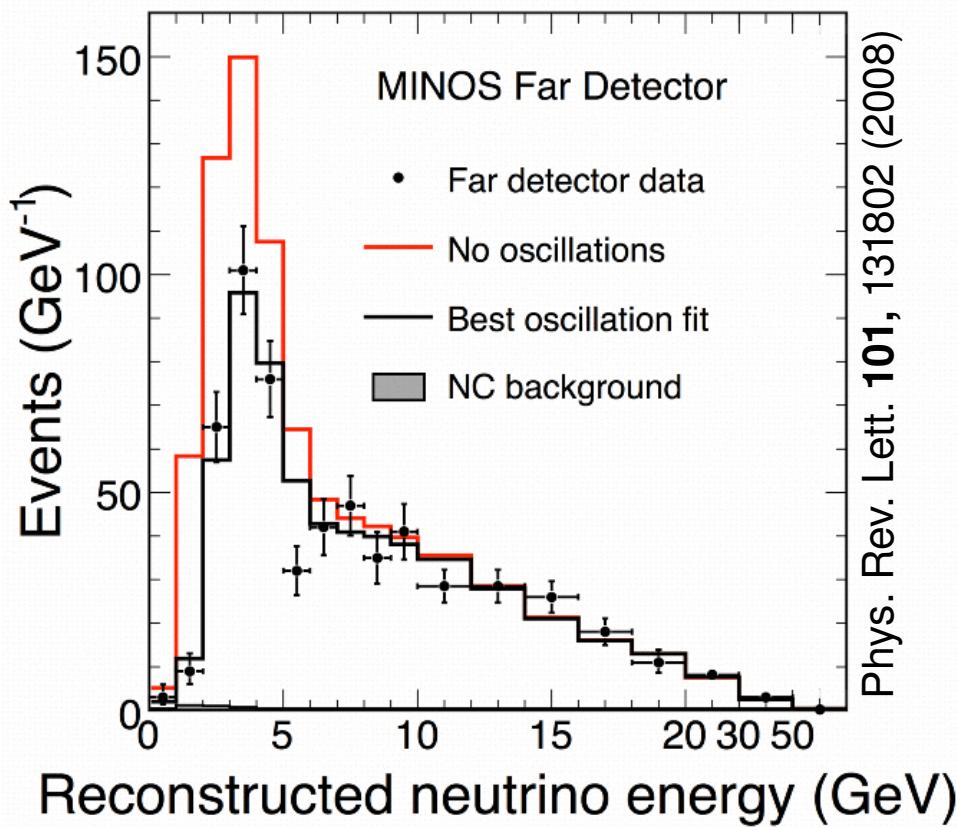
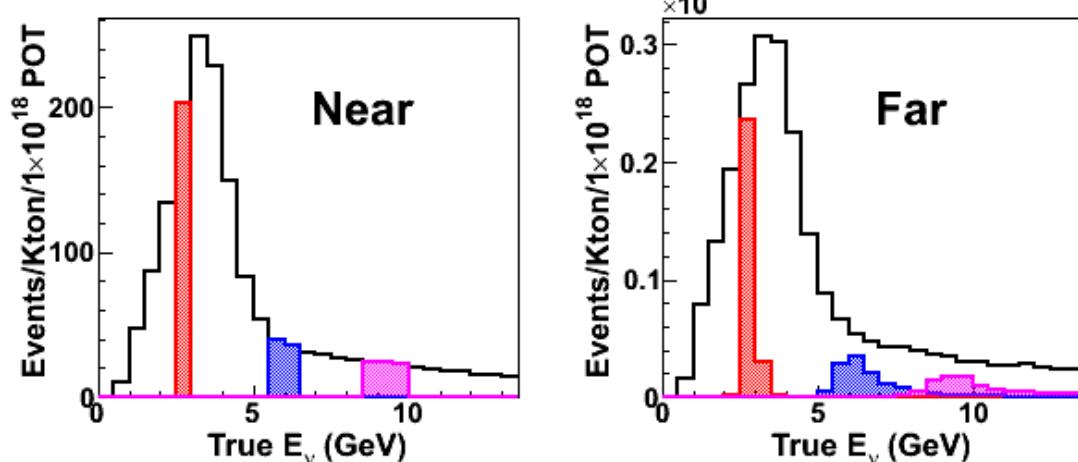
ν_e charged current [signal]

- ν_e is small component of initial flux
- Electron leaves characteristic deposition pattern (**compact shower**)



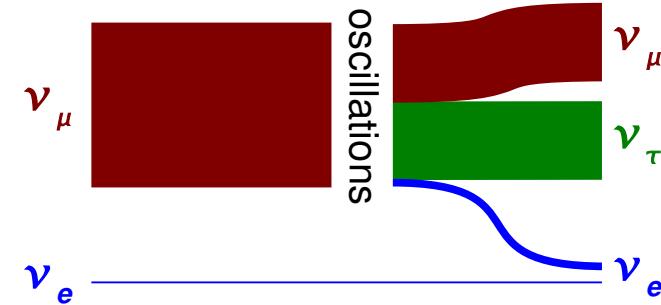
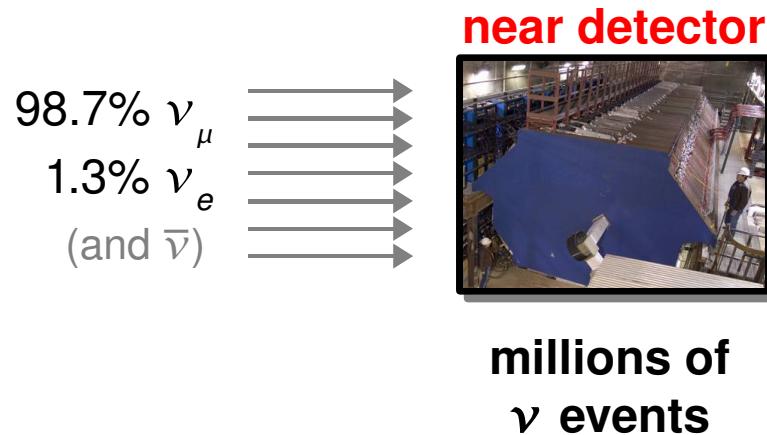
First, ν_μ survival

- Measure ν_μ CC interactions in the **near detector**
- **Transport** observed spectrum to create far detector **prediction**
- **Energy-dependent deficit** in far detector data $\Rightarrow \theta_{23}, |\Delta m^2_{32}|$



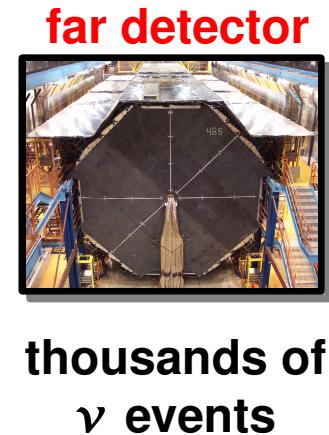
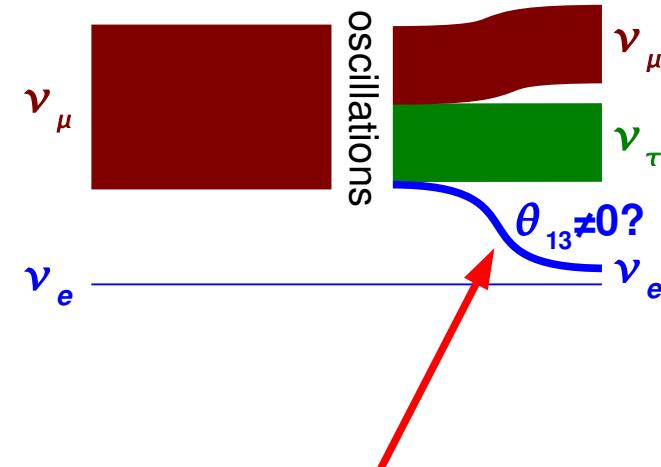
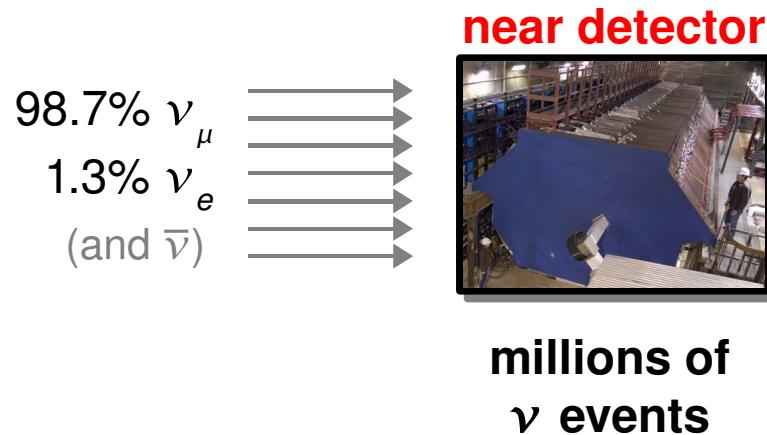
ν_e appearance

→ travel 735 km →



ν_e appearance

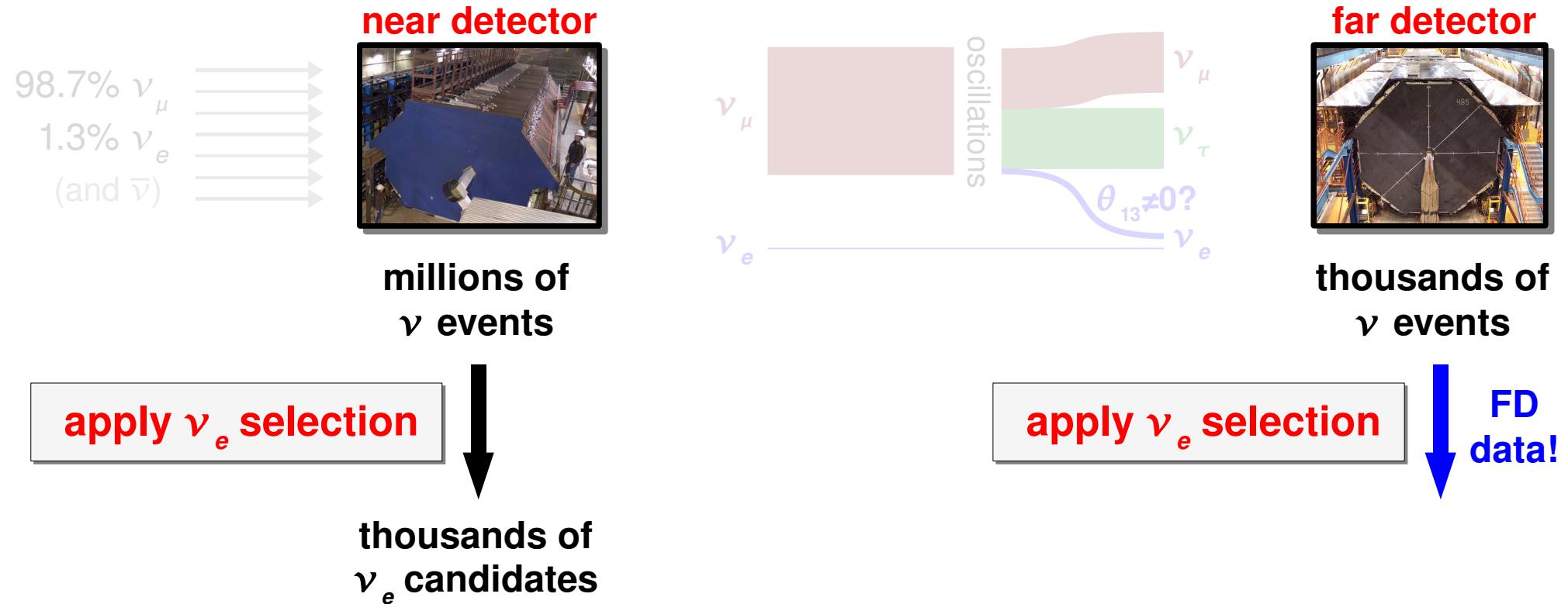
→ travel 735 km →



Looking for this!

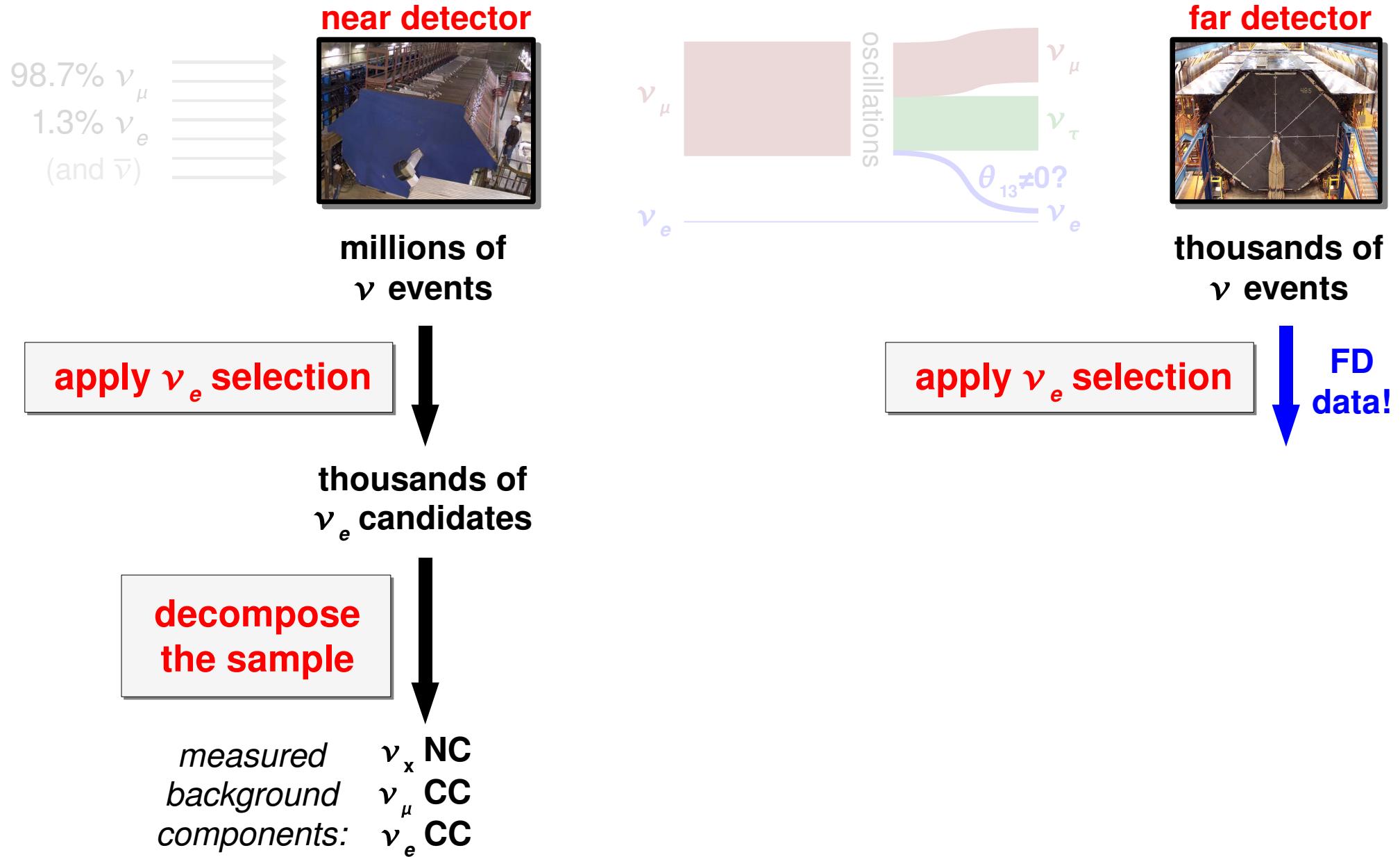
ν_e appearance

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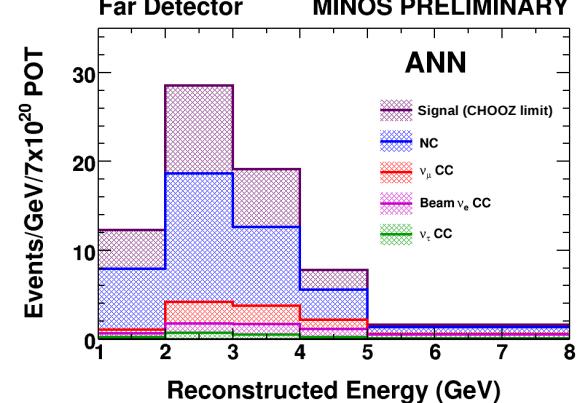
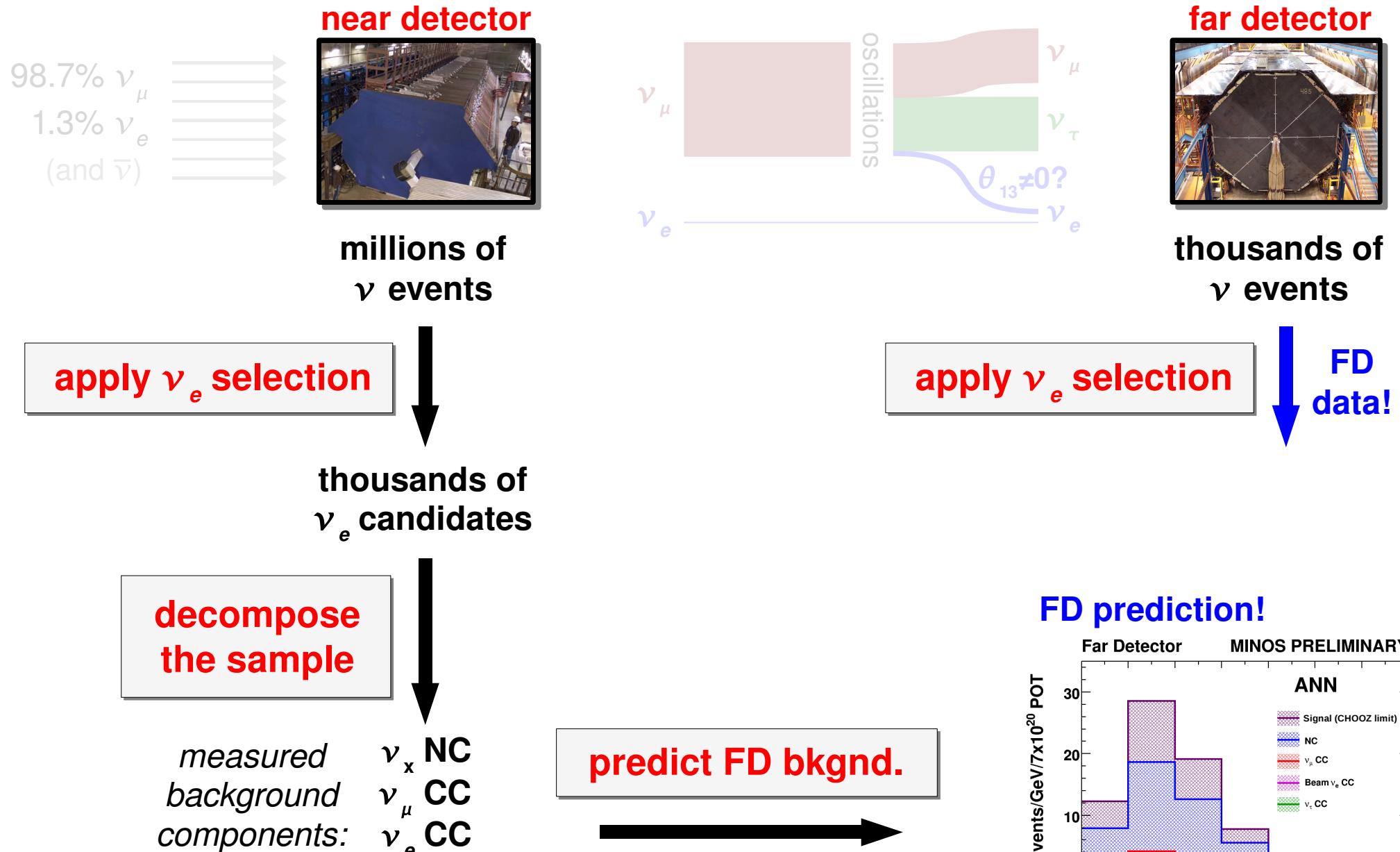
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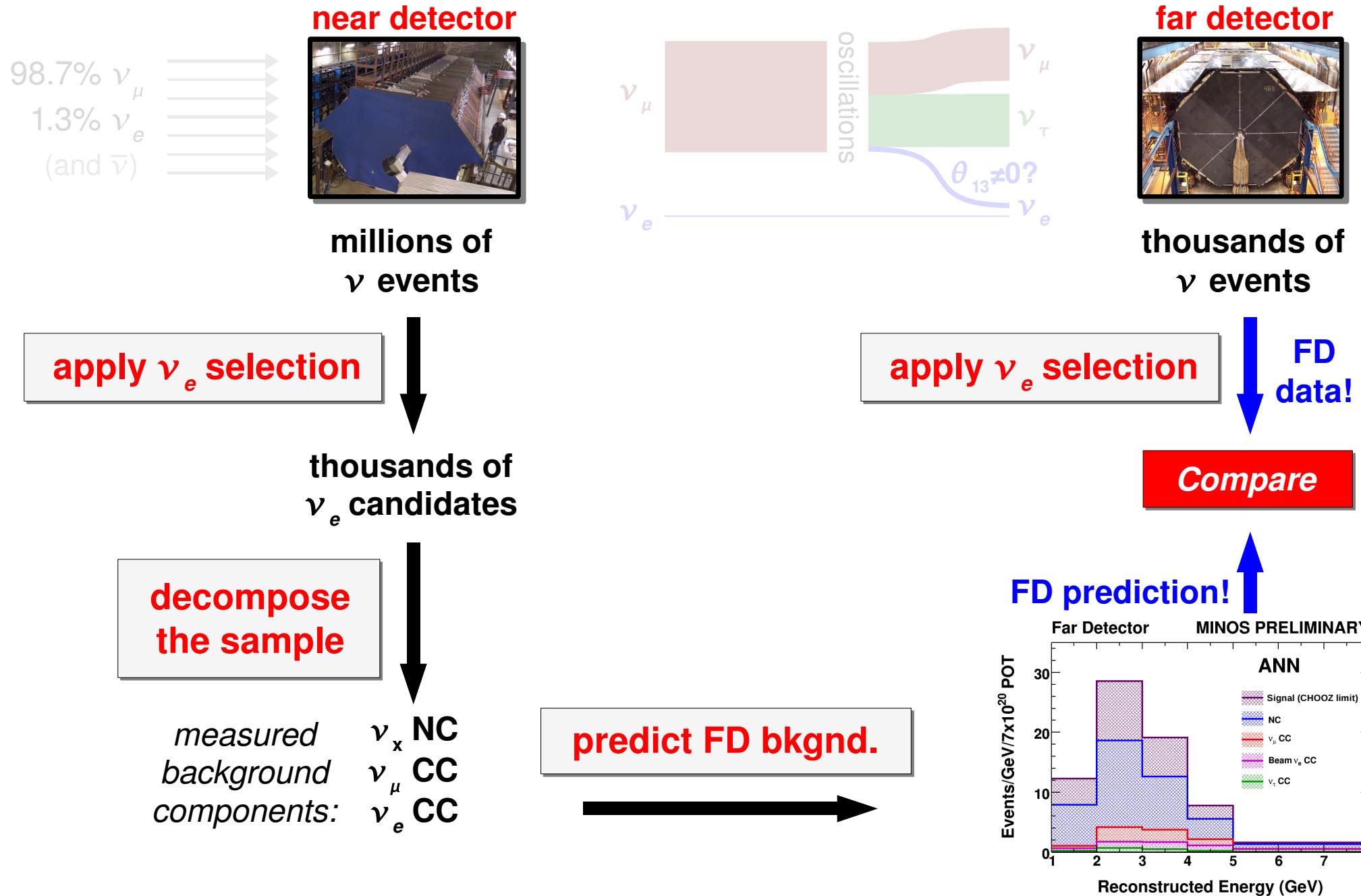
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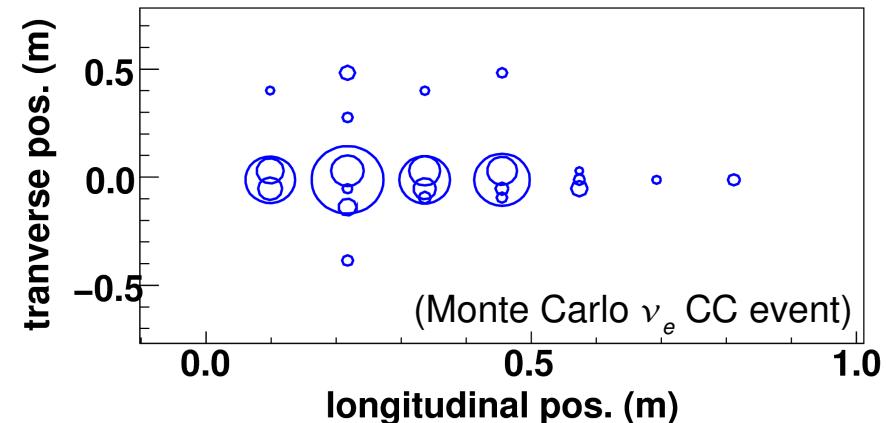


Preliminary cuts

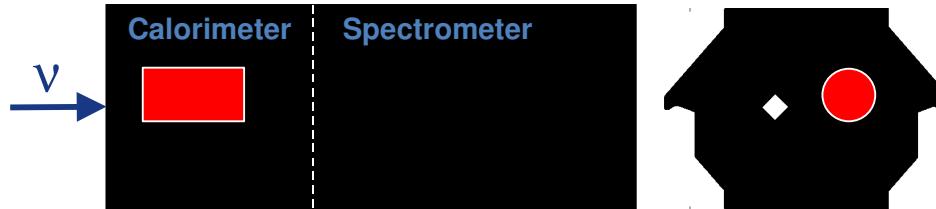
- Selecting ν_e CC candidates: start with some *preliminary cuts*...

1. Cleanly contained event vertex

- Tell-tale tracks can be seen
- Ensure reasonable calorimetry

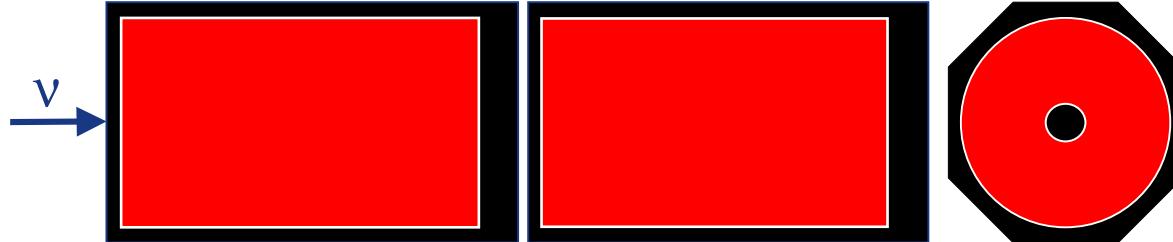


Near det.



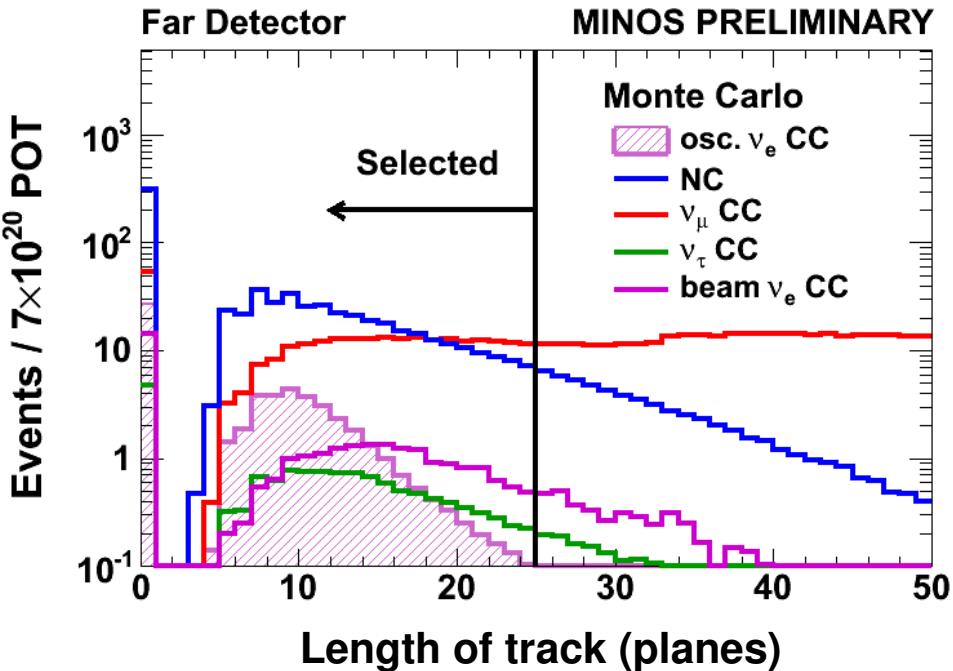
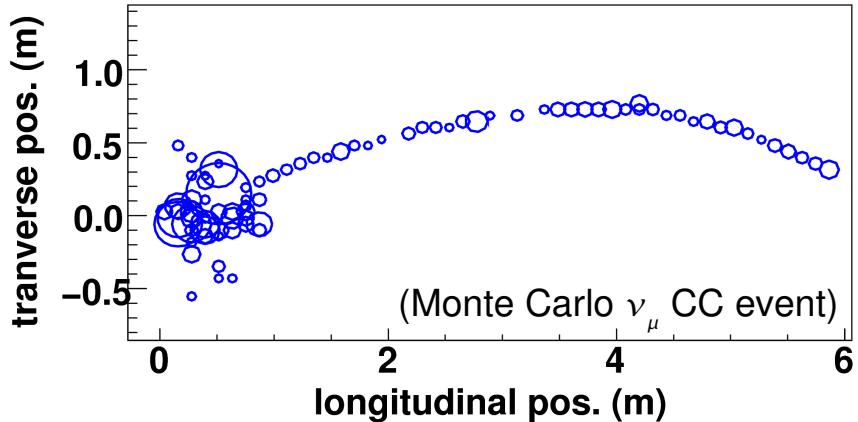
Fiducial regions in red

Far det.



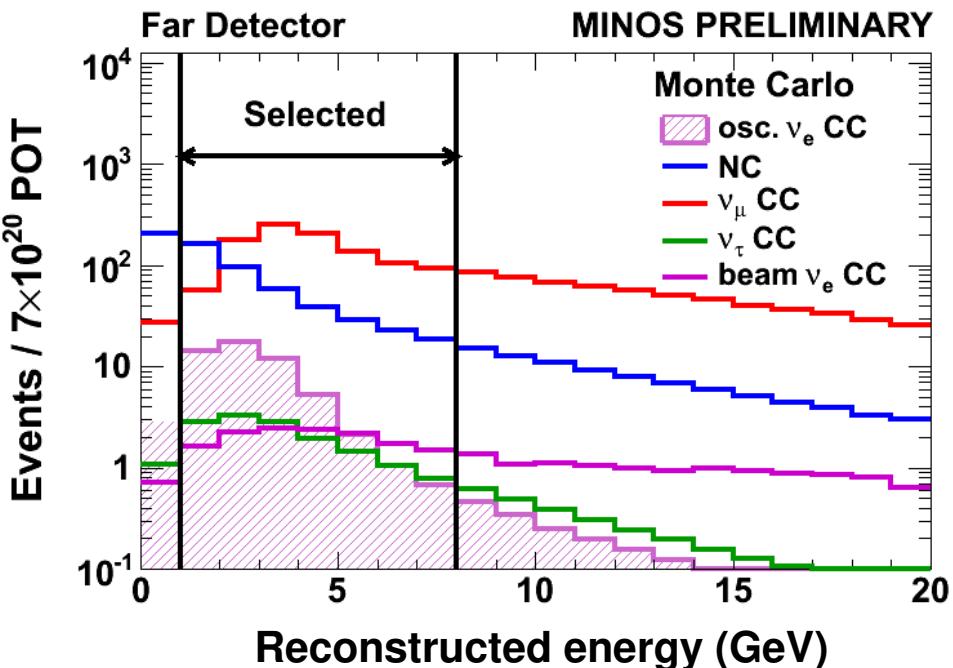
2. No long tracks

- Removes majority of ν_μ CC events



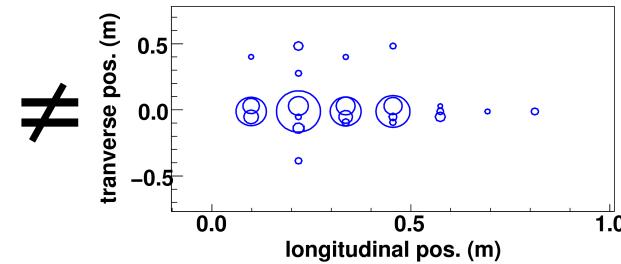
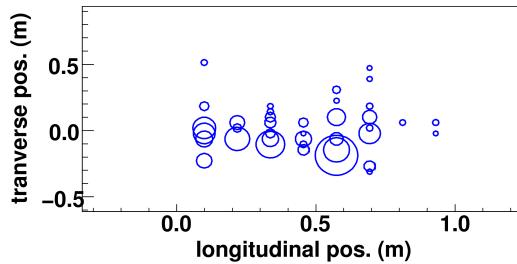
3. Appropriate energy

- No signal expected above 8 GeV
- Minimal signal and lots of NC background below 1 GeV

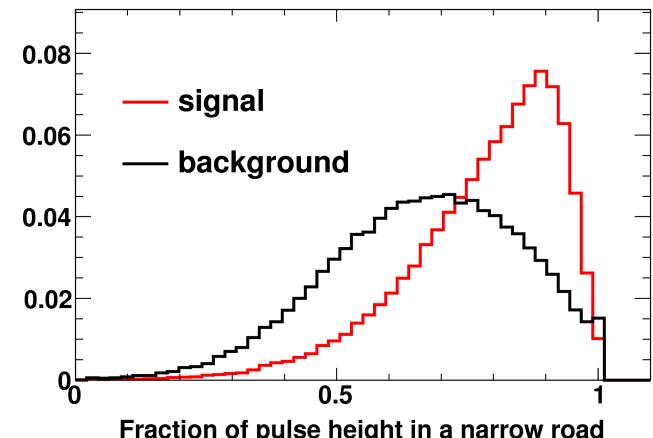
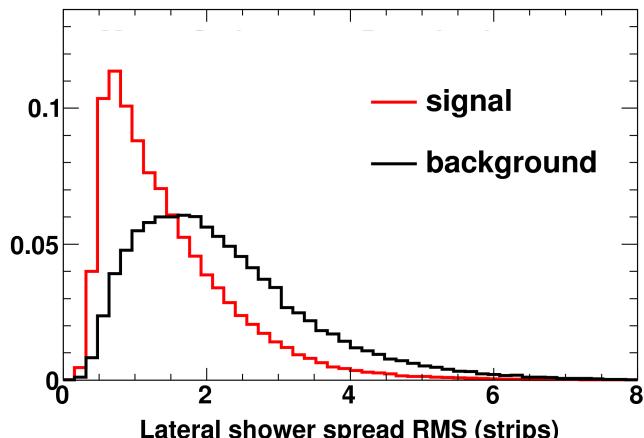
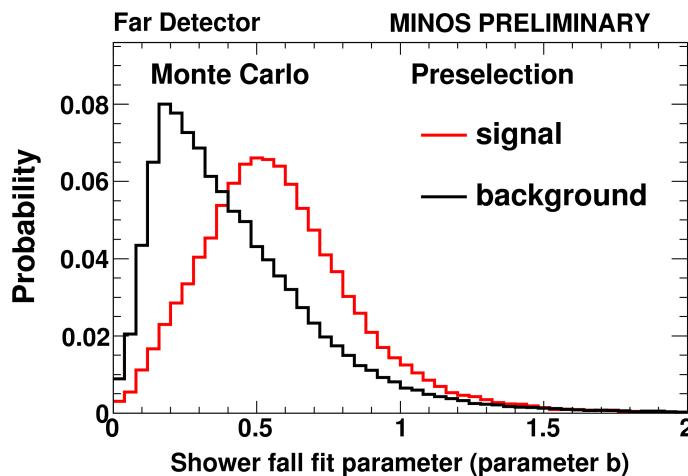
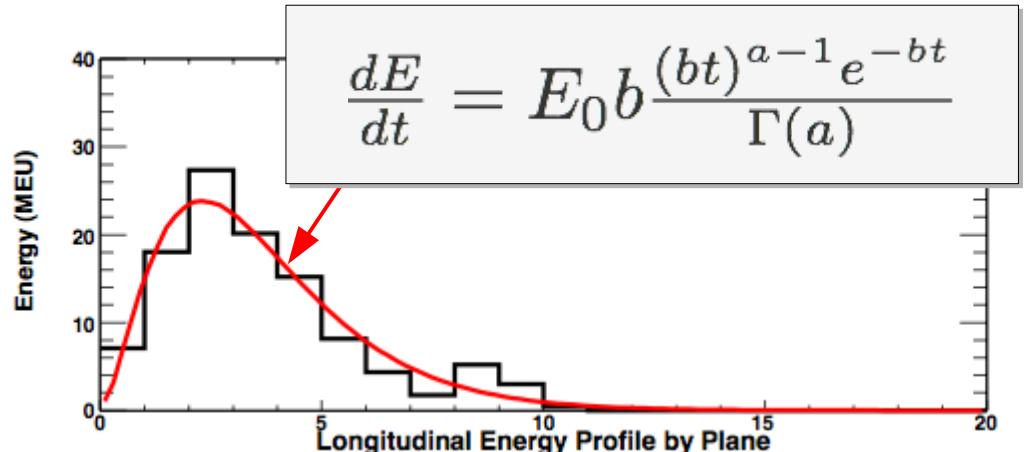


Shower identification

- Classify shower topology with an **ANN**
- 11 discriminant variables derived from longitudinal and transverse **profiles of energy deposition**
- *Examples:* shower fall off (parameter b), shower width metrics



Example EM shower profile



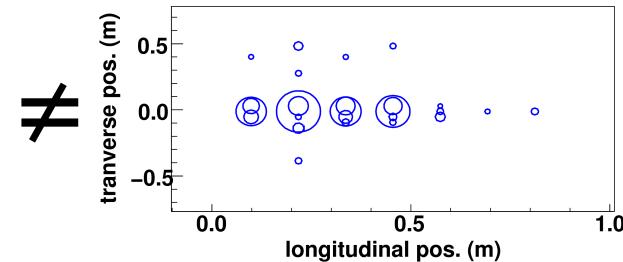
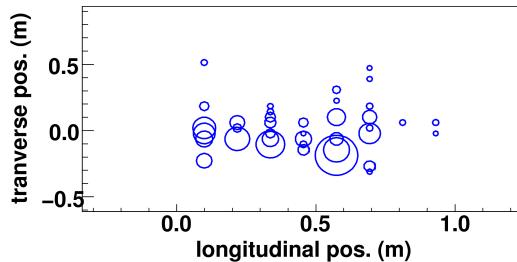
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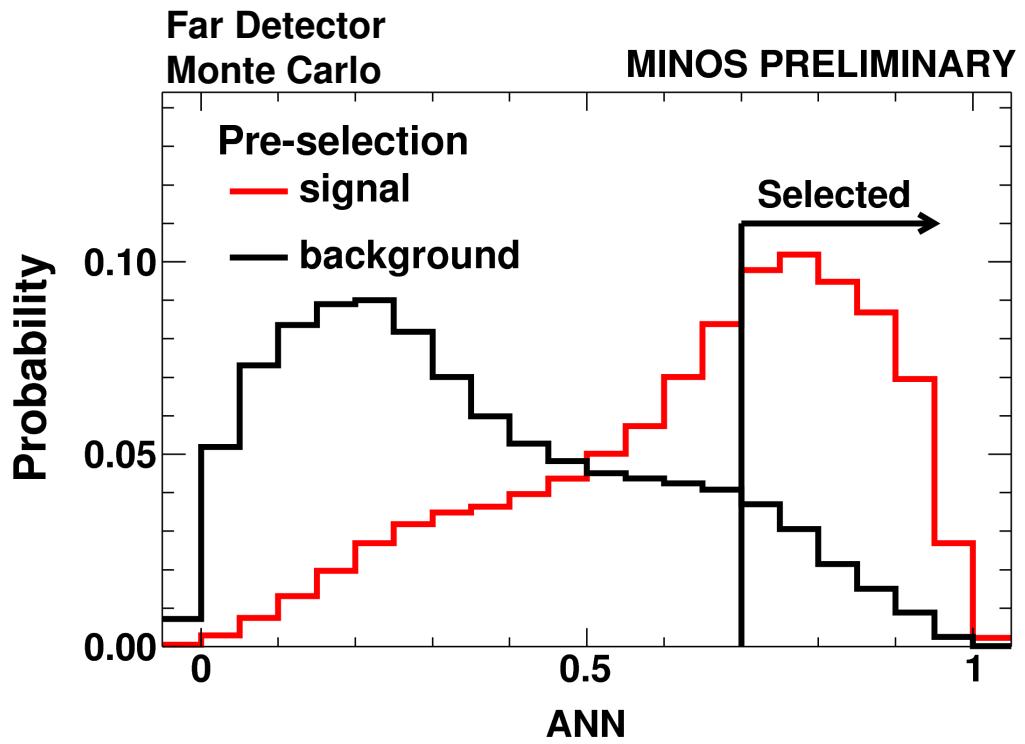
Selection efficiencies (presel.+ANN):

signal 42%

NC 5.4%
CC 0.4% } (i.e., mis-ID rates)



Resulting ANN discriminant:



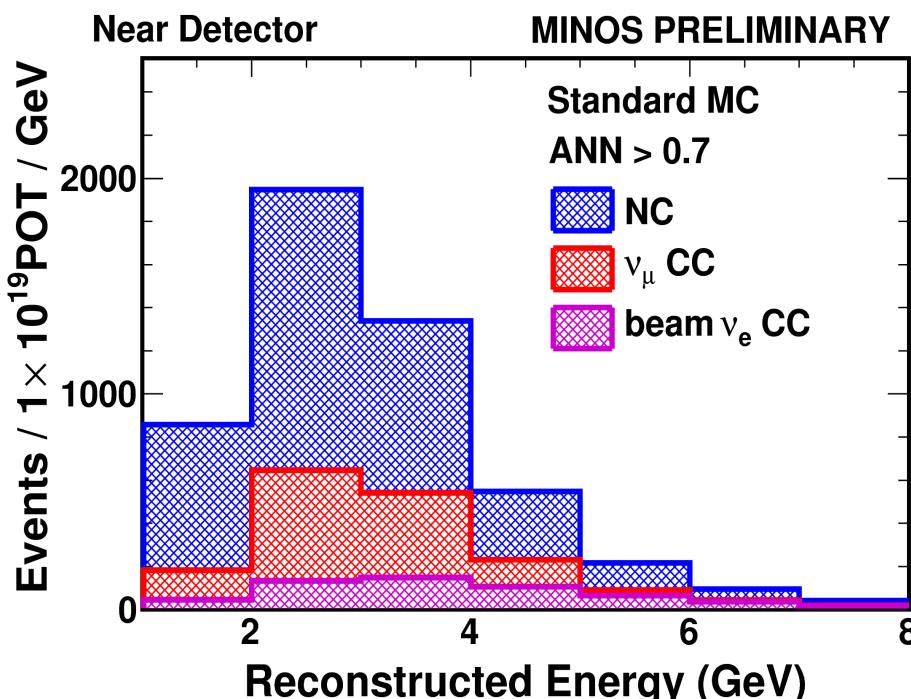
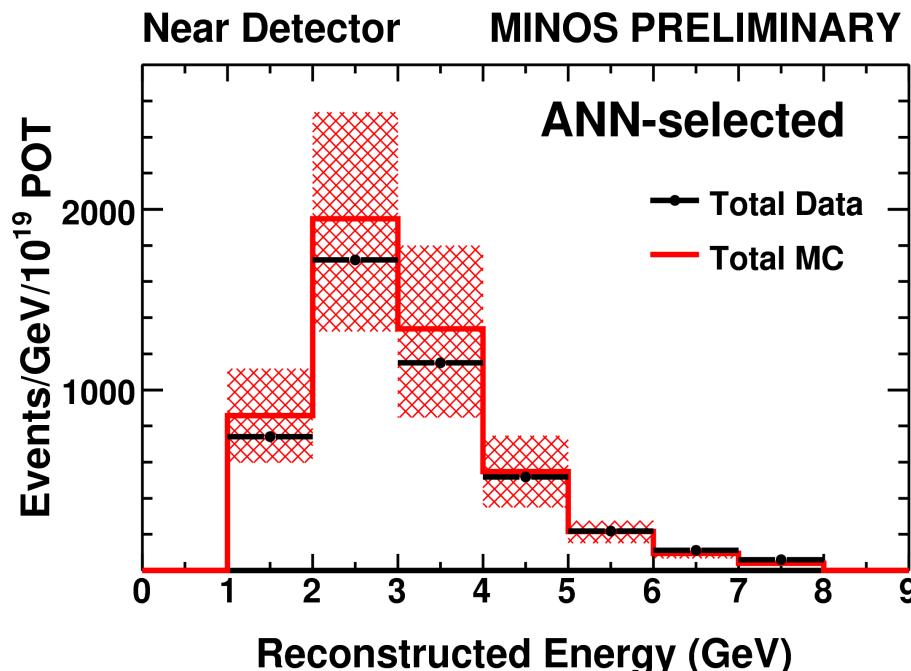
Near detector events

ν_e candidates in the near detector (ANN selection)

- Data and MC differ by up to 15%, but...
- ... are consistent given the large hadronic model uncertainties (*red error band*)

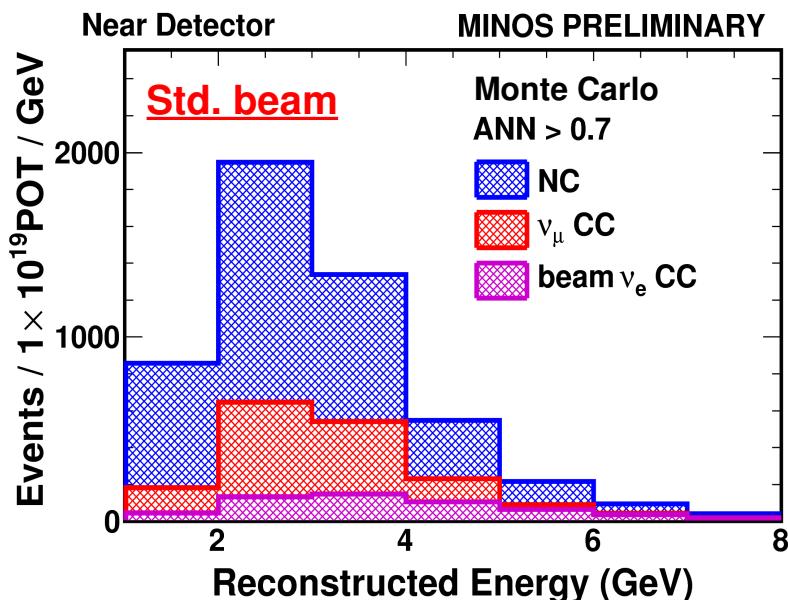
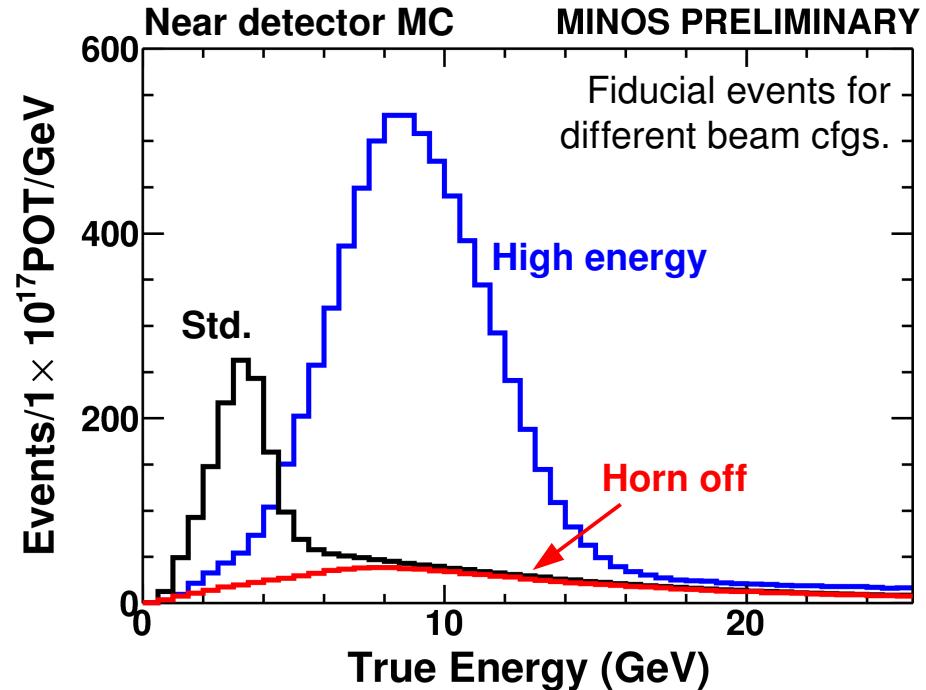
We need not rely on the simulation of the primary backgrounds (neither raw production nor selection efficiencies)

- Observed **near** detector rate is converted to a **far** detector prediction

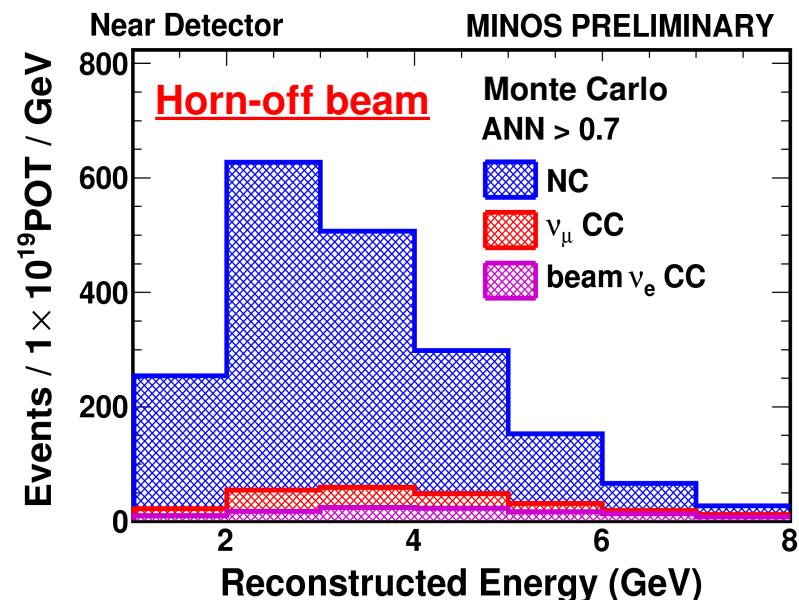


Background decomposition

- Transport of **CC components** to far det. requires application of $P_{\text{osc}}(\nu_\mu \rightarrow \nu_x)$
- Could use MC to estimate fraction of background that is ν_μ and ν_e CC
- *Better: measure NC, CC components* by adjusting horn focusing, modifying NC / CC fraction



Turn off
focusing horn

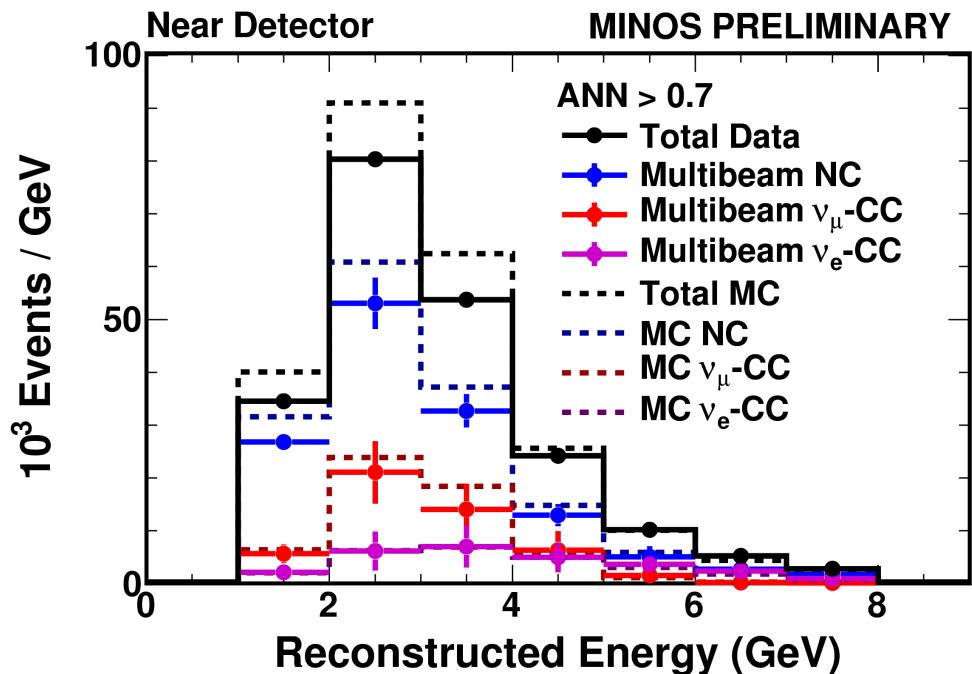


- Approximately...

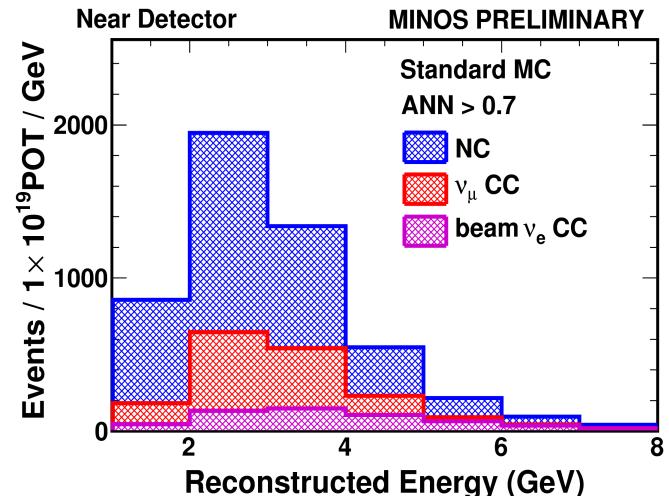
- horn-off and high- E \Rightarrow NC
- tuned MC \Rightarrow ν_e CC
- unitarity \Rightarrow ν_μ CC

- Actually...

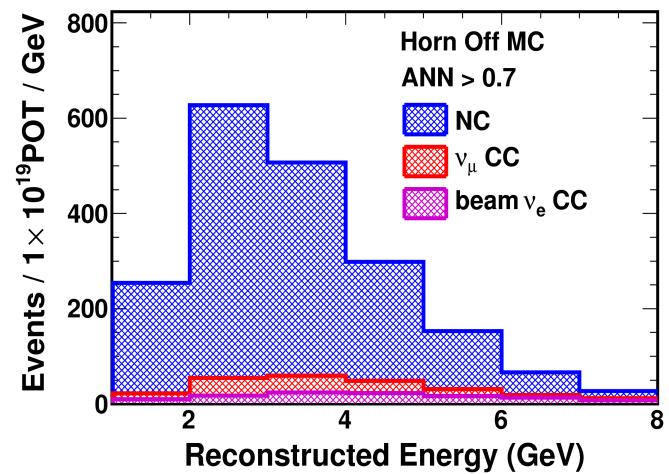
- Solve the full **linear system**
(3 beam configurations and
3 background components)
- ...with injected MC ν_e constraint



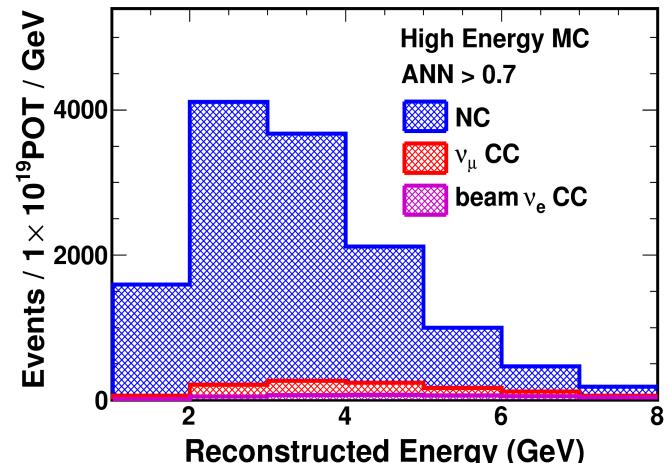
Standard



Horn off



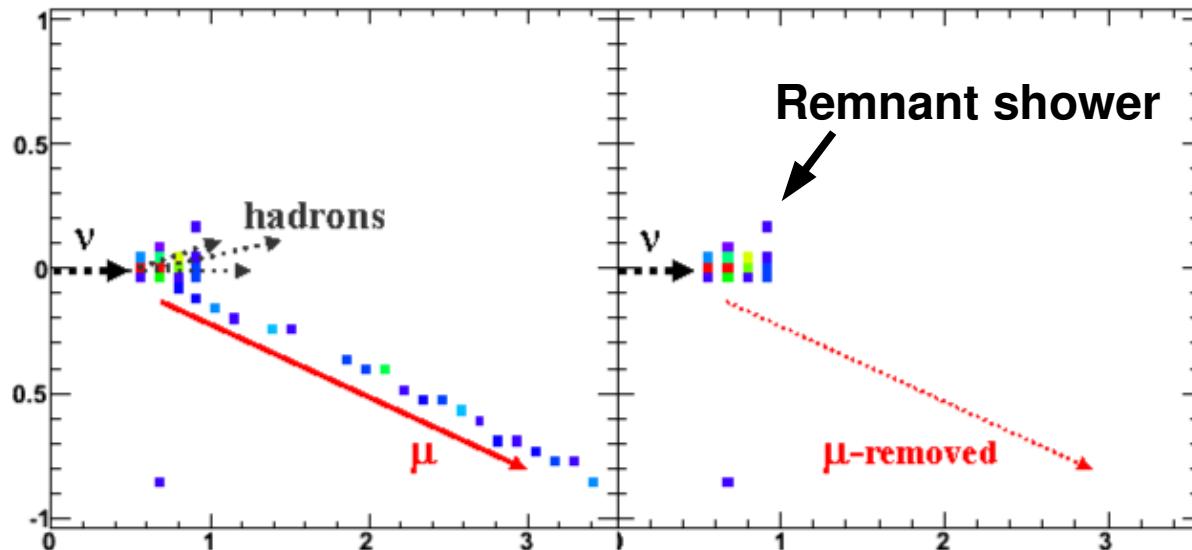
High energy



A cross check...

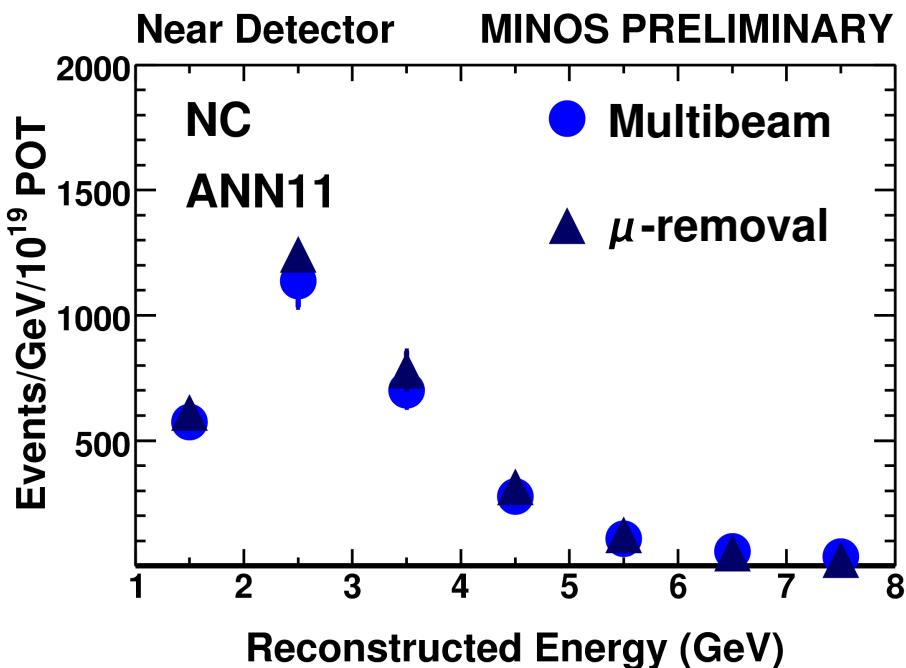
Alternatively

Use **mock NC events** to measure **NC acceptance**



Remove the hits from μ tracks in ν_μ CC events

Apply ν_e selection to the remnants (*mock NC events*)



Data/MC ratio with mock NC events
⇒ correction for real NC events

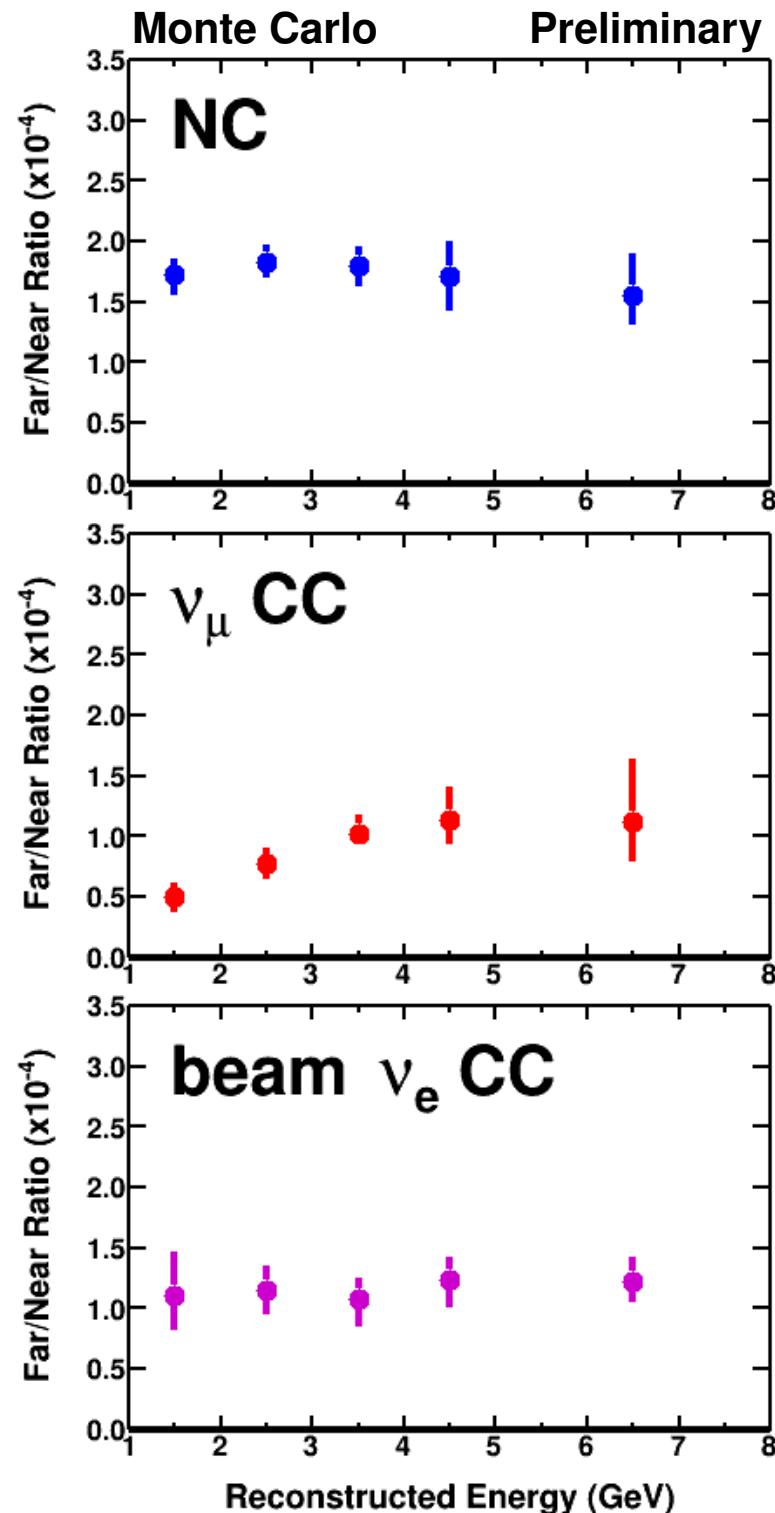
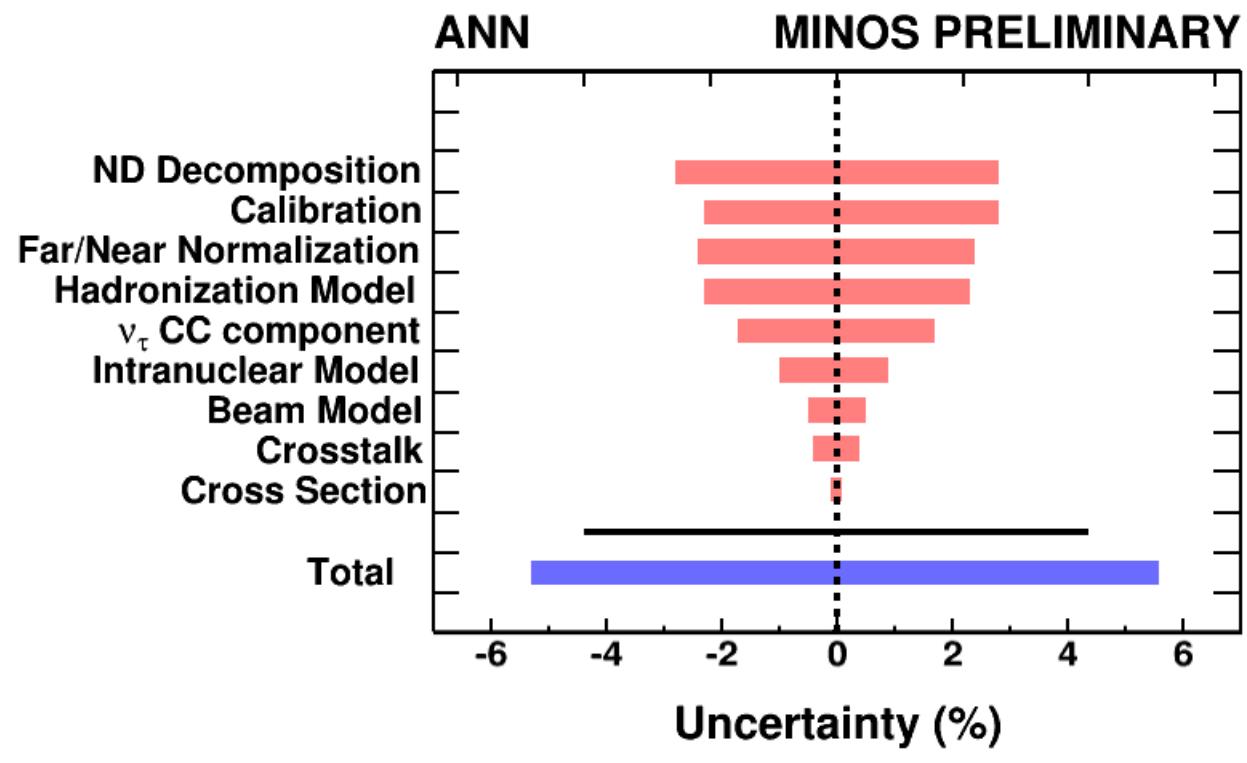
Resulting “measured” NC component agrees between the two methods

From near data to far prediction

Far/near ratios in bins of energy →

Detector differences lead to systematic errors in far/near ratio

*attenuation, readout (single vs. double),
PMT design, crosstalk, ...*

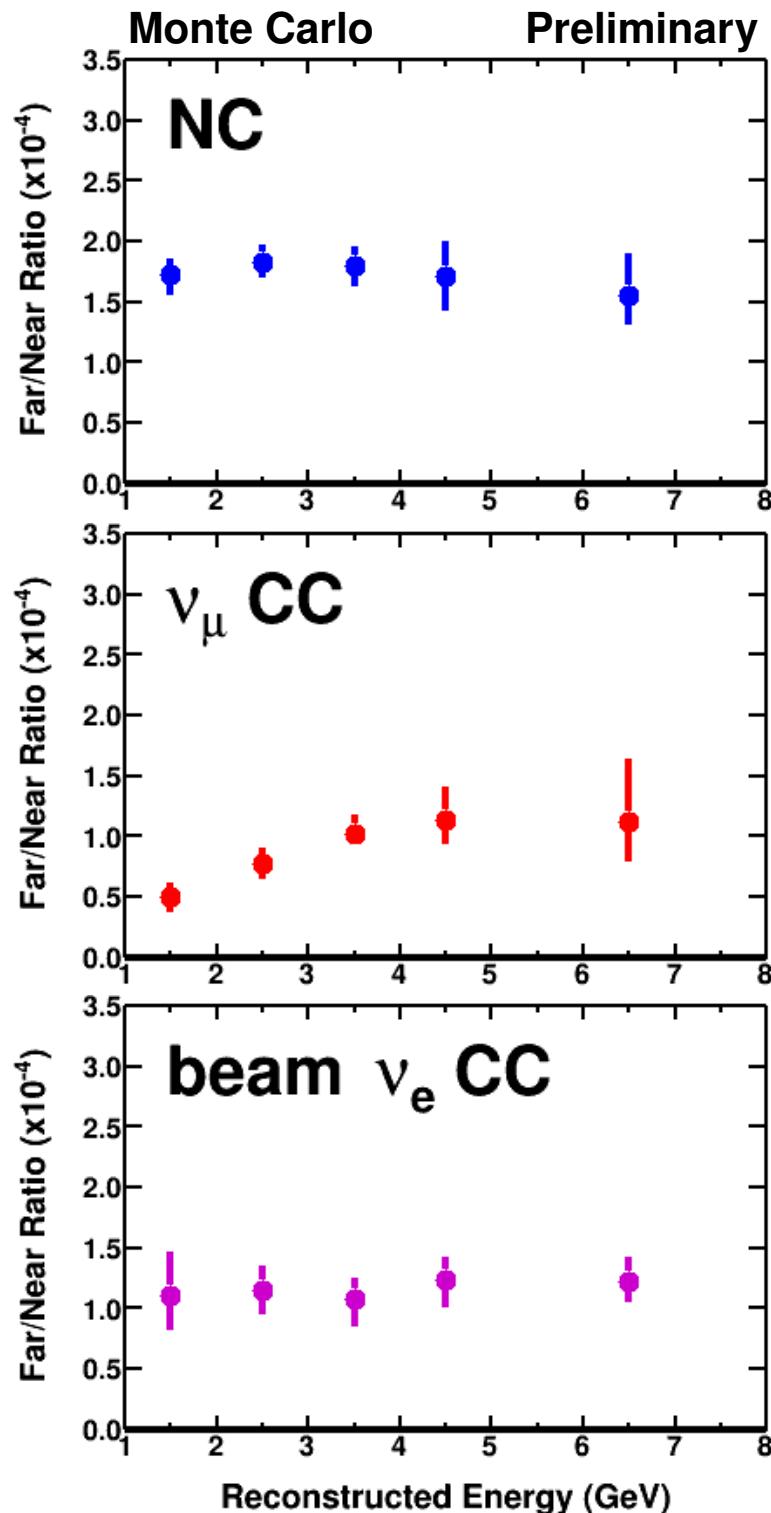
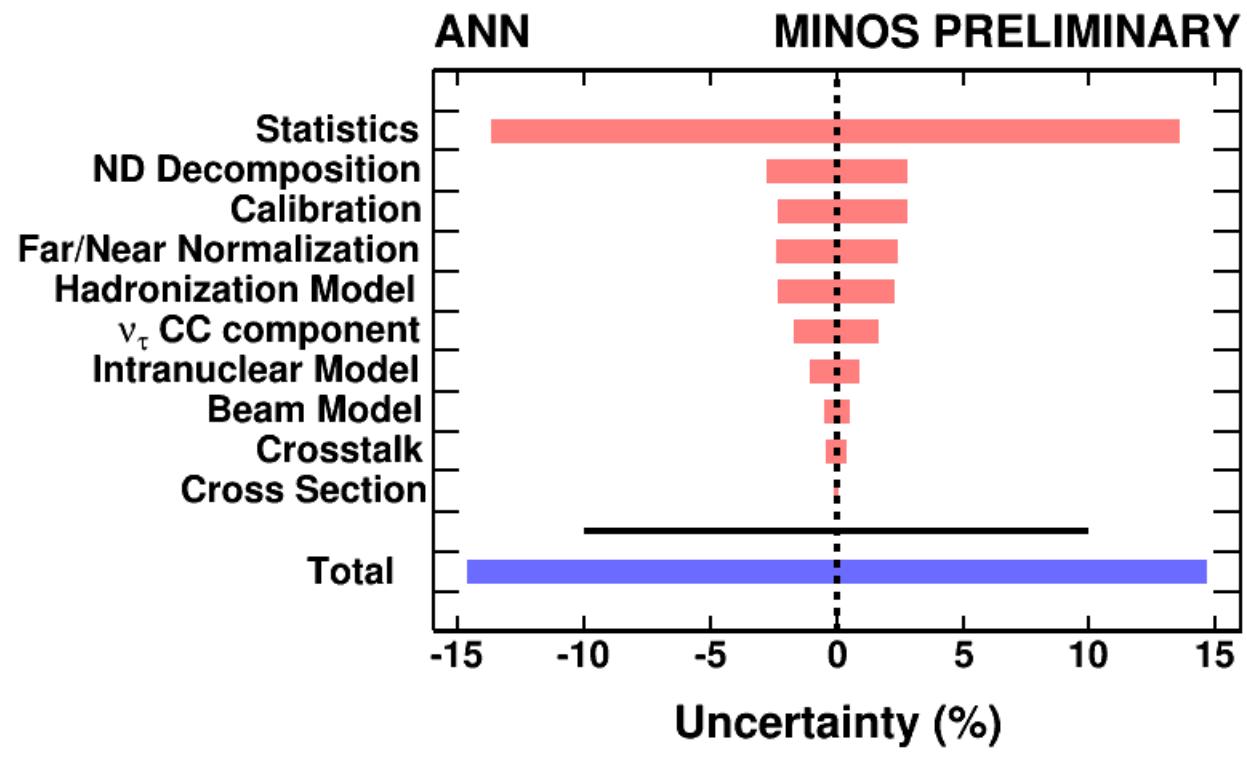


From near data to far prediction

Far/near ratios in bins of energy →

Detector differences lead to systematic errors in far/near ratio

*attenuation, readout (single vs. double),
PMT design, crosstalk, ...*



Far detector prediction

Small ν_τ CC background (*from $\nu_\mu \rightarrow \nu_\tau$ oscillations*) estimated from the Monte Carlo (with knowledge of ν_μ disappearance)

7×10^{20} p.o.t. predicted event counts:

Backgrounds:

Total 49.1

NC 35.8

ν_μ CC 6.3

ν_e CC 5.0

ν_τ CC 2.0

Signal (CHOOZ limit):

ν_e CC 24

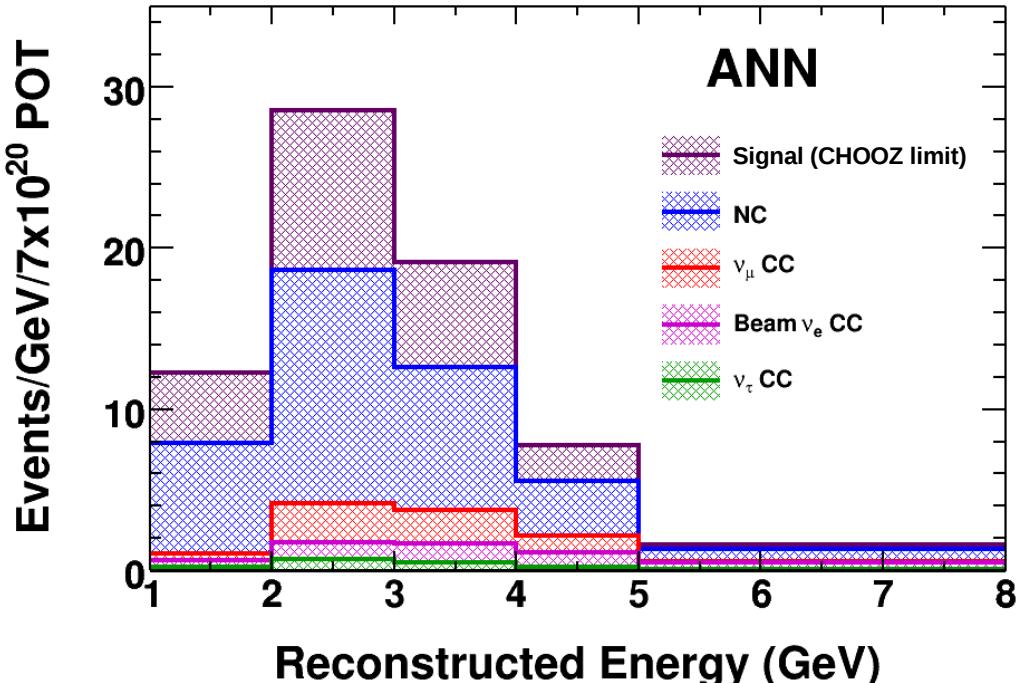
(stat. + syst. error on total BG is 7.4 events)

Signal: $\Delta m_{32}^2 = 2.43 \times 10^{-3} \text{ eV}^2$
 $\sin^2(2\theta_{13}) = 0.15$

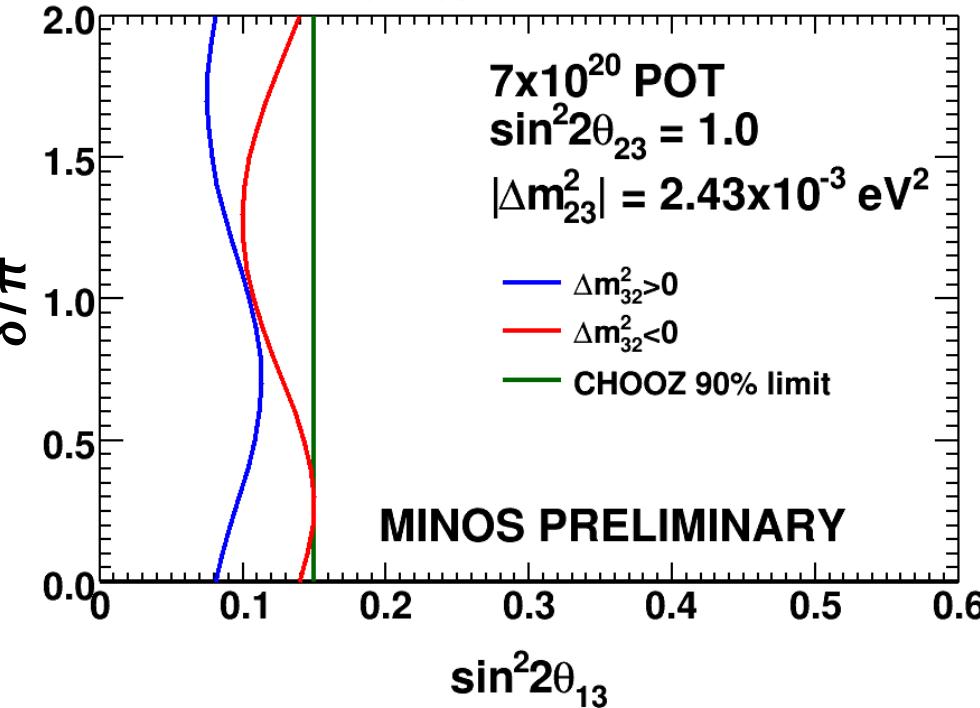
$\sin^2(2\theta_{23}) = 1.0$
 $\delta_{CP} = 0$

Far Detector

MINOS PRELIMINARY



ANN Sensitivity (Hypothetical 90% C.L. Limit)



Blind analysis checks

Check of the analysis – *invert the cut*

Perform ***the full analysis*** with one modification:

→ **invert the cut** on the ν_e discriminant: **ANN<0.5** rather than **ANN>0.7**

Expected events for ANN<0.5

For $\sin^2(2\theta_{13}) = 0$

Total **314**

NC 186

ν_μ CC 117

ν_e CC 5

ν_τ CC 6

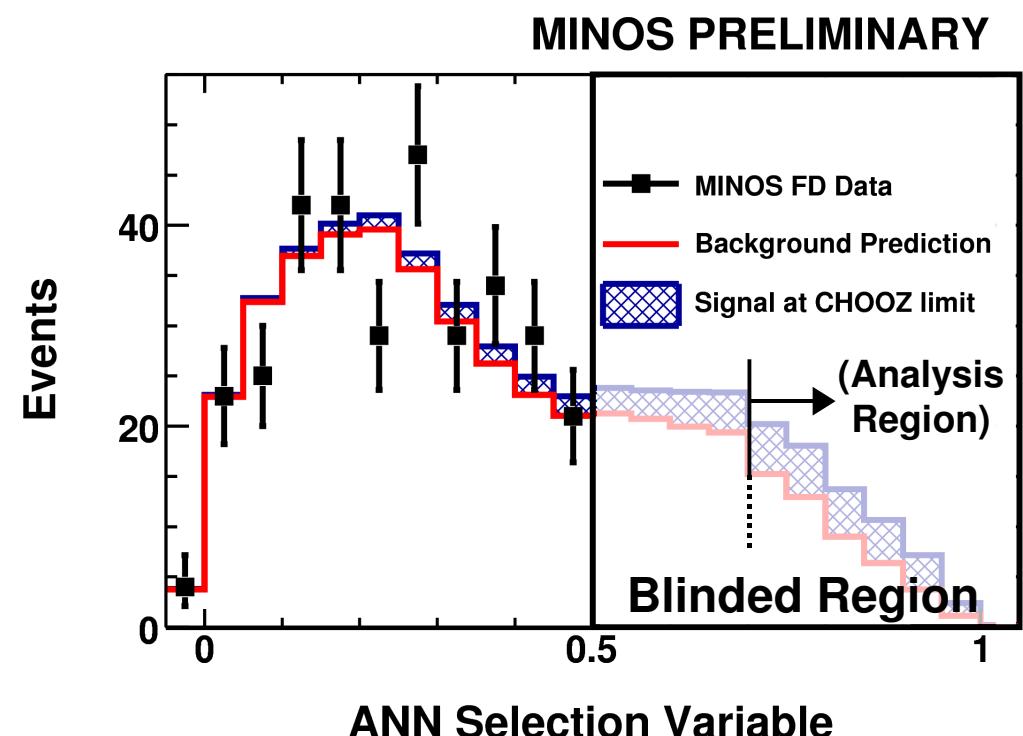
For $\sin^2(2\theta_{13}) = 0.15$

Total **314 + 13**

Observed for ANN<0.5:

327 events

(0.75σ excess if $\sin^2(2\theta_{13})=0$)



(uncertainty in this test) $\approx 0.4 \times$ (uncertainty in main analysis)

Check of the analysis – verify NC shower mis-ID

Revisit the **muon-removed events** (*mock NC events*)

- Measure **selection efficiency** for these showers in **Far Detector data**
- Compare with **Near-Detector-driven prediction**

Predicted and measured efficiencies (vs. energy) agree

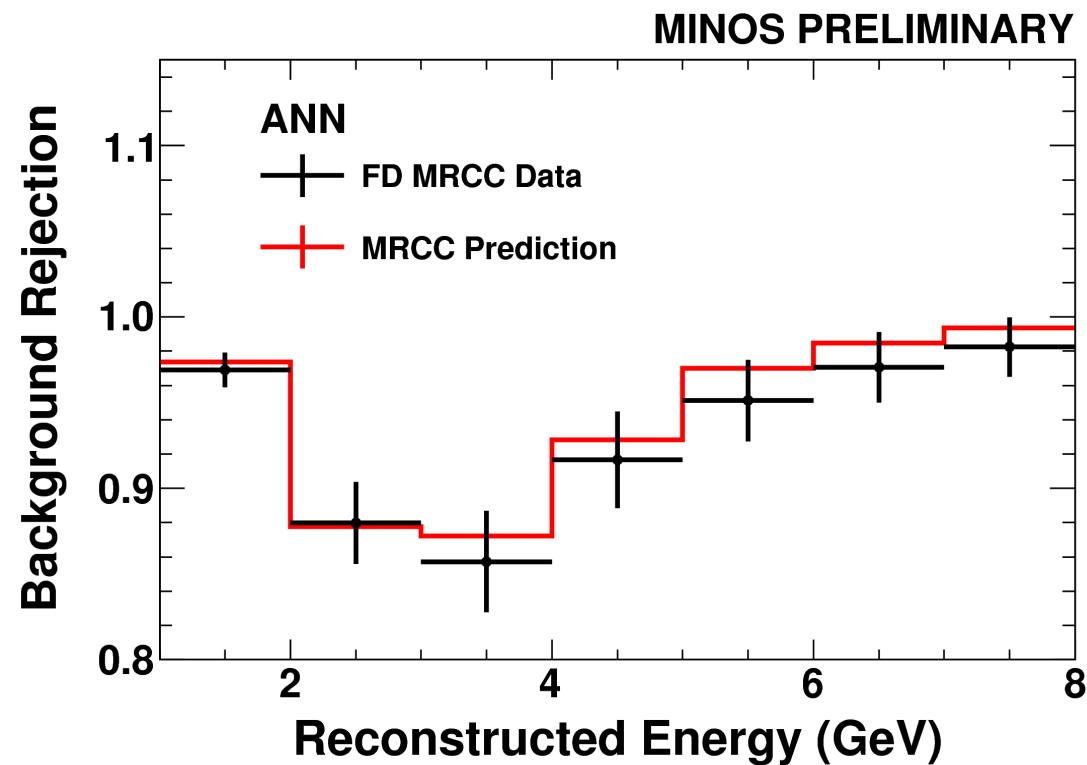
Mis-ID rate for μ -removed events:

data: $(7.2 \pm 0.9)\%$

pred: $(6.42 \pm 0.05)\%$

[stat. errors only]

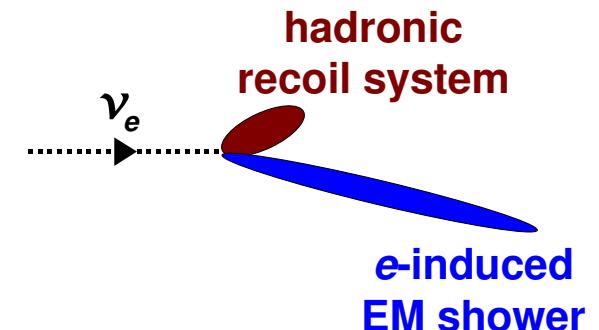
⇒ compatible at 0.86σ



(uncertainty in this test) \approx (uncertainty in main analysis)

Check of the analysis – verify signal efficiency

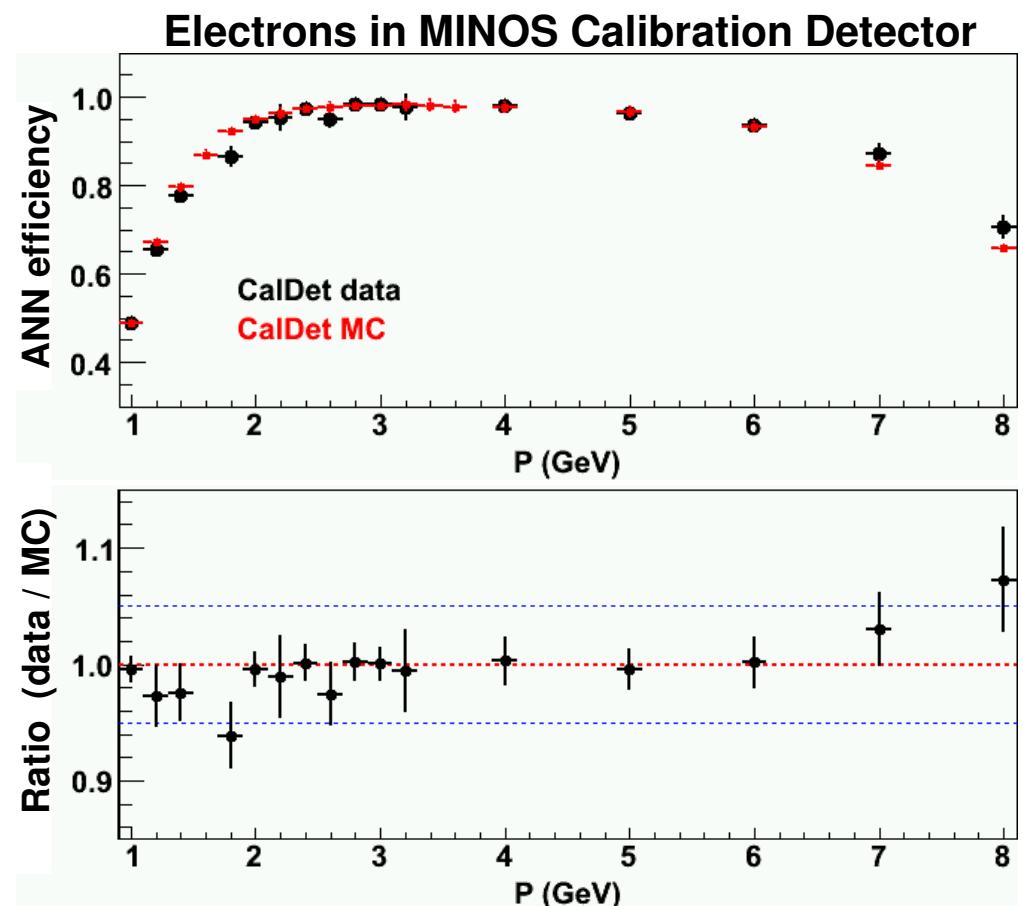
ν_e CC events: **EM shower + hadronic shower**



1st: Test selection on **pure EM showers** from CalDet

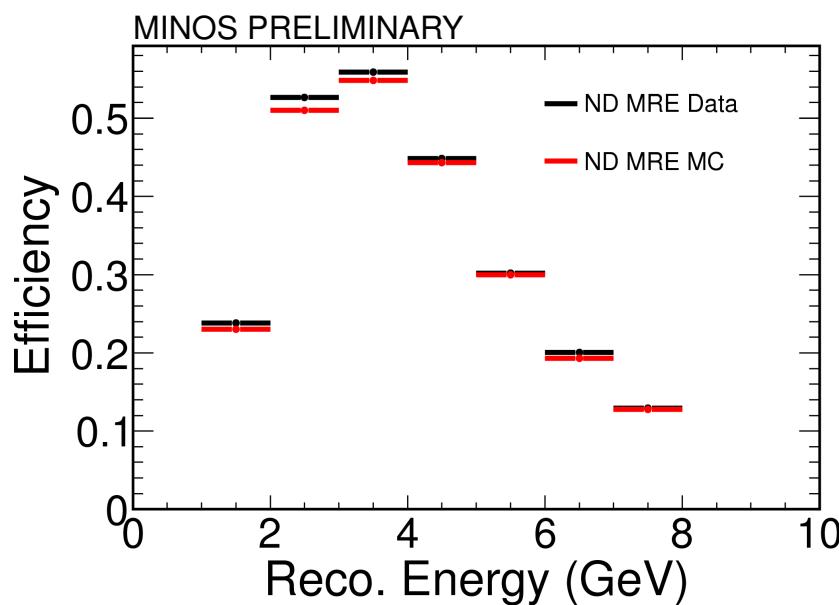
Efficiency understood to better than 3%

Apply a systematic error based on the limits of this comparison

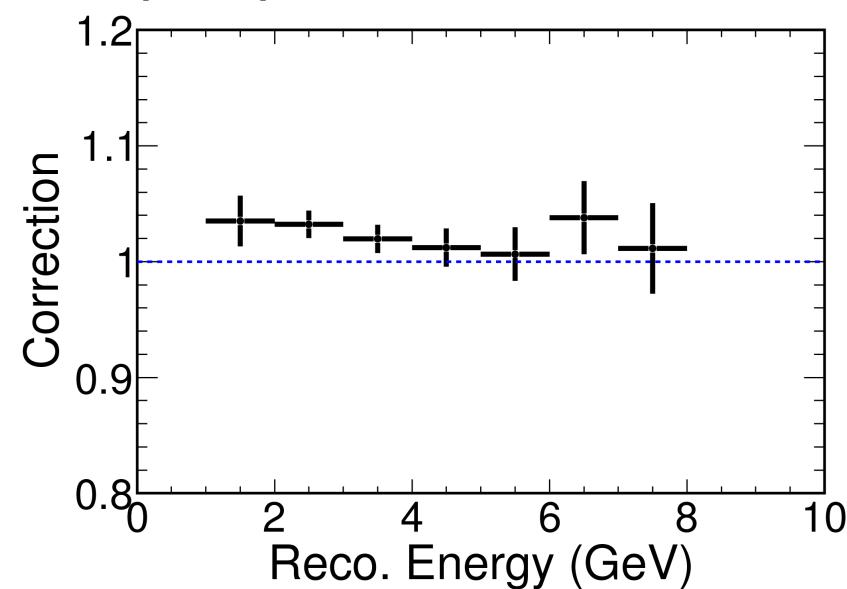


2nd: Test importance of **hadronic shower** modeling:

→ use μ -removed sample to get **hadronic showers** from data or MC



Small efficiency difference (~3%) taken as a correction



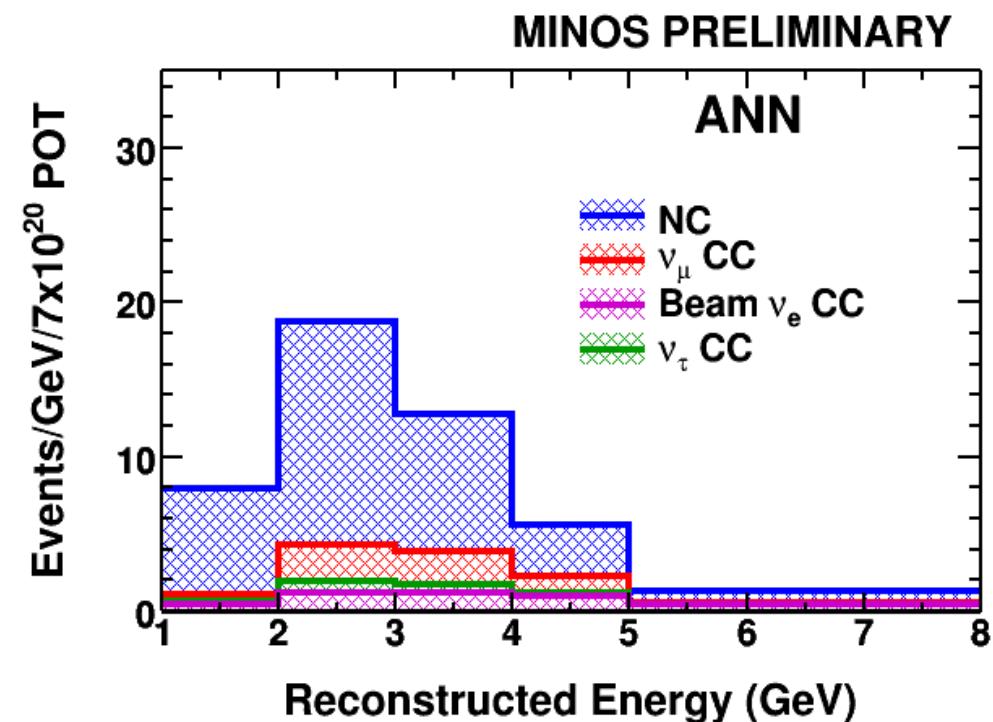
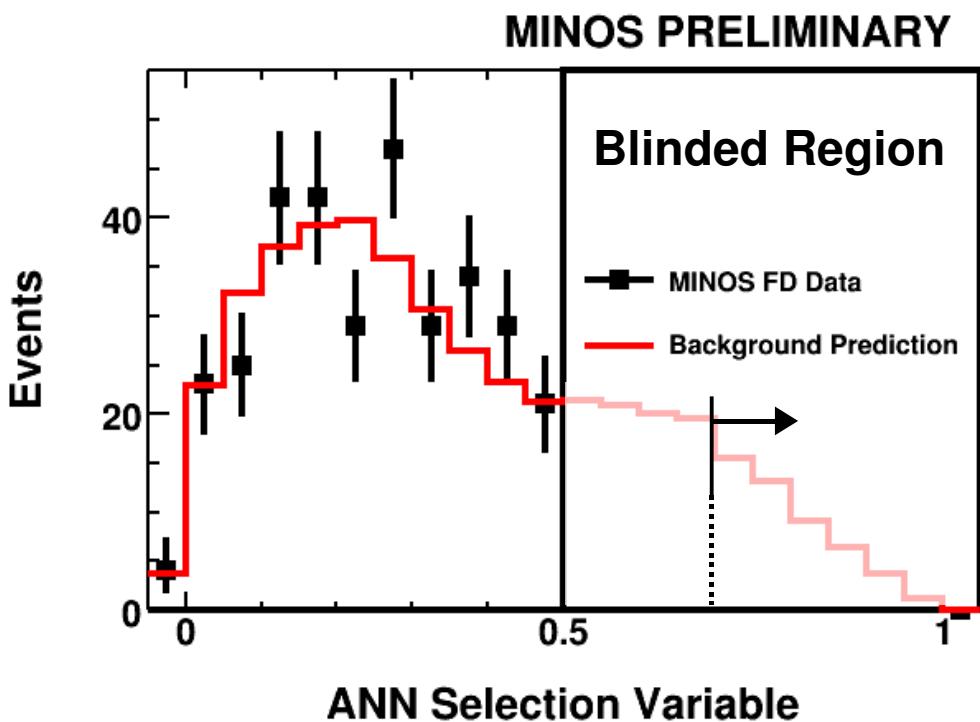
Results

ν_e appearance result

With 7×10^{20} protons on the NuMI target

ν_e charged current candidate events:

background expectation: $49.1 \pm 7.0(\text{stat.}) \pm 2.7(\text{syst.})$

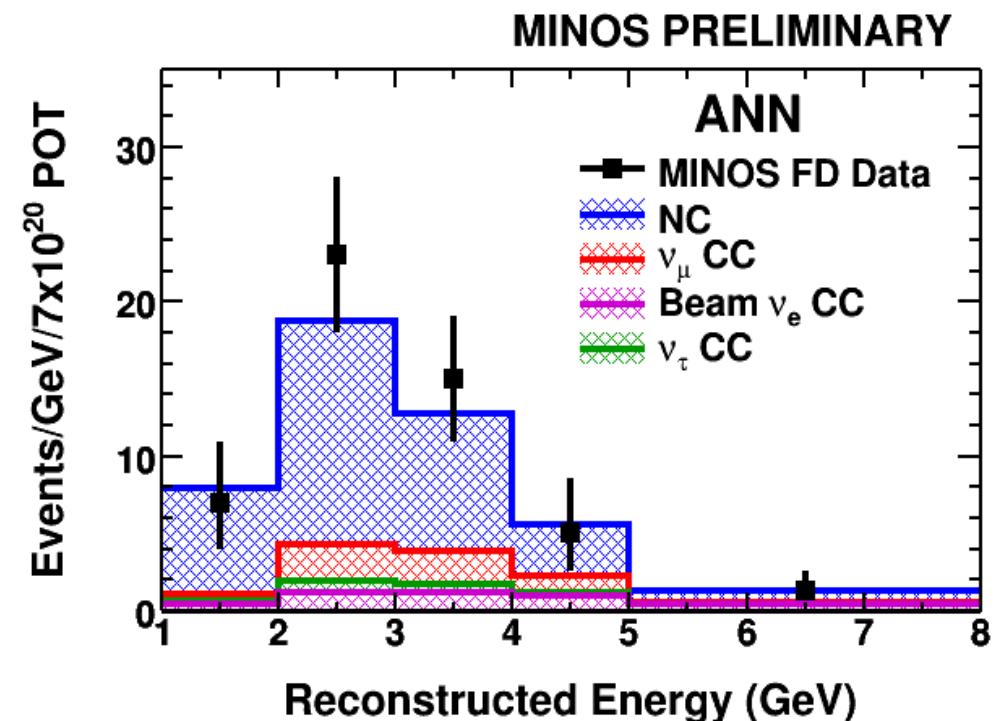
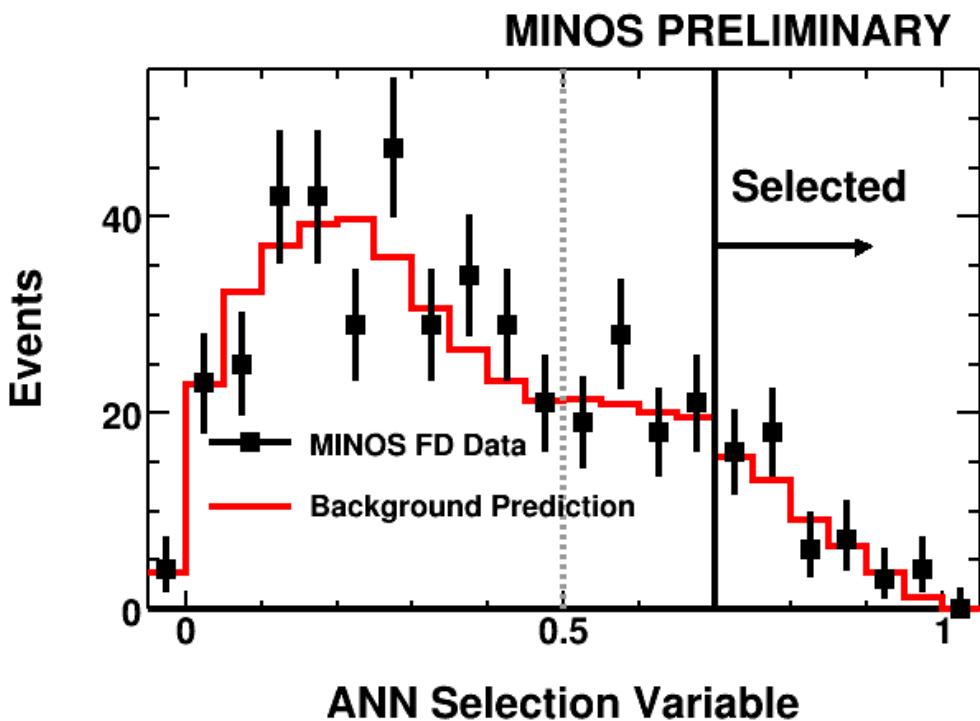


ν_e appearance result

With 7×10^{20} protons on the NuMI target

ν_e charged current candidate events:

background expectation: $49.1 \pm 7.0(\text{stat.}) \pm 2.7(\text{syst.})$
observed: 54
(0.7σ excess)

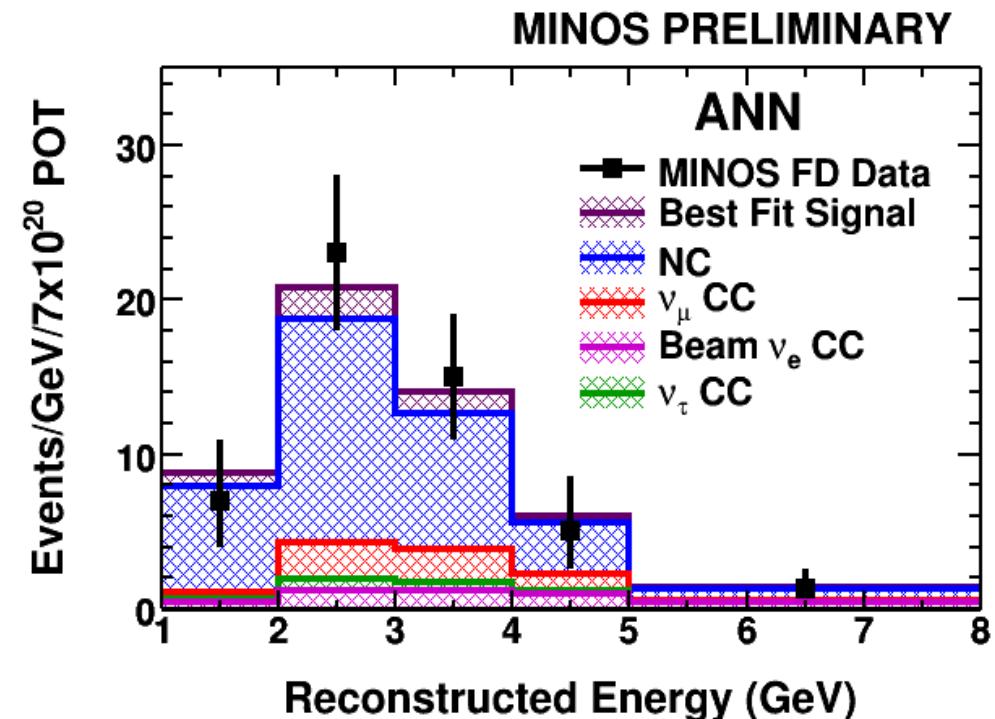
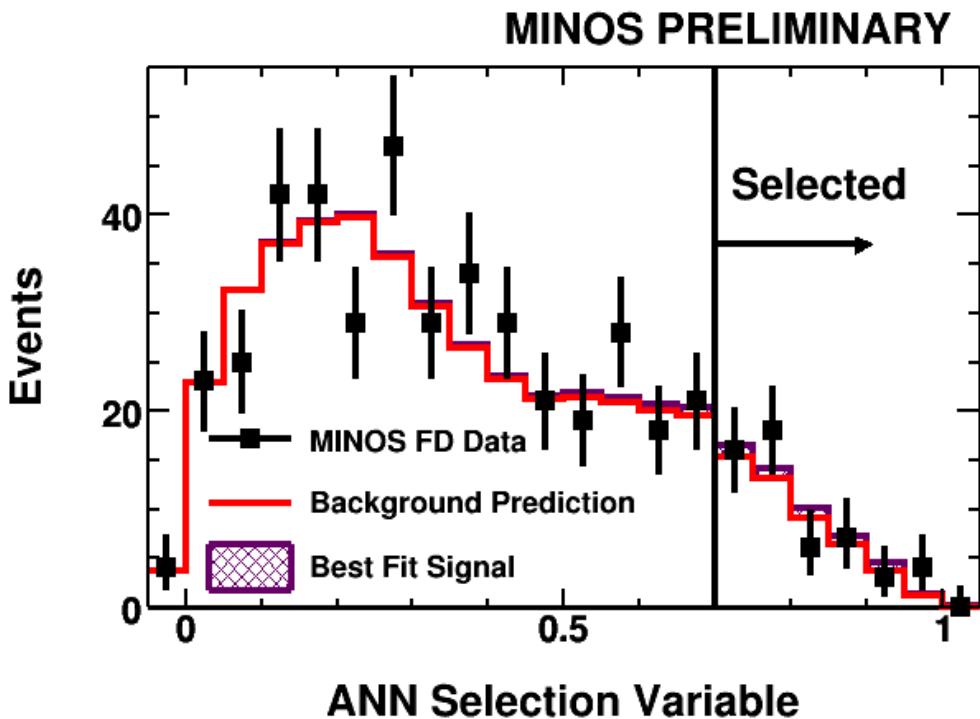


ν_e appearance result

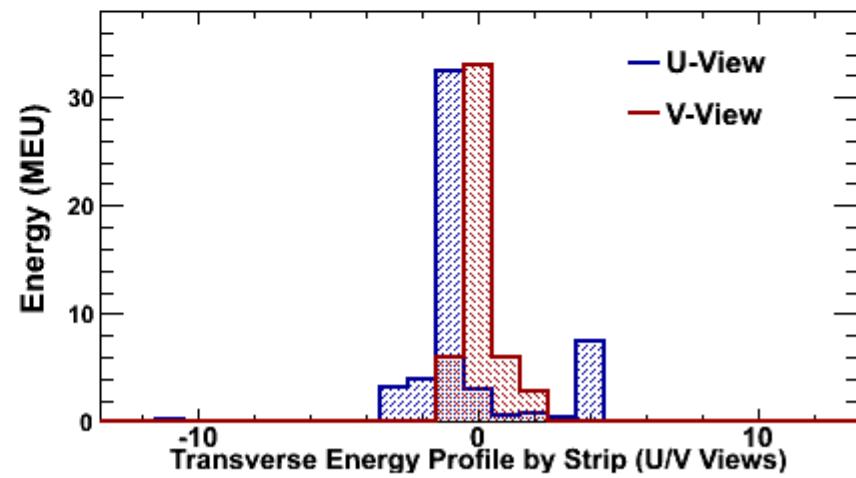
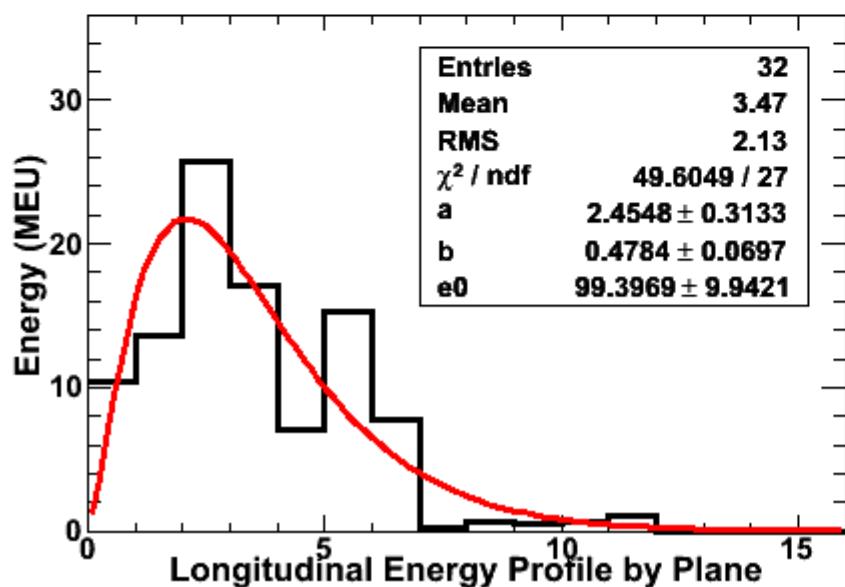
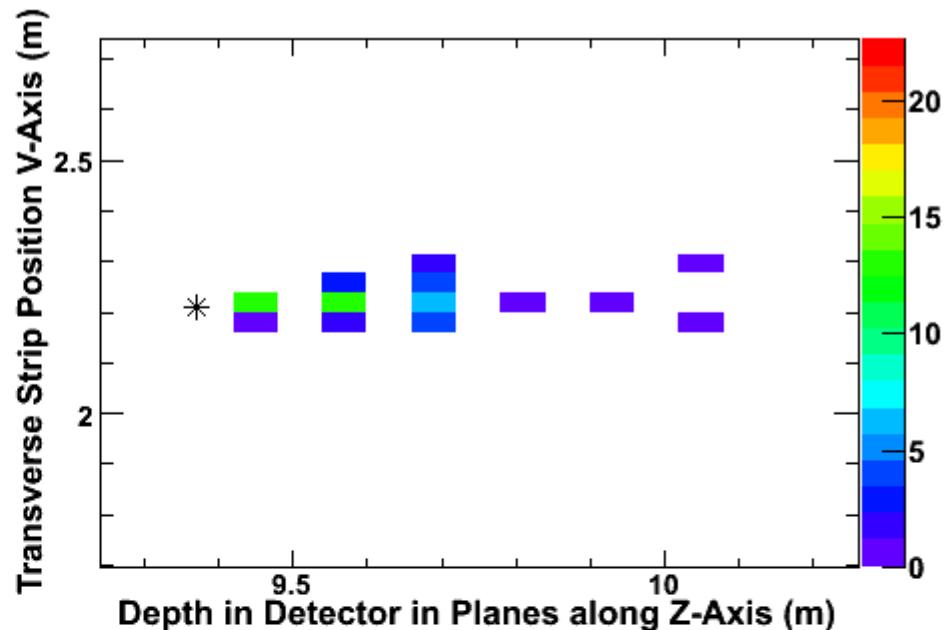
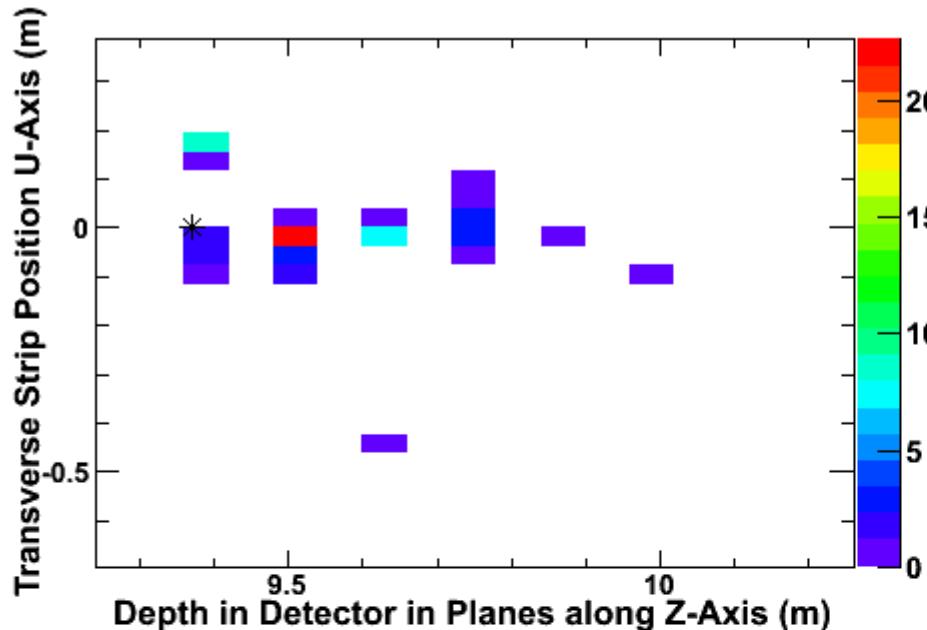
With 7×10^{20} protons on the NuMI target

ν_e charged current candidate events:

background expectation: $49.1 \pm 7.0(\text{stat.}) \pm 2.7(\text{syst.})$
observed: 54
(0.7σ excess)



ν_e CC candidate event in the far detector...



FD RUN = 43563

EVENT ID = 95374

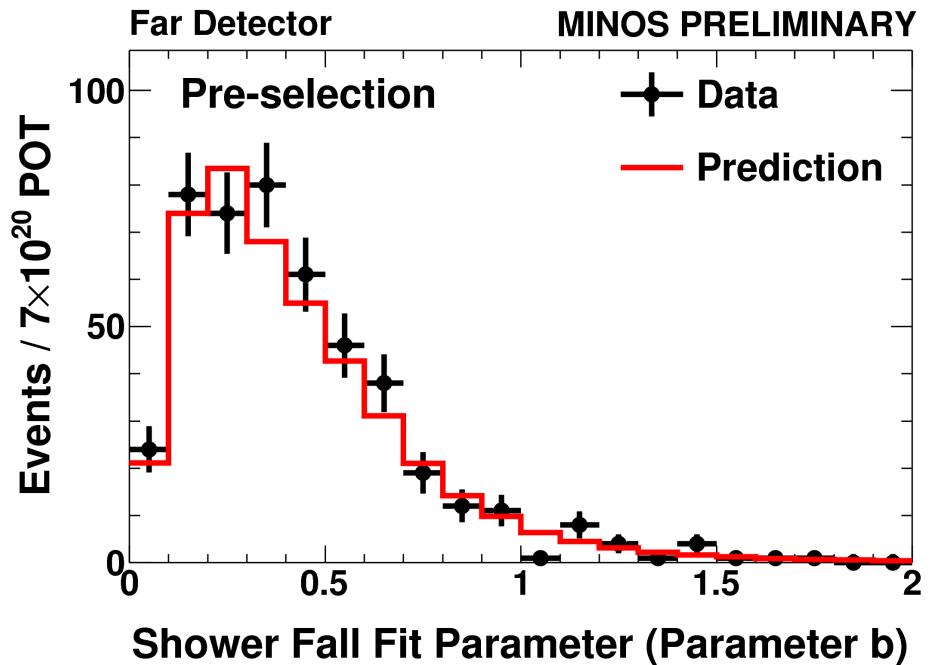
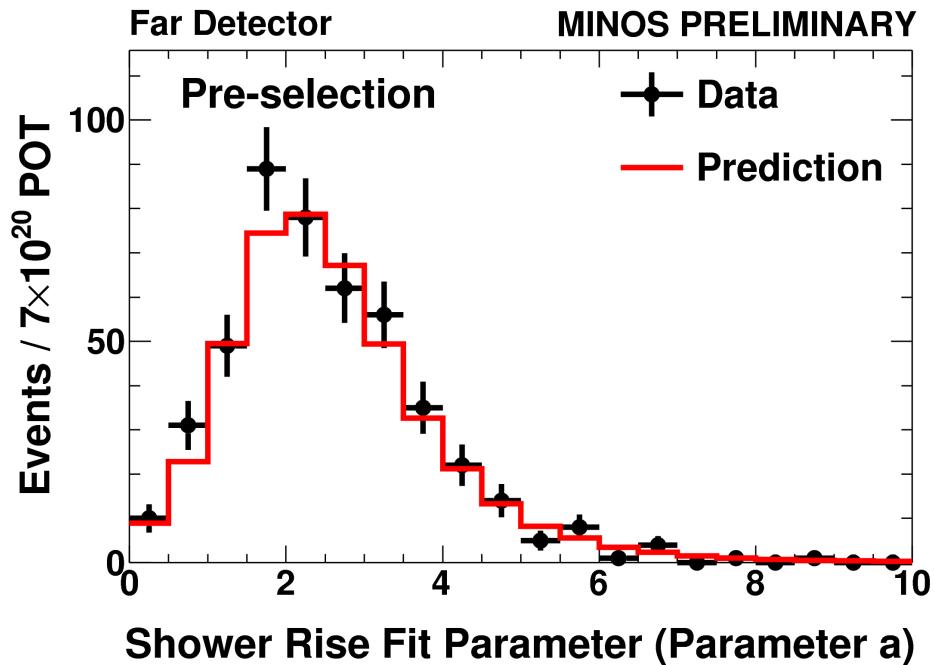
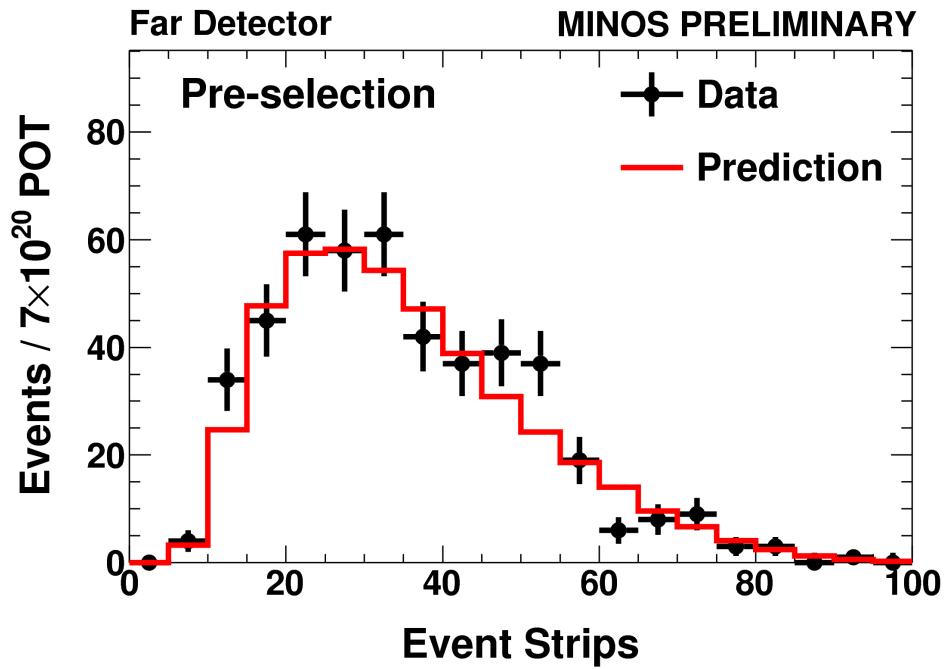
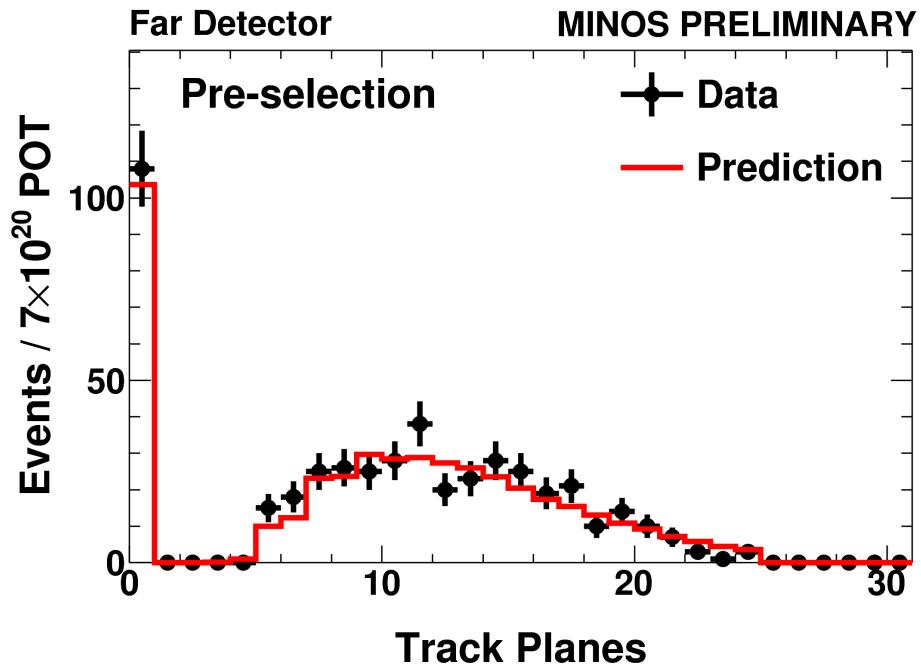
Reco. Energy = 4.35 GeV

ANN11 PID = 0.81

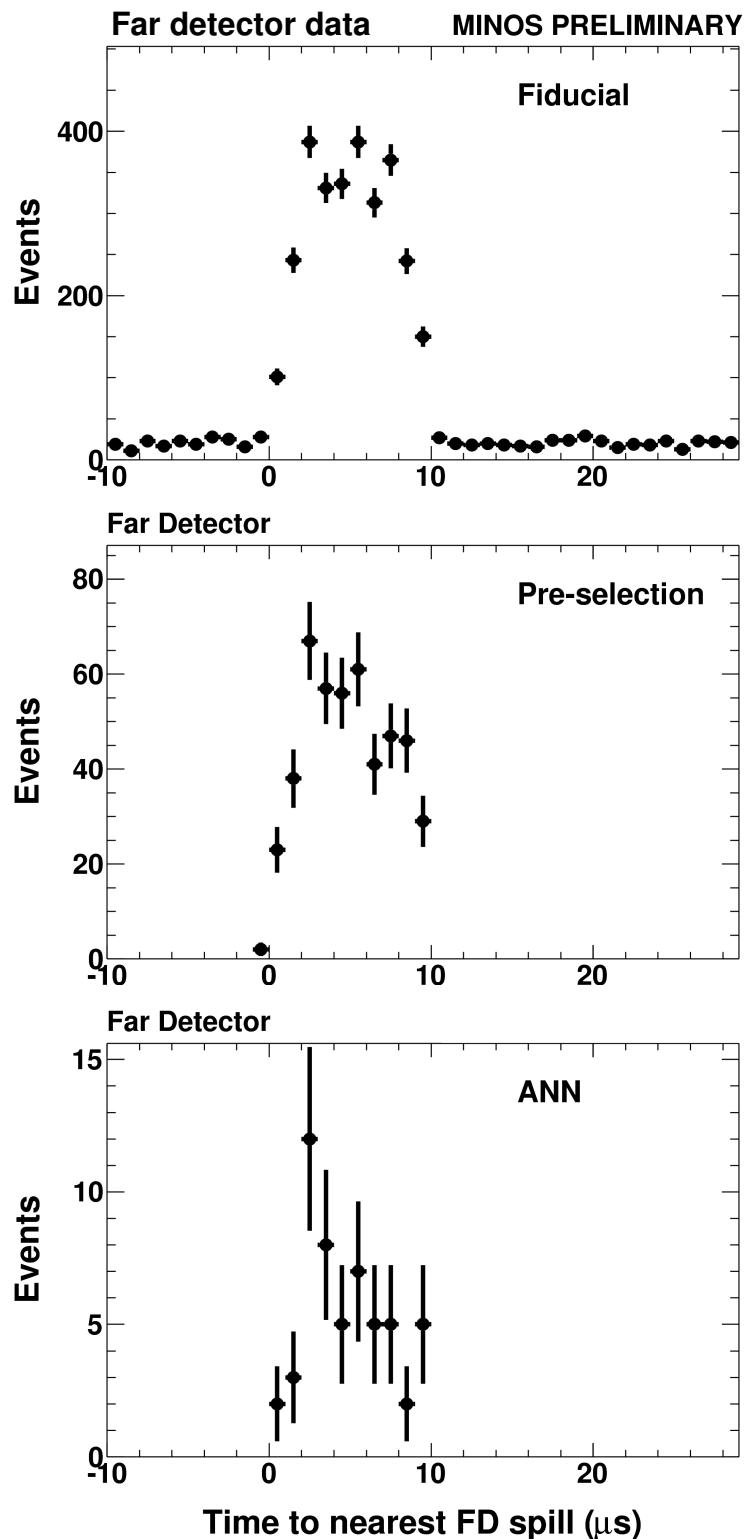
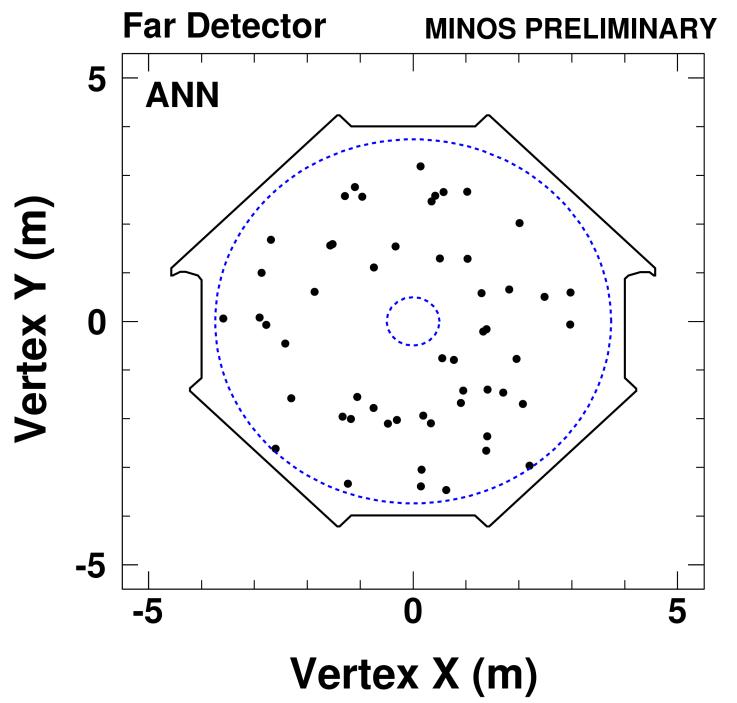
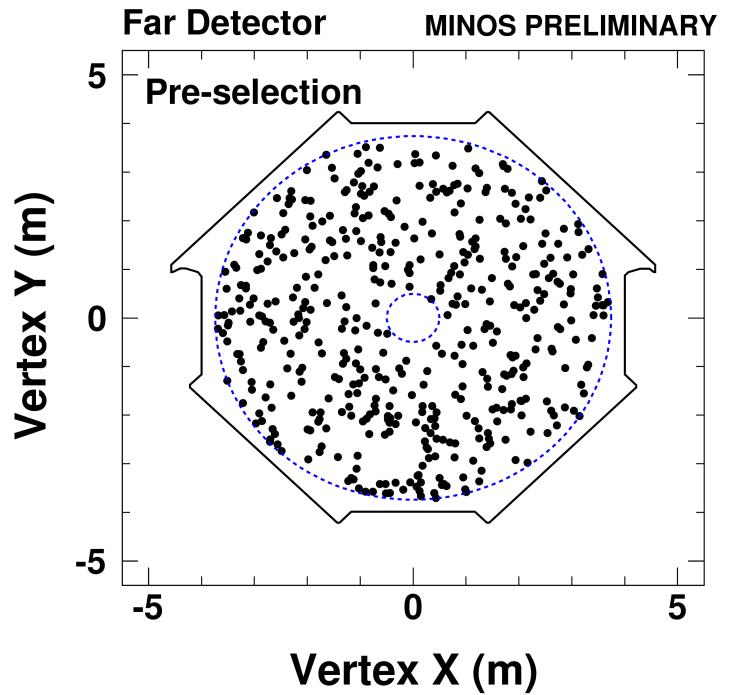
PAR PID = 0.85

ANN14 PID = 0.86

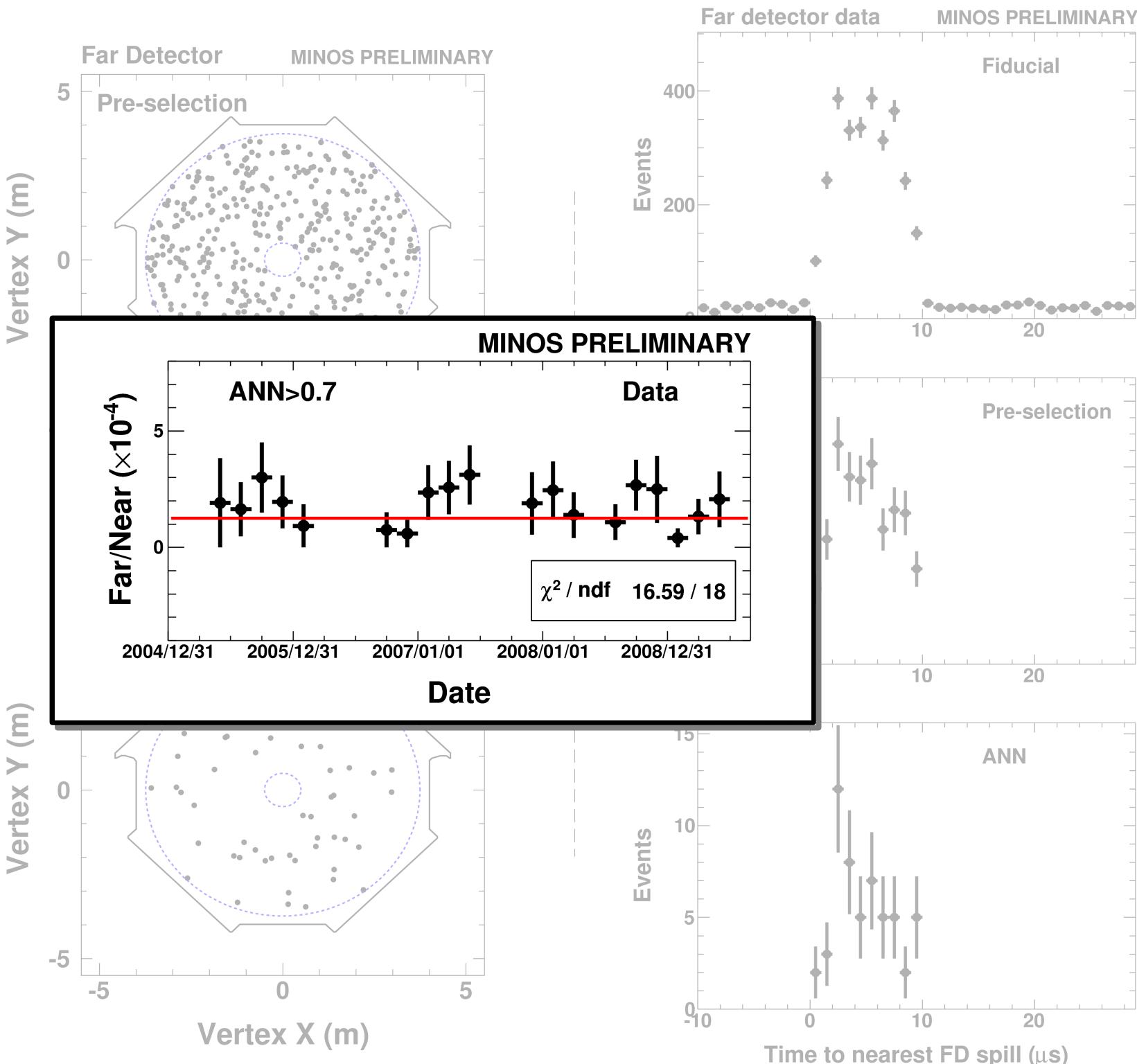
Event shapes in the far detector: *Data and expectations match*



Positions and times of events



Positions and times of events

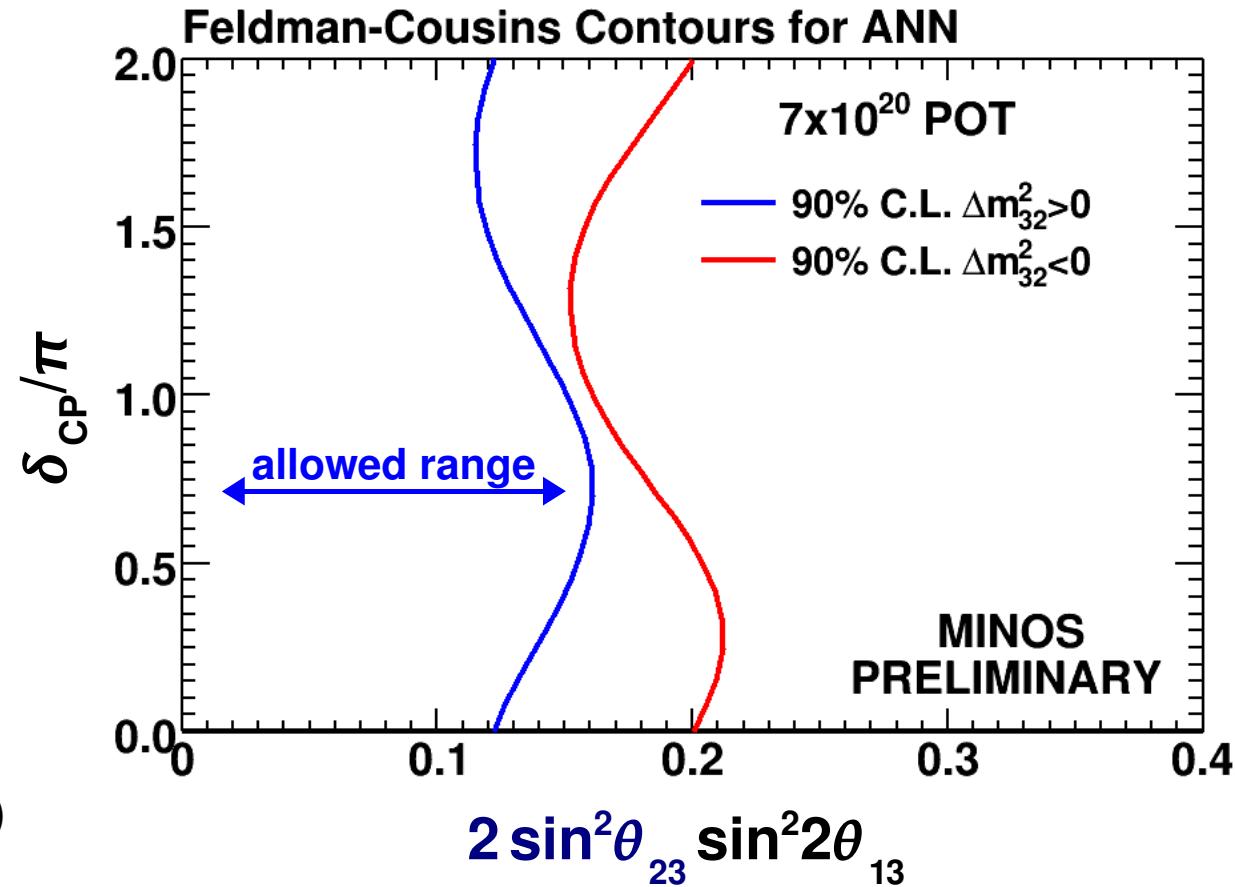


Oscillation interpretation of the data

$\sin^2(2\theta_{13})$ allowed range
depends on CP-phase δ and
mass hierarchy [sign(Δm^2)]

90% C.L. allowed ranges →

(other oscillation parameters taken
with current experimental uncertainties)



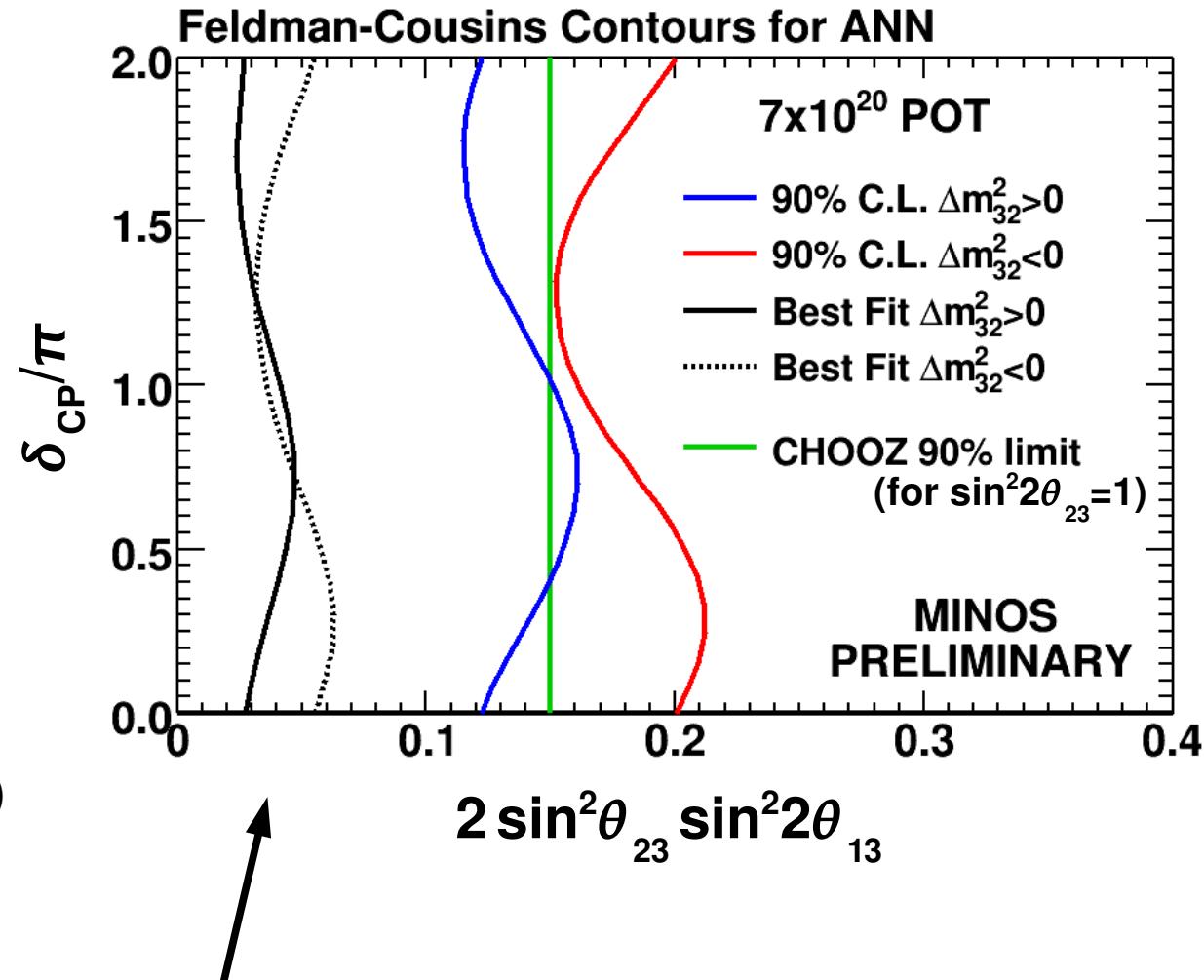
↑
(unity if $\theta_{23} = \pi/4$)

Oscillation interpretation of the data

$\sin^2(2\theta_{13})$ allowed range
depends on CP-phase δ and
mass hierarchy [sign(Δm^2)]

90% C.L. allowed ranges →

(other oscillation parameters taken
with current experimental uncertainties)



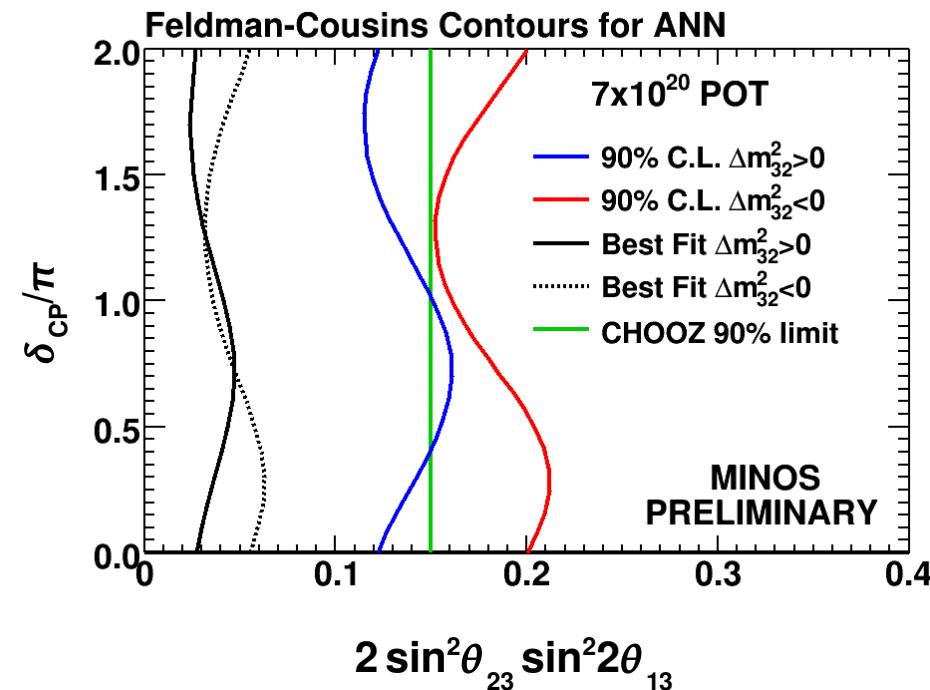
MINOS best-fit [**black curves**] along with
CHOOZ upper limit (for $\theta_{23} = \pi/4$) [**green line**].

Summary

- No significant excess of candidate ν_e CC events:

expected bg: $49.1 \pm 7.0_{\text{stat}} \pm 2.7_{\text{syst}}$
events obs'd: 54

- Consistent results obtained with two additional event selectors (*not shown here*)



Representative 90% C.L. limits: put $\delta_{\text{CP}}=0$, $\theta_{23}=\pi/4$

normal hier. $\Rightarrow \sin^2 2\theta_{13} < 0.12$

inverted hier. $\Rightarrow \sin^2 2\theta_{13} < 0.20$

- **For the near future:**

additional neutrino data and significant new analysis techniques
⇒ potentially large sensitivity gains

Acknowledgments

The MINOS Collaboration would like to thank the **many Fermilab groups** who provided **technical expertise and support** in the design, construction, installation and operation of the MINOS experiment.

Thanks to the **Accelerator Division** for the neutrinos!

We also acknowledge the financial support from DOE; NSF; STFC(UK); the University of Athens, Greece; Brazil's FAPESP, CNPq, and CAPES.

We are grateful to the **University of Minnesota** and the Minnesota **Department of Natural Resources** for hosting us.

